# Classical \& Quantum Chaos In a Non-Linearly Kicked Harmonic Oscillator 

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I hereby certify that this material, which I now submit for assessment on the programme of study leading to the award of $\mathbf{P h D}$ is entirely my own work and has not been taken from the work of others save and to the extent that such work has been cited and acknowledged within the text of my work

## Signed-

I.D. No.:

Candidate

Date.

In Memory of Major Thomas J. Corley (1909-1985)

Requıēscās in Pace, Praeceptor Meus


#### Abstract

The work described m this thesis is based on a detailed analysis of the classical and quantum non linear dynamics of a kıcked oscillator This system belongs to a class of kıcked physical systems (time dependent Hamltonaans) whose dynamics have unversal properties We begin the analysis by considering the classical mapping (recursive relationship) derived from the parent system equations The analysis covers the system's phase space and its evolution as parameters are changed The detaled orbit structure is obtaned and the break-up of this orbit structure in the phase space, influenced by presence of periodic orbits, is examined thoroughly We also show the exastence of two types of orbital diffusion (normal diffusion and a resonance enhanced diffusion) The results from this classical analysis are then compared with the quantum mapping The complexity of this quantum mapping is considerable but, with some necessary numerical considerations, we have used it to generate the tıme evolution of the quantum probability amphtudes of the system's eigenfunctions These amphtudes permit the calculation of the system's energy as time progresses and enable us to compare the quasi-phase space given by the Wigner distrubution with the classical manifold structure to check for scarring of the quantum wavefunctions The quantum mapping we derive has not been defined in any of the literature so that all the results obtained in the quantum regime are original In the classical regime our work on penodic orbits and resonance enhanced diffusion is also original We have adopted some technqques and methods from other kicked systems and modified them for our system to complete the investigation of the kicked oscillator


## ACKNOWLEDGEMENTS

The work contaned in thesis cannot be justly attributed solely to myself but also to all those who in some way or other made my tume as a post-graduate so much more interesting than it would have been without them I don't like presenting lists of who did thus and that etc, etc What I want to present here is a collection of the messages I would like to pass on to those who shared these last five years with me, both in D C U and St Patrick's College, Maynooth and also to those without whom I would never have started or finshed this work

Thus thesis is dedicated to one person who unfortunately will never see it The measure of one's greatness is not how great they consider themselves but how fortunate people think themselves to be having known such a person This can never be more true in the case of Major Thomas J Corley without whom the hves of many would have been so very much duller and without whom I would never have had the opportunity nor the ability to complete this research I hope your trust m me was not ill-placed

Laughter is such a definitive tonuc that if it was possible to capture its essence and bottle it would most certanly be called The Carter Girls The potency of this essence is without equal as can be judged by the effect these two people have on all around them when they get started I can only say this you two have cheered me up on so many occasions that if a laugh was worth money I'd owe more than most developed countries You two are fantastic and I love you both for all you've given me over the years Cheers'

To my family, close relatives and friends I give my thanks for the times I've been encouraged and comforted by you Its good to know there's always a place called home waiting for you with those you love willing to help when times are bad or joimng in the laughter when tımes are good If I had it all to do again I wouldn't change a thing Its from the mistakes I've made that I've learned so much more about people, life and loyalty than I would have known if I hadn't made any in the first place To be welcomed with open arms and not judged prematurely is surely a most blessed gift and one which I know is there for me at home

Life is all about taking risks and chances It was such when Professor Danel

Heffernan took me on to do research for hum back in '89 If it wasn't for his willingness to take a chance with me and take me on as a post-grad then the past five years would not have unfolded the way they did and my hfe would certanly be very much different now than is the case It $2 s$ only fitting that I take this opportumity to personally thank Professor Daniel Heffernan for his foresight in giving me this chance to do some real physics Thanks Danny for all the help and gudance you've given me

I was not the only person in the same boat when a post-graduate in D C U and it is to my fellow past (and present) colleagues that I dedicate this paragraph The number of people I have come to know whule a post-graduate precludes my presenting a complete list here so I'll just thank you all for the good tımes and the bad There are a few I would like to single out for special mention To Billy, for the laughs, strategy games, good conversation and all the bottles of Muller you owe me, thanks' The last few years wouldn't have been near as good as they have been without you To Catherine for all the laughs and good chats we've had To the staff of the School of Physical Sciences in D C U , especially Prof Eugene Kennedy, and the staff in the Dept of Mathematical Physics in Maynooth I want to thank you for all the help you've given me in the past I also want to single out Dr Charles Nash for giving me all the help he has both with the administration of the network in Maynooth and his help in unravelling the compexties of TEX Finally to all those, too numerous to mention (so I won't), thank you for the past five years

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Finally, if there's anyone else out there I haven't mentioned or who doesn't fall into any of the above categories TOUGH'

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## OVERVIEW OF THESIS LAYOUT \& CONTENTS

The work presented in this thesis is in two parts, the classical regime of the kicked harmonic oscillator and the corresponding quantum regime The classical regime is presented first and the results from the detailed analysis are presented in chapters 3 \& 4 We then proceed to the more abstract fully quantum regime and compare the results obtaned, in chapters 6 \& 7 , with those in the earher chapters from the classical regime Our intention is to relate the dynamics \& phase space structures in the classical regime to the evolution of wavefunctions, energes \& qausi-phase space distributions in the fully quantum regime Consequently, we will illustrate how remnants of the classical regime are visible in the fully quantum regime without us having to explore such limits as $\hbar \rightarrow 0$

The principle reasons for our choosing the kicked harmonic oscillator are that this system is physically real (and not mathematically motivated) and classically chaotic Furthermore, it has been postulated by Berman et al (1991) that the kucked rotator, which is known to have suppression of chaos in the quantum regime, is not a genenc kucked system They argue that the presence of an extra time scale in the kicked oscillator can give rise to wavefunction de-localisation and hence non suppression of chaos in the quantum regime

In chapter 1 we review the area of Hamiltoman chaos which is pertment to the research carried out by us The chapter also introduces the physical signficance of the system we examune (Chernikov et al (1987)) and shows how the equations governing the system are obtaned Some definitions are included at the end of the chapter of the more important expressions and terms used in the later analysis

Chapter 2 consists solely of the derivation of the classical mapping from the Hamiltoman describing the kicked harmonic oscillator We proceed by expressing the Hamultoman in terms of the systems anmiblation \& creation operators with no specific kick potential is given Our reason for this is to show the generic nature of the mapping thus obtaned through comparison with that obtained Goggin et al (1990) for their formulation of the quantum logistic map We then proceed with a specific potential so that the system corresponds to the physical system presented in section 13 of this chapter (Chermkov et
al (1987)) Once the final expression is arrived at, the mapping is transposed to from a quantum operator regime to a classical scalar regime and the final classical mapping is presented once the real and complex terms of the expression are separated The special case of a resonant kucking, $\imath e$ when the kıck period is commensurate with the natural period of oscillation of the free harmonc oscillator, is highlighted specifically as this is the case we are primarly considering Some phase space portraits are also included for the purpose of illustrating the diversity of dynamic behaviour present

The analysis of the derived classical mapping takes place in chapters 3 \& 4 In chapter 3 we examine the structure of the phase space and use the subsequent results to haghlight the mechanisms of orbit travel in the diffuse stohcastic layer withn this phase space We also obtain an expression which allows us to predrct the orders of penodic points existing on the layer boundary and show using the KAM theorem how these affect the break up of stable orbits in the invariant cells in the phase space which define the stable motion of the system Due to the fundamental mportance of the stochastic layer in defining the non-linear properties of our kicked system, we examine the layer itself specifically its width variation as a function of the kuck strength and the deffusion of orbits within the layer itself This latter examination is presented in chapter 4 where we undertake an analysis sımular to that of Rechester \& White (1890) We examune the effect of correlations, between iterates, on the daffusion of orbits and we modify the approach of Rechester \& White so that it can successfully model the behaviour of our kicked system We also highlight how strong correlations between iterates can lead to a resonance effect which manufests itself as spikes on the diffusion curve A comparison is made between our expressions for this enhancement and the predictions of Ishizakı et al (1989) for their investigation of a sımular phenomenon for the kicked rotator

The derivation of the quantum mapping signufies the beginning of the quantum analysis of our kicked oscıllator The approach we take is sımılar to Fox et al ((1990), (1994)) which maps the probabllty amphtudes of the undriven oscillator's elgenfunctions from knck to kack This allows us to construct wavefunctions, energy evolutions, phase space distributions and quasi-energy levels The approach uses a fixed evolution matrix which multiphes a vector, consisting of the amplitude strengths of one time step, to give
a vector consisting of the amphtudes at the subsequent time step Thus approach allows all the analysis to be carried out, for a given set of parameter values, by just calculating a single matrix However, for each parameter change, a new matrix must be evaluated The derivation proper and the numerical short-cuts we considered when computing these evolution matrices are included in chapter 5

The results from our quantum mapping are analysed in chapter $6 \& 7$ We take a three-tiered approach to the analysis, namely a neghgble kick strength input, a moderate kıck strength input and a large kick strength input Our reasons for thes are made clear in chapter 6 The vast choice of initial distributions of probability amplitude strengths is overcome by considering two types a pure state of the undriven oscillator and a mixed state consisting of a weighted distribution of the undriven oscillator's eigenfunctions These plus our three-tiered approach constitute the scope of our quantum analysis inputs In chapter 6 we examine the amplitude distribution as it evolves in tıme for our three kick strengths, the energy evolution for the corresponding cases and the relevant wavefunction evolution We examine the diffusion of orbits which exhubit a linear energy increase and also the quasi-energy levels manifested in some of the results as steady state levels in the energy evolution curves The conclusions drawn are compared with the quasi-phase space portraits presented in chapter 7

Chapter 7 represents the final stage in our analysis We present, initially, a deriva$t_{10 n}$ of the Wigner distribution (Wigner (1932)) based on the equations given in the review by Hullery et al (1984) The resultant expression is also shown to be real valued despite the appearance of complex terms The latter half of the chapter is devoted to several parss of figures, each consisting of an image of the wigner distribution above the corresponding contour plot of the same distribution Our intention is to illustrate how the system evolves temporally in its position-momentum space and compare it to the classical results The results of this are as startling as they are beautiful Our main conclusions are summarised in chapter 8 where inter-comparisons are made and future paths, along which research could proceed, highlighted

## CHAPTER 1 <br> HAMILTONIAN CHAOS

The pursuit of an understanding of the mechanisms behind prevously indeterminable processes has been made possıble with the development of fast, rehable, cost efficient sold state electronic computers Today, through the use of computers, both the scientific and engineenng communties are benefiting from the research into controlling the non-linear behaviour of critical systems, the onset of turbulence and chaos This chapter will attempt to review briefly the main developments in non-hnear dynamics which are relevant to the work being presented in this thesis Many books exast which will give the reader a much more detaled picture of this whole area of science (Cvitanovic (1989), Bergé et al (1987), Devaney (1989), Ott (1993))

## 11 Introduction

Historically, scientists, mathematicians and, more recently, engineers have attempted to impose a simplified picture on our perceptions of the world about us For hundreds of years the nature of everything about us has been explaned away as modrfications on a theme of linearity in everything from maths to astronomy it was using Euchdian geometry that astronomers \& phlosophers back in antiquity tried unsuccessfully to model the Universe with the earth at its center using a series of concentric circular orbits for the moon, planets \& sun (Ptolemanc System) Even the most complex of systems have had attempts made on them to be reined $i n$ by this predulection to linearity and to be brushed off as impossible when proved unsuccessful Yet even despite this blinkening there were some observant enough to see something hidden in the nature of these systems, a complexaty which defied the simple, artificial rules that were used to explain them and flaunting a rich beauty beyond the realms of the mediocrity of simple linearity Leonardo Da Vincı humself sketched the flow of water coming from a fountain in Florence in the $16^{\text {th }}$ century and was observant enough to include the small structure eddy currents within the larger ones as he sketched the water spewing out This structure within structure is what makes everything around us so interesting the fact that something could be there which
we've overlooked Non-lineanty is the nature of our being and chaos the theory behind that nature

In classical systems chaos is considered present if the system exhibits a sensttive dependence on mitial conditions What this means is that any two mitial conditions separated by even by an infintesimal amount diverge exponentially with time it was Hadamard (1898) who discovered this exponential divergence of nearby trajectories for a particle flow over a surface of negative curvature Not long after, Duhem (1906) and Poincaré (1908) determuned that such a property in any system precludes the making of long term predictions about nature of the system's dynamics The scene was thus set for chaos All that was requred was the proper impetus to bring this work out of the abstract realm of mathematics and to a wider audence (Heffernan et al (1992) \& the references quoted therein)

It was not until the invention of the electronc computer in the 1950s and, more importantly, the further developement of these into the sold state transistorised computers of today that numerical analysis and modelling took off in a big way Many researchers then looked to using these number crunchers as means of evaluating the complex differential equations and iterative mappings It was Lorenz (1963) who, while modelling fluid flow in a simplfied atmospheric model, found the sensitive dependence on inatial conditions Smale (1967) showed that this complex nonhnear behaviour existed in a number of systems However, it was with the advent of such people as Feigenbaum $(1978,1979)$ and his scaling numbers (to explain the structure in a simple iterative mapping on the unit interval), Mandelbrot (1982) and his fractals and Ruelle \& Takens (1971) with their strange attractors that this area of research rapidly took off with the litany of chaoticians including such prominent others as Pomeau, Vıdal, Manneville, Ott, York, Greobogı, Proccacia, Jensen and so on The vast number of pubhcations on this subject over the past decade and a half is as starthng as the area is encompassing The whole dynamicism of nature itself seems contanned within the influence of chaos, everything from the fluid dynamics (Libchaber and Maurer (1982)) to organic cardac cells (Guevara et al (1981)), from non linear electronic crrcuts (Testa et al (1982)) to optical systems (Arrech et al(1982)) and chemical reactions (Hudson and Mankın (1981)) All of the above systems are, of course, classical and easily
observable in experiments In fact the experments are so vaned that the time-scales and length scales of many of the chaotic systems analysed in the classical limit vary over many orders of magnitudes, from munutes for the Libchaber \& Maurer experiment to tens of mucro-seconds for the Testa et al experiment Classical systems were the obvious choice for examung chaos as these apply directly to the world about us and are in the same ballpark when it comes to our normally observable lives yet there were those who wondered about the quantum limit and the classical-quantum correspondence and this led some to ask What happens on the quantum level when the classical system is chaotic or turbulent ${ }^{\rho}$ It was this thought about the quantum limit and how classical chaos mught manifest itself in this hmit that led to the birth of what is now referred to as Quantum Chaology

## 12 Conservative Systems \& Chaos

The area of quantum chaos is quite different to most of the systems mentioned above because quantum systems are described by within the Hamiltonaans framework and therr phase spaces are area-conserved This property forces any analysis of such systems along different routes to the dissipative systems discussed in section 1 The previous systems were considered dissipative because structures in their phase spaces tend to contract and tend to be attracted to tıme invariant haghly complex structures called strange attractors These are usually fractal with the dynamics of the systems being confined to the region of phase space where these attractors reside In our conservative Hamultonian systems we do not have such dissipation or strange attractors present Any structure retains its area in the multi-dimensional phase space of the system We will restrict our arguments to the $1 D$ case as we will only be dealing with a 1 Dimensional oscillator in what is to follow in later chapters

As we have already mentioned, the system's dynamics are determined by its Hamiltoman, $H(\mathbf{p}, \mathbf{q}, t)$ The $\mathbf{p}$ describes the momentum and $\mathbf{q}$ describes position and both are normally vectors of the same dimensionality as the number of degrees of freedom in the system, $N$ As $N=1$ for our system the the vectors $\mathbf{p} \& \mathbf{q}$ will be replaced by the scalars
$p \& q$ The time evolution of $p$ and $q$ is described by Hamulton's equations whech relates the trajectory $p(t) \& q(t)$ trace out $m$ the $2 D$ phase space of our system

$$
\begin{gather*}
\frac{d p}{d t}=-\frac{\partial H(p, q, t)}{\partial q}  \tag{array}\\
\frac{d q}{d t}=\frac{\partial H(p, q, t)}{\partial p} \tag{array}
\end{gather*}
$$

If we form a vector $\tilde{x}$ from the $p \& q(\operatorname{Ott}(1993))$ ie

$$
\overline{\mathbf{x}}=\binom{p}{q}
$$

we can write the two Hamilton equations as one vector equation

$$
\begin{equation*}
\frac{d \tilde{\mathrm{x}}}{d t}=\mathbf{F}(\overline{\mathbf{x}}, t) \tag{array}
\end{equation*}
$$

The advantage of thes is not just the ease of use of one equation but that we can, by taking the divergence of $\mathbf{F}$, show how Hamlton's equations predict the conservation of area in our $2 D$ phase space,

$$
\begin{equation*}
\frac{\partial}{\partial x} \mathbf{F}=\frac{\partial}{\partial p}\left(-\frac{\partial H}{\partial q}\right)+\frac{\partial}{\partial q}\left(-\frac{\partial H}{\partial p}\right)=0 \tag{124}
\end{equation*}
$$

Furthermore, it can be shown (Ott (1993)) that for any closed curve in the $2 D$ phase space, $S_{o}$, the time evolved curve $S_{1}$ encompasses the same area as does $S_{o}$ This follows from differentiating wrt time the area integral and showing it to be zero

$$
\begin{equation*}
\frac{d}{d t} \int_{S_{1}} d^{2} \overline{\mathbf{x}}=\oint_{S_{1}} \frac{d \tilde{\mathbf{x}}}{d t} d S=\oint_{S_{1}} \mathbf{F} d S=\int_{S_{1}} \frac{\partial}{\partial \overline{\mathbf{x}}} \mathbf{F} d^{2} \tilde{\mathbf{x}}=0 \tag{array}
\end{equation*}
$$

This property of having a phase space incompressibilty, Liouvilles Theorem, means that attractors do not exist for Hamiltonian systems in the same sense they do for dissipative classical systems with phase space contraction We have deliberately laboured this point so that this difference is clear Despite the area conservation of the phase space, complex structures can (and do) exast there We will show later that, using the K A M theorem, the break up of orbits in our $2 D$ phase space can become very complex, even fractal, in that the structure has detalled structure even at arbitranly small length scales

The absence of strange attractors with infinitely complex structures, caused by successive foldings and stretchings of the phase space, in classically conservative systems
requares the exastence of another mechansm to give nise to the highly non-hnear behaviour that is expected from a system which is classically chaotic The incompressibility of the phase space also allows the system to occupy the whole of the phase space This prompts the question can some orbits diffuse around some (or all) of the whole phase space and exhzbit sensative dependence on inztzal condztions in a manner analogous to the behaviour of orbits on a dissipative strange attractor ${ }^{p}$ The answer is yes if the system has a surface layer which spreads out over part of the phase space and if this layer permits orbits to diffuse over it The exastence of this layer gives rise to what we now term Hamiltonzan or Conservatzve Chaos In fig 11 we illustrate such a layer which exasts in the phase space of our driven harmonic oscullator The layer is called stochastic because of the diffuse nature of the orbits within it One can see from the figure that the layer occupies the space between zslands of bounded motion in the phase space Any orbits within these islands cannot leave them and are thus bounded in energy, momentum and position The orbits present in the stochastic layer are not so bounded and can diffuse freely over the entire $2 D$ phase space since the web like layer connects all inter-1sland spaces in the phase space (a more quantitative explanation is presented in chapter 3)

The parameters used in fig 11 are explaned in later chapters and the figure is used here for qualitative purposes only, the quantitative analysis will be undertaken in chapters $3 \& 4$ for the classical mapping So for classical conservative systems the exastence of thus layer permits complex non-linear dyanamics to exist The quantum regime is so different from the classical that methods are used which for the most part do not resemble those in the classical regime Energy is a big factor in determining how diffuse an orbit is becoming and so, as energy levels exist m the quantum, it makes sense for energy and energy levels to be scrutimsed in depth Some, such as Berry \& Robnk (1984) and Lewenkopf (1991), have analysed the energy level statistics of tıme independent Hamltonuan systems to check for a possible phase change when the corresponding classical limit is going from regular to chaotic motion Others, Israelev (1990), tend to construct quasienergy levels for kicked systems and analyse these levels with a hope to uncovering some of the mystery surrounding quantum chaos. Both are, to date, inconclusive, Lewenkopf (1991) proved that exceptions can exast to the change m energy level statistics proposed


Fig 11 Plot of the diffuse (or stochastic) layer for the kicked harmome oscullator for $\mu_{c l}=65, K=01 \& \beta=\frac{\pi}{2}$ The meaning of the above parameters is explained in chapters $2 \& 3$
as a mechansm for hughlighting chaos, but work is continuing The evolution of energy in time (in time dependent Hamultomans) and the subsequent comparison with the classical limit is also used This technique was apphed to the kucked rotor (Casatı et al (1979)) and shown to have a natural saturation in the energy Though indicative of suppression of quantum chaos in the kecked rotator, this type of energy saturation is not unversal in quantum systems In fact Berman et al (1991) have postulated that the presence of an extra time scale ( $\alpha \ln \left(\hbar^{-1}\right)$ ) in the quantum regime of the kicked harmonic oscillator mught allow for the non suppression of chaos and that a quantum analogue of the classical crystalline phase space structure mught allow for the delocalisation of wavefunctions allowing the kucked harmonic oscillator to be chaotic in the quantum regime (Daly \& Heffernan (1994)) In chapter 6 of thus thesis we present results which appear to show the energy increasing linearly without saturation indicating non-suppression of chaos Our matrix size and the limit of computing power avalable limit our time evolution but the evidence points to no suppression of chaos in the quantum limit

For quantum systems trajectories do not exast in the manner that we are accus-
tomed to dealing with in the classical regime because of the uncertainty principle Therefore a phase space as such cannot be constructed because of the uncertanty in trajectories We can however construct a quasi-phase space using the distribution introduced by Wigner (1932) and further developed by Cohen (Hillery et al (1984)) Thus Wigner distribution sets up a probability measure over the momentum-postion space allowing contour hnes to be drawn joining regions of equal probabilty Thus one can construct a picture of the probability distribution in the Q-P plane and thus a quasi-phase space for these quantum systems (Hannay \& Berry (1980)) To more readuly compare the quantum and classical phase spaces the Husimu distribution (Husim (1940)) is often invoked This distribution is a coherent state representation of the Wigner distirbution (c $f$ chapter 7 and the review article by Hillery et al (1984)) As the coherent state has the least uncertannty of any wavefunction the hope is to more closely compare the two phase spaces This is the more common distribution encountered in the literature (Balazs (1990), Radons \& Prange (1990), Scharf \& Sundaram (1991), Kus et al (1991) and Scharf \& Sundaram (1992))

The discrete nature of the kıck used in penodically kıcked systems permits the construction of mappings relating the system's behaviour from one kuck period to the next For some of the more simple systems, the kucked quantum rotor (Scharf \& Sundaram (1991)) or the Baker's map (Balazs (1990)), the methods used appear simple enough because of the simplicity of the elgenfunctions of the systems However for the more complex systems, such as our kicked harmonic oscullator, the more direct method of relating probablity amplitudes, of the undriven eigenfunctions, from kick to kıck (Fox \& Lan (1990), Fox \& Elston (1994)) is used with success We consider this latter technqque more realistic as it relates physical quantities (the probability amplitudes of the eigenfunctions of the undriven harmonic oscillator) to the tıme evolution of the kicked system It will become apparent from the results in chapters $6 \& 7$ that from just the evolution of the amplitudes we can calculate the energy evolution, the quasi-energy levels \& the Wigner distribution all as functions of time The usefulness of thus technique depends greatly, however, on the eigenfunction of the system itself and the existence of anlaytical solutions to integrals involving such functions (cf chapter 5 eq (5126)) We end with the comment that each system has its own characteristics and it is by fundamentally understanding these that the
right method for quantising the Hamiltonian can be found
The last few paragraphs were intended to highlight the differences between the quantum regime and the more famuliar classical regime The conceptual difficulties assocrated with quantum phenomena are not eased by the complex non-hnear nature of these systems but we hope that we have answered more questions in this very brief discussion of quantum chaos More detaled discussions (ones which do the subject more justice than could be attempted here) can be found m such books as Quantum Chaos edited by Cerderra, Ramaswamy, Gutzwiller \& Casatı and Edward Ott's Chaos in Dynamical Systems We complete this introductory chapter with a brief outhne of some of the nomenclature used in this work and how they are defined in the context of this work

## 13 The System Proper \& Some Definitions

The kacked harmonc oscillator, in the form we present its Hamiltonian in chapters $2 \& 5$, can be motivated from physical systems Chernikov et al (1987) chose that of particle motion in an external magnetic field $B_{o}$ with a disturbing plane wave packet present which propagates orthogonally to the magnetic field (which propagates along the $x$ axis) The plane wave packet contans a very large number of plane waves

$$
\begin{equation*}
\frac{d^{2} x}{d t^{2}}+\omega_{o}^{2} \sin (x)=\frac{e}{m} E_{o} \sum_{m} \sin \left(k_{m} x-\omega_{m} t\right) \tag{array}
\end{equation*}
$$

where $\omega_{o}=e B_{o} / m c$ is the cyclotron frequency By setting $k_{m}=$ const $=k_{o}$ and $\omega_{m}=$ $n \Delta \omega$ with $\Delta \omega$ equal to the frequency separation of neighbouring modes and $n=0, \pm 1$, we find that, as the number of modes $m$ tends to infinity, the summation term can be spht to give

$$
\begin{equation*}
\sum_{m} \sin \left(k_{m} x-\omega_{m} t\right)=\sin \left(k_{o} x\right) \sum_{n=-\infty}^{\infty} \cos \left(\frac{2 n \pi t}{T}\right) \tag{array}
\end{equation*}
$$

The summation over $n$ of the cosines gives a periodic train of delta pulse so that eq (131) becomes a more familiar equation

$$
\frac{d^{2} x}{d t^{2}}+\omega_{o}^{2} x=\frac{e}{m} E_{o} \sin \left(k_{o} x\right) \sum_{n=-\infty}^{\infty} \delta(t-n T)
$$

From thus point it is quite starightforward to reconstruct the system Hamultonian Noting that $d q / d t=-p / m$ and that $d^{2} q / d t^{2}=-\left(\frac{1}{m}\right) d p / d t$ we find that by replacing $x$ with $q$ we obtain the time evolution equations of $p$ and $q$

$$
\begin{gather*}
\frac{d p}{d t}=m \omega_{o}^{2} x-e E_{o} \sin \left(k_{o} x\right) \sum_{n=-\infty}^{\infty} \delta(t-n T)  \tag{134a}\\
\frac{d q}{d t}=-\frac{p}{m} \tag{134b}
\end{gather*}
$$

where, upon invoking Hamilton's equations (121) \& (12 22 ) we end up with the hamultonuan we require for our analysis and that which is used from now on when referning to this system

$$
\begin{equation*}
H(p, q, t)=\frac{p^{2}}{2 m}+\frac{1}{2} m \omega_{o}^{2} q^{2}+\frac{e E_{o}}{k_{o}} \cos \left(k_{o} x\right) \sum_{n=-\infty}^{\infty} \delta(t-n T) \tag{135}
\end{equation*}
$$

Some of the more commonly used expressions and words in any area of research can very rapidly become almost colloqual so that their original intended definition or function takes on a much broader term or in some case is used to refer to something only vaguely related to the orginal meaning To prevent any misunderstanding here we will make a few loose definitions to essentially highlight the meaning we intend in this thesis for the following terms

Integrability When we refer to integrablity in this system, the kicked harmonic oscillator, we mean that the Hamltonıan, $H(p, q)$, can be expressed, by means of a canonical change of variables $(p, q) \rightarrow(J, \theta)$, solely in terms of one of the variables in our system the introduction of the action-angle variables means the undriven harmonc oscillator can be expressed solely interms of $J$ (the action variable) and is therefore integrable The kicked system cannot be so expressed and is therefore non-integrable The analysis involving integrability is carned out in chapter 3
K.A M. Theorem As we understand this very complex theorem, the tori (in our $2 D$ phase space these are just orbits) in the phase space are subject to distortion and eventually break up if the rotation number of the orbit on this torus is sufficiently commensurate with the perzodicity of the perturbing term This is how we apply the theorem to our system In


Fig 12 Unperturbed circular orbits of a typical integrable system The dashed orbits are the two mentioned which have rotational number of 3 \& 4 and are among the first orbits to break up under K A M From Chaos in Dynamical Systems by E Ott, Chapter 7

$F_{2 g} 19$ The perturbed orbits corresponding to fig 12 The two commensurate orbits have broken into hyperbolic \& elliptic points of the order of the rotational number From Chaos in Dynamical Systems by E Ott, Chapter 7


Fig $1 \&$ The two different types of fixed points shown here in the phase plot of the pendulum, $\ddot{\mathrm{x}}=-\sin (x)$ The separatrix is the line separating bound motion (oscllation) and unbound (rotation) From "Self.Generated Chaotzc Behavzour in Nonlinear Mechanics" by $R$ H Gelleman in Universality ${ }_{2 n}$ Chaos edited by $P$ Cvitanovic
fig 12 we show the circular type orbits for an integrable system (as a surface of section of a torus) and this can be compared to the case in fig 21 for our system with no kicking (1 e integrable) In fig 13 we show how some orbits of rotation numbers $3 \& 4$ have broken up into two sets of four points (one set hyperbohc, the other ellhptic) The elliptic points of these broken up orbits have a circular structure simular to the original integrable system and these too experience break up in a simlar manner resulting in a highly complex structure for large kicking This is discussed in detal in chapter 3

Periodic Fixed Points What we want to discuss here is the two types of periodic fixed points encountered in the type of conservative system we're considering The first is the elluptic fixed point This is distingushed by its complex eigenvalues and by the fact that nearby points are rotated by the presence of this point into ellipses (or circles) Therefore the presence of circular concentric orbits in our phase space highlights the presence of an ellhptic fixed point at the centre of these concentric orbits Hyperbolic fixed points have real eigenvalues and are distingushable by the fact that orbits near them can be attracted
towards them along their stable directions (given by the eigenvalue less than one) and repelled along the unstable directions (given by the eigenvalue greater than one) Orbits also generally have the shape of hyperbolae near such fixed points A schematic of the two types is given in fig 14

This introductory chapter is just a brief review of the areas relevent to the work in this thesis and a more in depth introduction can be found in the reference books previously mentioned We will now discuss $m$ detaul the physics of the non-lnearly kıcked harmonic oscillator

## CHAPTER 2 THE CLASSICAL MAPPING

This chapter is intended as an introduction to the classical mapping which we obtained from the quantum mechancal equations describing the kicked harmonc oscillator The ideas and methods by which we obtain the final form for the mapping are outhned along with a brief description of the the derivation itself To lllustrate the complexity of the resultant mapping, we include some phase space portraits along with a plot obtaned from the differential equations of the system for comparison

## 21 The Classical Derıvation

As the title of this section suggests, our purpose here is to derive the classical mapping from the quantum operator equations The complete derivation is quite long and tedrous so only those steps that are necessary for cohesion will be presented Furthermore, as the undriven system is a standard under-graduate problem, we will assume the reader is familar with the operators, $a \& a^{\dagger}$, and with the basic properties of the undriven system A more complete introduction can be found in any standard quantum mechanics book such as that by Messiah (1976) It is the classical map, resulting from this derivation, which we will use extensively in the classical analysis of the kicked quantum harmonic oscillator undertaken in chapters three and four

In order for us to proceed with this derivation it is necessary to explann briefly how we intend, here, to transpose from the quantum to the classical In quantum space the exastence of wavefunctions necessitates the exastence of operators to operate on these wavefunctions and obtain measurable quantities (once the expectation value is sought) In classical space these operators become simple scalars or, using a more formal notation, we simply take the $c$ number representation for any operators used If some operator function, $f\left(a, a^{\dagger}\right)$, represents the system in the quantum limit then a scalar function $\bar{f}\left(\alpha, \alpha^{*}\right)$ can be defined to represent the system $m$ the classical lumit where the $a \& a^{\dagger}$ are operators and form a Hermitian conjugate pair with the $\alpha \& \alpha^{*}$ being scalar and forming a complex conjugate pair It is by this transformation that the Hermitian conjugate operators become
simple complex conjugate numbers and hence how a classical system can be formed from a given quantum system (Lousell (1965))

Let us now introduce the system to be analysed It consists of a quantum harmonc oscllator, of mass $M$ and natural frequency $\omega_{o}$, which is driven by a potential $V(p, q)$ whose temporally discrete nature is brought about by its product with a periodic delta function Thus the Hamiltonaan, $H_{T}$, for this system is the sum of the Hamultonan for the undriven harmonc oscillator, $H_{o}$, and the driving (or kacking) potential term $H_{1}$ The form of the potential , $V(p, q)$, is critical for the overall system's behaviour and its exact form will be given later so as to keep the following derivation as generic as possible

The Hamultoman, $H_{o}(t)$, for the ssolated quantum harmonic oscullator is given below in eq (211)

$$
\begin{equation*}
H_{o}(t)=\frac{p(t)^{2}}{2 M}+\frac{1}{2} M \omega_{o}^{2} q(t)^{2} \tag{array}
\end{equation*}
$$

where $p(t)$ signfies momentum and $q(t)$ position and both are functions of time, $t$ (Messiah, 1976) As we will demonstrate in chapter 3, this system is integrable

To account for the kicking, the driving term, denoted $H_{1}(t)$, is added to $H_{o}(t)$, forming the hamultonian $H_{T}$ for the complete system, and the form that this kicking takes is given in eq (2 12 )

$$
\begin{equation*}
H_{1}(t)=V(p, q) \sum_{n=-\infty}^{\infty} \delta(t-n \tau) \tag{array}
\end{equation*}
$$

As can be seen from the equation above, $H_{1}(t)$ is the product of a continuous potential $V(p, q)$ with a periodic train of discrete delta pulses The addition of $H_{1}(t)$ not only makes an integrable system, $H_{o}(t)$, non-integrable but also allow highly complex dynamics to exast in the system even when $H_{1}(t)$ is more of a perturbation than a dominant term It is this property that makes this class of system particularly interesting

To obtain the classical mapping for this system it is necessary, and preferred, to express the above equations in terms of a set of anmhlation and creation operators, given by $a(t)$ and $a^{\dagger}(t)$ respectively Those famulhar with the isolated harmonic oscillator will be aware that this substitution of $a(t)$ and $a^{\dagger}(t)$ for $p(t)$ and $q(t)$ makes the hamultoman $H_{o}(t)$ much more manageable and this property is apphed to the compound hamiltonian
$H_{T}(t)$ to sımplify it As an intermediate step in the substitution the operators, $P(t) \&$ $Q(t)$, are defined by

$$
\begin{equation*}
p(t)=P(t) \sqrt{M \hbar \omega_{o}} \quad q(t)=Q(t) \sqrt{\frac{\hbar}{M \omega_{o}}} \tag{array}
\end{equation*}
$$

By then defining the creation \& anmbulation operators as

$$
\begin{equation*}
a(t)=\frac{1}{\sqrt{2}}(Q(t)+\iota P(t)) \quad a^{\dagger}(t)=\frac{1}{\sqrt{2}}(Q(t)-\iota P(t)) \tag{214}
\end{equation*}
$$

the Hamiltoman, $H_{o}(t)$, for the isolated oscillator, and the Hamiltonan $H_{1}(t)$ for the driving term become

$$
\begin{equation*}
H_{o}(t)=\frac{\hbar \omega_{o}}{2}\left(a(t) a^{\dagger}(t)+a^{\dagger}(t) a(t)\right) \tag{215a}
\end{equation*}
$$

and

$$
\begin{equation*}
H_{1}(t)=V\left(a(t), a^{\dagger}(t)\right) \sum_{n=-\infty}^{\infty} \delta(t-n \tau) \tag{215b}
\end{equation*}
$$

respectively with the complete system Hamiltoman, $H_{T}(t)$, being just the sum of eq (215a) with eq (215b), 1e

$$
\begin{equation*}
H_{T}(t)=\frac{\hbar \omega}{2}\left(a^{\dagger}(t) a(t)+a(t) a^{\dagger}(t)\right)+V\left(a(t), a^{\dagger}(t)\right) \sum_{n=-\infty}^{\infty} \delta(t-n \tau) \tag{216}
\end{equation*}
$$

With the system hamiltonan thus expressed we will now proceed to examine the time evolution of the operators, $a(t) \& a^{\dagger}(t)$, so that a recursive mapping can be obtained The tıme evolution of any operator, $\Psi$, can be described by the equation

$$
\begin{equation*}
\frac{\partial \Psi}{\partial t}=-\frac{\iota}{h}(\Psi, H) \tag{array}
\end{equation*}
$$

where $\Psi \equiv \Psi(t)$ is time dependent The brackets, (), denote the commutation of any operators contaned theren As it is necessary to commute operators to obtain their time evolution then it is appropriate to review some commutation rules which are used to great advantage in the simplification of the equations resulting from the apphcation of eq (217)
$1(a, a)=0 \quad$ (any operator commuted with itself is zero)
$2 \quad\left(a, a^{\dagger}\right)=1, \quad\left(a^{\dagger}, a\right)=-1$
$3 \quad(a, b c)=(a, b) c+(a, c) b$
$4 \quad\left(a, b^{2 n}\right)=\sum_{s=0}^{n-1} b^{s}(a, b) b^{n-s-1}$
The a and $a^{\dagger}$ in the four equations above are of course functions of tıme, $a \equiv a(t)$ and $a^{\dagger} \equiv a^{\dagger}(t)$ For $a(t) \& a^{\dagger}(t)$, eq (218d) simplfies to $n b^{2 n-1}$ because the commutation of $a(t)$ with $a^{\dagger}(t)$, and vice versa, is scalar, $c f$ eq (218b)

The time evolution of the operator $a(t) 15$, using rules $1-4 \mathrm{in}$ eqs (218a) to (218d) and eq (217)

$$
\begin{equation*}
\frac{\partial a(t)}{\partial t}=-\iota \omega a(t)-\frac{\iota}{\hbar}\left(a(t), V\left(a(t), a^{\dagger}(t)\right)\right) \sum_{n=-\infty}^{\infty} \delta(t-n \tau) \tag{array}
\end{equation*}
$$

For $a^{\dagger}(t)$ a simular equation is obtanned

$$
\begin{equation*}
\frac{\partial a^{\dagger}(t)}{\partial t}=\omega \omega a^{\dagger}(t)-\frac{\iota}{\hbar}\left(a^{\dagger}(t), V\left(a(t), a^{\dagger}(t)\right) \sum_{n=-\infty}^{\infty} \delta(t-n \tau)\right. \tag{array}
\end{equation*}
$$

As the explicit form for the driving potential has not yet been revealed, then for convenence $-\frac{2}{\hbar}\left(a(t), V\left(a(t), a^{\dagger}(t)\right)\right)$ shall be written as $f\left(a, a^{\dagger}\right)$ and $f^{\dagger}\left(a, a^{\dagger}\right)$ for the corresponding $a^{\dagger}$ commutation Putting $b \equiv \iota \omega_{o}$ and by integrating the equation over one kıck cycle from just before the $N$ to just before the $(N+1)$ kıck ( $t=N \tau-\epsilon$ to $t=(N+1) \tau-\epsilon$ ) then, (as we let $\epsilon \rightarrow 0$ ),

$$
\begin{equation*}
a(N+1) e^{(b(N+1) \tau)}=a(N) e^{(b N \tau)}+f\left(a(N), a^{\dagger}(N)\right) e^{(b N \tau)} \tag{2111}
\end{equation*}
$$

Dividing across by $e^{(b(N+1) r)}$, to isolate the $a(N+1)$ term on the left hand side, the resulting equation is

$$
\begin{equation*}
a(N+1)=a(N) e^{(-b r)}+f\left(a(N), a^{\dagger}(N)\right) e^{(-b \tau)} \tag{2112a}
\end{equation*}
$$

with the equation for the $a^{\dagger}(t)$ operator sımular in form to the above except for the substitution of $a(t)$ for $a^{\dagger}(t)$ and a few sign changes

$$
\begin{equation*}
a^{\dagger}(N+1)=a^{\dagger}(N) e^{(b \tau)}+f^{\dagger}\left(a(N), a^{\dagger}(N)\right) e^{(b \tau)} \tag{array}
\end{equation*}
$$

Defining $\beta$ to be equal to $b \tau$, the final genenc forms for the anmhlation and creation operators, $a \& a^{\dagger}$, can now be written These two equations are called generic because
they hold true for any temporally discrete potential of the form given in eq (212) The equations below are remarkably sımlar to those of Goggin et Al (1990) in his formulation of the quantum logistic map This form is quite general and can be apphed to any drıven oscillator system regardless of the kıck term as well as other kıcked systems where an anmhlation - creation type operator set exasts and where this set exhibits the same properties as those for our system

$$
\begin{align*}
a(N+1) & =\left[a(N)+f\left(a(N), a^{\dagger}(N)\right)\right] e^{-\beta}  \tag{2113a}\\
a^{\dagger}(N+1) & =\left[a^{\dagger}(N)+f^{\dagger}\left(a(N), a^{\dagger}(N)\right)\right] e^{\beta} \tag{2113b}
\end{align*}
$$

Despite having obtained such a set of generic maps it is impossible to predict the system's behaviour until a specific form for the potential is given The form of the potential we have chosen is

$$
\begin{equation*}
V(p, q)=\mu_{q} \cos (k q) \tag{2114}
\end{equation*}
$$

after Berman et al (1991) In thus chonce of potential the parameter $\mu_{q}$ is the quantum kuck strength and $k$ is a characteristic length scale The importance of writing the potential like this is to enable a direct comparison to be made between the classical parameters $m$ our current analysis and the quantum parameters in the forthcoming quantum analysis in chapters 5 and 6 This choice is also in keeping with the physcial significance of the system as explained in chapter 1 Upon substituting the operators a and $a^{\dagger}$ for $q$, our equation for the potential becomes

$$
\begin{equation*}
V\left(a, a^{\dagger}\right)=\mu_{q} \cos \left(K\left(a^{\dagger}(t)+a(t)\right)\right) \tag{2115}
\end{equation*}
$$

where for convenience $K$ is defined to be $k \sqrt{\hbar /\left(2 M \omega_{o}\right)}$ Using the trigonometric identity $\operatorname{Cos}(A+B)=\operatorname{Cos}(A) \operatorname{Cos}(B)-\operatorname{Sin}(A) \operatorname{Sin}(B)$ with the properties of the operators, a and $a^{\dagger}$, and the series expansions of the sine and cosine functions, the operator functions $f\left(a, a^{\dagger}\right)$ and $f^{\dagger}\left(a, a^{\dagger}\right)$ both take on the same form

$$
\begin{equation*}
f\left(a, a^{\dagger}\right)=f^{\dagger}\left(a, a^{\dagger}\right)=+\iota \mu_{c l} \operatorname{sm}\left[K\left(a^{\dagger}(t)+a(t)\right)\right] \tag{array}
\end{equation*}
$$

where $\mu_{c l}$ is the classical kick strength and is defined to be $\mu_{q} K / \hbar$ A detanled step by step derivation of how we obtain eq (2 1 16) from eq (2 1 15) is included in Appendux A

It is now possible to construct the quantum operator mapping for a and $a^{\dagger}$ for the chosen potential These two mappings are identical in form to the generic ones defined by eqs (2113a-b) but they are specific in the sense that the general functions $f\left(a, a^{\dagger}\right)$ \& $f^{\dagger}\left(a, a^{\dagger}\right)$ have been replaced by therr specific form as given by eq (2116) The final operator maps are

$$
\begin{align*}
& a(N+1)=a(N) e^{-\beta}+\iota \mu_{c l} \sin \left[K\left(a^{\dagger}(N)+a(N)\right)\right] e^{-\beta}  \tag{2117a}\\
& \quad a^{\dagger}(N+1)=a^{\dagger}(N) e^{\beta}-\iota \mu_{c l} \sin \left[K\left(a^{\dagger}(N)+a(N)\right)\right] e^{\beta} \tag{array}
\end{align*}
$$

It should be pointed out here that the $a(N) \& a^{\dagger}(N)$ constitute a Hermitian conjugate pair In order to transpose to the classical regime we multiply $a(N), a^{\dagger}(N) \& \mu_{c l}$ by a parameter $K_{0}\left(\equiv k_{0} \sqrt{\hbar / 2 M \omega_{0}}\right)$ defined to be equal to one Also $K$ is divided by $K_{0}$ for consistency in the sine argument Thus no parameters or vanables have any exphcit dependence on $\hbar$ We can now transpose to the classical It is possible to express $a(N)$ in the $c$ number representation as $\alpha(n)$ and $a^{\dagger}(N)$ as $\alpha^{*}(n)$ where the $\alpha^{\prime} s$ constitute a complex conjugate pair and are not operators As the $\alpha(n)^{\prime} s$ and $\alpha^{*}(n)^{\prime} s$ are just numbers then we can, quite justıfiably, write $\alpha(n)$ as $x(n)+\iota y(n)$ and $\alpha^{*}(n)$ as $x(n)-\iota y(n)$

Therefore eqs (2117a) and (2117b) become

$$
\begin{equation*}
x(n+1)+\iota y(n+1)=\left[x(n)+\iota y(n)+\iota \mu_{c l} \sin (2 K x(n))\right] e^{-\beta} \tag{2118a}
\end{equation*}
$$

for the transposition of the $a(N)$ operator mapping and

$$
\begin{equation*}
x(n+1)-\iota y(n+1)=\left[x(n)-\iota y(n)-\iota \mu_{c l} \sin (2 K x(n))\right] e^{+\beta} \tag{2118b}
\end{equation*}
$$

for the $a^{\dagger}(N)$ operator mapping
It is interesting to note that the real and imaginary parts of the two above equations are identical We now spht the real and imaginary parts to yeld a mapping for $x(n)$ (the real part) and $y(n)$ (the corresponding imaginary part) The two equations we obtain are, for the real parts of eqs (2 1 18a) \& (2 1 18b)

$$
\begin{equation*}
x(n+1)=x(n) \cos \left(\omega_{o} T\right)+y(n) \sin \left(\omega_{o} T\right)+\mu_{c l} \sin [2 K x(n)] \sin \left(\omega_{o} T\right) \tag{2119a}
\end{equation*}
$$

and for the imaginary parts

$$
\begin{equation*}
y(n+1)=y(n) \cos \left(\omega_{o} T\right)-x(n) \sin \left(\omega_{o} T\right)+\mu_{c l} \sin [2 K x(n)] \cos \left(\omega_{o} T\right) \tag{2.1.19b}
\end{equation*}
$$

where we have replaced the parameter $\beta$ with $\omega \omega_{o} T$ and simplified the equations. These are the equations which constitute our classical mapping as derived from the quantum mechanical operator equations. Their form is different from that normally observed in the literature such as the articles by Chernikov et al. (1989), Israilev (1990), Berman et al. (1991) and others (Schmera et al 1992, Zaslavsky and Filonenko, 1968). The reason for this is our inclusion explicitly of the $K$ in the sine argument whereas most others re-scale their $x(n)^{\prime} s$ and $y(n)^{\prime} s$ such that this term is excluded. Our mapping is equivalent to that in the literature when $K=0.5$. Our inclusion of $K$ allows us to expand the phase space as the system's periodicity is directly proportional to the parameter $K$. This ability to expand the phase space allows detailed analyses to be made of the small scale structures without necessarily having to worry about computer precision. The trade off, however, is the complexity of an extra parameter.

These equations are also dimensionless as the $a \& a^{\dagger}$ operators are themselves dimensionless from their definitions in eqs. (2.1.3) \& (2.1.4) and the transposition to the classical representation adds no dimension to the variables $\alpha$ and $\alpha^{*}$. Attention should also be brought to the fact that the $x$ in the classical mappings can be traced directly back to the $Q$ operator in quantum space and likewise the $y$ to the $P$ : i.e. from eq. (2.1.4) and the definitions of $\alpha$ and $\alpha^{*}$

$$
a=\frac{1}{\sqrt{2}}(Q+\iota P) \rightarrow \alpha=(x+\iota y) \& a^{\dagger}=\frac{1}{\sqrt{2}}(Q-\iota P) \rightarrow \alpha^{*}=(x-\iota y)
$$

so that

$$
\frac{Q}{\sqrt{2}} \rightarrow x \quad \& \quad \frac{P}{\sqrt{2}} \rightarrow y
$$

with $\rightarrow$ signifying here the classical transposition.
Taking eqs. (2.1.19a) \& (2.1.19b) a much simplified mapping is found for $\omega_{0} T=\frac{\pi}{2}$ which is the specific form chosen in the subsequent analysis as this form despite being more
manageable has one less parameter to deal with This form is given below in eqs ( 2120 ) and (2121) and will henceforth be referred to as the Resonant Kıcked Oscllator Map (RKOM)

$$
\begin{gather*}
x(n+1)=y(n)+\mu_{c l} \sin [2 K x(n)]  \tag{array}\\
y(n+1)=-x(n) \tag{array}
\end{gather*}
$$

To test whether the above mappings did in fact describe the behaviour of the onginal Hamultonan, eqs (211) \& (2 12 ) were modelled using the time evolution approach for the $p \& q$ operators, not the $a \& a^{\dagger}$ as before The Hamltoman $H_{T}(t)$ given by the sum of eqs (211) and (212) gives the time evolution equations below for the $P$ \& $Q$ operators (defined by eq (213))

$$
\begin{equation*}
\frac{\partial P}{\partial t}=-\frac{\iota}{h}\left(P, H_{T}\right)=\omega_{o}^{2} Q+\mu_{c l} \sin (K Q) \sum_{n=1}^{\infty} \delta(t-n \tau) \tag{2122}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\partial Q}{\partial t}=-\frac{\iota}{h}\left(Q, H_{T}\right)=-P \tag{array}
\end{equation*}
$$

These were integrated numencally and the results were in agreement with the mappings above Therefore we can confidently state that the dynamics present in the classical mapping are representative of the continuous system

Some phase space plots, from the mappings in eqs (2120) \& (2121) and from the continuous system in eqs (2122) \& (2123), are shown in the proceeding section to illustrate how complex the phase space for this system can be and how the mapping is equivalent to the system's differential equations

## 22 Some Phase-Space Portrats

This section is solely dedicated to illustrating the phase space of the classical mapping derived in the preceding section and as such we do not intend here to analyse the phase space This section is solely demonstrative The portraits chosen are considered


Fig 21 Phase space plot for $\beta=\frac{(1+\sqrt{5}) \pi}{2}, K=01$ and $\mu_{c l}=0,1$ e no kıckıng representative of the system at large and show how the parameters affect the structure and, subsequently, the dynamics of the kicked oscillator

Figure 21 shows the undriven oscillator phase space with the circular orbits characteristic of the energy levels for the solated system The energy levels of the undriven system are given by (Messiah, 1976)

$$
\begin{equation*}
E_{n}=(n+1 / 2) h \omega_{o} \equiv r_{n} \tag{221}
\end{equation*}
$$

As the $x$ and $y$ are essentially the position and momentum respectively, then the energy of the system at any given point $(x, y)$ in the phase space is directly proportional to the sum of the squares of $x$ and $y$

$$
\begin{equation*}
E \propto x^{2}+y^{2} \tag{array}
\end{equation*}
$$

For the undriven system the energy of any orbit, $n$, is constant $\left(=r_{n}\right)$, from eq (2 21 ), so


Fıg 22 Phase space plot for $\beta=\frac{\pi}{2}, K=01$ and $\mu_{c l}=1$,


Fıg 23 Phase space plot for $\beta=\frac{\pi}{2}, K=01$ and $\mu_{c l}=95$,
the relation between the energy and $x \& y$ is a circle hence the circular orbits The phase space properties are examıned further in chapter 3


Fıg 24 Phase space plot for $\beta=\frac{(1+\sqrt{5}) \pi}{2}, K=01$ and $\mu_{c l}=65$ This is the off resonance case

Figs $22 \& 23$ show the phase spaces for two values of the kıcking when the system is on resonance, i e $\beta \equiv \frac{p}{q} \pi$ where $p \& q \in Z$ As evident from these two figures, the higher the kicking, the wider the diffuse layer between cells A quantitative analysis of this layer is undertaken in chapters $3 \& 4$

The final parr of diagrams illustrate the system for non-resonant kicking, $\beta$ not equal to $\frac{p}{q} \pi \forall p, q \in \mathbf{Z}$, and the system as modelled by the differential equations (2122) \& (2123) The lack of symmetry in the former is striking with the diffuse layer spreading over the whole plane as the driving is increased The latter llustrates just how well the mapping represents the system and how, using some straight-forward identities, the system

$F_{2 g} 25$ Phase space plot for the differential equations proper
can be described totally by a simple set of recursive mappings instead of a set of coupled differential equations It's these mappings, not the differential equations, that allow the following analysis to be carried out

The five phase space portraits shown in this section are only illustrative whereas, for any true characterisation to be undertaken, it is necessary to quantitatively analyse the components of the system's phase space This is the task we undertake in the following two chapters

## CHAPTER 3 <br> ANALYSIS OF THE PHASE SPACE

The phase space of the system may at first ghmpse appear to be just nace pactures and little else However buried in these nace pactures is the essence of the system's behaviour and evolution both as a function of its parameters and time Therefore the principle dzagnostic tool available is this phase space and its analysis is tantamount to understanding the system's behaviour

In this chapter we analyse the phase space of the classical mapping derived in chapter 2 and characterise it by considering its symmetry, by using the K A M theory, by looking at the effect of periodic points on invanant orbits (Cleary, 1990) and by examining the stochastic web's width and structure Through the adaptation of these well established techniques and by developing new ones it is our intention to obtain a greater understanding of the system proper using its principal diagnostic tool the phase space

## 31 The Symmetry Of The Phase Space

The symmetry of the system's phase space depends solely on the parameter $\beta$ which relates the natural frequency of the undriven oscillator, $\omega_{o}$, to the periodic time, $\tau$, of the drac delta function From the phase portraits in the last chapter it is possible to see just what an effect $\beta$ has on the phase space A striking example is the difference between the phase spaces in figs $23 \& 24$ which show the system for $\beta=\frac{\pi}{2}$ and $\frac{1+\sqrt{5}}{2}$ respectively The former clearly shows a four fold symmetry which we attribute to the 1dentity

$$
\begin{equation*}
\beta=\omega_{o} \tau=\frac{\pi}{2} \quad \Leftrightarrow \quad 4 \omega_{o}=\frac{2 \pi}{\tau}\left(\equiv \omega_{1}\right) \tag{array}
\end{equation*}
$$

with the ratio of $\omega_{o}$ to $\omega_{1}$ being 41 The latter figure (fig 24 ) shows no obvious symmetry primarily because the relationship between $\omega_{o}$ and $\omega_{1}$ is incommensurate The advantage of using a resonance driving is that the phase space is easier to analyse using periodic orbits with their associated stabilities and manfolds However it is possible to successfully apply some of the results of the resonance case to the non-resonance case as has been done for
diffusion m the stochastic layer (Chirkov (1979), Lachtenberg and Wood (1989), Schmera et al (1992)) Our task here is to use return maps as a technque for analysing the four fold symmetry in the resonance case we are considering This technqque is limited in that it is only apphcable to a subset of all resonance cases

The return map is one of the principle methods for quantitatively examining and lllustrating the symmetry of the system at a particular resonance value of $\beta$ This technique has been widely used to analyse classically chaotic low dımensional mappings (May (1976), Feigenbaum (1980)) However, this method is restricted in that only positive integer ratios of $\omega_{o}$ to $\omega_{1}$ (1 e $\beta=\frac{1}{q}$ where $q \in \mathbf{Z}^{+}$) can have comparable return maps Therefore if the ratio between $\omega_{o}$ and $\omega_{1}$ is $n$ then the $n^{\text {th }}$ return map is required which exists only if $n$ is a positive integer $\beta$, in most of the analysis to follow, is fixed at $\pi / 2$ This requires us to obtain an expression for the fourth return map

The fourth return map is that version of the classical mapping which relates $y(n+1)$ to $y(n-3)$ and simularly $x(n+1)$ to $x(n-3)$ Thus we can explicitly relate every fourth point The second and thurd return maps, denoted (2RM) and (3 R M) respectively, must be formulated as steps to getting an exphcit form for the fourth return map ( 4 R M ) To obtain these maps it is necessary to recall the defintion of the RKOM defined in eqs (2120) \& (2121) Takıng these equations for the mapping and expressing $x(n+1)$ and $y(n+1)$ in terms of $x(n-1)$ and $y(n-1)$ respectively our expression for the second return $\operatorname{map}(2 R M)$ is

$$
\begin{gather*}
x(n+1)=-x(n-1)+\mu_{c l} \sin \left(2 K y(n-1)+2 K \mu_{c l} \sin (2 K x(n-1))\right)  \tag{312a}\\
y(n+1)=-y(n-1)-\mu_{c l} \sin (2 K x(n-1)) \tag{312b}
\end{gather*}
$$

It follows that those points of the form $\frac{m \pi}{2 K}$ constitute steady state solutions of the above equations because penod two fixed points have, from their very definition, the property that $y(n-1)=y(n+1)$ and $x(n-1)=x(n+1) \quad$ By steady state we refer to those solutions of any return map such that $x_{n}=x_{n-k}$ where $k$ is the order of the return map and is an integer In what is to follow these solutions are those which set all the sine terms to zero The period one fixed points, of whinch they is but one the origin,
are also steady state solutions of the above as they are for all the return maps because $x(n+1)=x(n)=x(n-1)=\quad=x(0)$

The thurd return map (3RM) is obtanned in a manner simular for that of 2 R M where this tıme $y(n+1)$ is related to $y(n-2)$ and $x(n+1)$ to $x(n-2)$ However as thes return map is not required for the present analysis its form is not presented explicitly The fourth return map ( $4 \mathrm{R} M$ ) is given below

$$
\begin{align*}
& x(n+1)=x(n-3)-\mu_{c l} \sin \left(\left(2 K y(n-3)+2 K \mu_{c l} \sin (2 K x(n-3))\right)\right) \\
& -\mu_{c l} \sin \left(2 K y(n-3)+2 K \mu_{c l} \sin (2 K x(n-3))+2 K \mu_{c l} \sin (2 K x(n-3)-\right. \\
& \left.\left.2 K \mu_{c l} \sin \left(2 K y(n-3)+2 K \mu_{c l} \sin (2 K x(n-3))\right)\right)\right)  \tag{313a}\\
& y(n+1)=y(n-3)+\mu_{c l} \sin (2 K x(n-3)) \\
& \quad+\mu_{c l} \sin \left(2 K x(n-3)-2 K \mu_{c l} \sin \left(2 K y(n-3)+2 K \mu_{c l} \sin (2 K x(n-3))\right)\right) \tag{313b}
\end{align*}
$$

The penod four fixed points are those points satisfying the identities

$$
\begin{equation*}
x(n+1)=x(n-3) \quad y(n+1)=y(n-3) \tag{314}
\end{equation*}
$$

So the steady state solutions of the above two are the fixed points of period four, two or one Any multipher of four is a hypothetical steady state solution of the $4 \mathrm{R} M$ hence the inclusion of fixed points of period two and one For a solution of the form $x(n+1)=x(n-3)$ etc to occur all terms after the first on the right hand side of eqns (313a) \& (313b) must go to zero This means that all the sine arguments must all go to zero and so the penod four fixed points must satisfy

$$
\begin{align*}
& 2 K x(n+1)=2 K x(n-3)=2 K x=2 m \pi  \tag{array}\\
& 2 K y(n+1)=2 K y(n-3)=2 K y=2 n \pi \tag{array}
\end{align*}
$$

As $K$ has been chosen to be 01 for most of the analysis then it follows that the period four fixed points are given by

$$
\begin{equation*}
x=5 m \pi \quad y=5 n \pi \tag{317}
\end{equation*}
$$

where $n, m \in Z$ As the multiphers, $n$ and $m$, can take on any integer value then the period four fixed points exast as a unform grid over the entire $2 D$ plane of the phase space This has far reaching consequences for 1 t allows the unbounded transport of certan phase space orbits and also allows a diffuse layer to exist over the entire phase space Both of these consequences will be studied in detall in proceeding sections and chapters but suffice to say that this grid is essential for the stochastic behaviour m the system

Now that the exastence of this grid of points has been estabhshed it remans for us to consider whether these fixed points are hyperbolic or elliptic The reason for thes consideration is to determine in which regions of the phase space the unstable directions of the hyperbohc points exist as these determine the unbounded, diffuse behaviour previously mentioned Why this is so will become clear later The behaviour of orbits in the vicinity of elliptic and hyperbolic points determines how the phase space itself evolves both as a function of time and as a function of any of the three parameters, $\mu_{c l}, K$ and $\beta$ In order for us to determine whether a periodic fixed point is elliptic or hyperbolic it is necessary to find the fixed point's eigenvalues as given by the equation

$$
\begin{equation*}
|J-\lambda I|=0 \tag{318}
\end{equation*}
$$

where, in this $2 D$ mapping, $\lambda$ is a $2 D$ vector contanning the eigenvalues, $\left(\lambda_{1}, \lambda_{2}\right), J$ is the Jacobian of the $n^{\text {th }}$ return map (where $n$ is the periodicity of the fixed point) and is a $(2 \times 2)$ matrix with I being just the $(2 \times 2)$ identity matrix The eigenvalues of an elliptical point are in the form a complex conjugate pair such that the matrix $M$, contanning the real and imaginary parts of the eigenvalue solutions of eq (318) for this specific fixed point, is a rotation matrix about the point itself (Bergé et al (1987)) For the $2 D$ system $M$ is of the form

$$
M=\left(\begin{array}{cc}
\cos (\gamma) & -\sin (\gamma)  \tag{array}\\
\sin (\gamma) & \cos (\gamma)
\end{array}\right)
$$

For these ellhptic fixed points the complex eigenvlaues can be expressed in terms of two real numbers $\alpha \& \beta$ such that,

$$
\begin{gather*}
\lambda_{1}=\alpha+\iota \beta  \tag{3110a}\\
\lambda_{2}=\bar{\lambda}_{1}=\alpha-\iota \beta \tag{3110b}
\end{gather*}
$$

$$
\begin{equation*}
\left\|\lambda_{1}\right\|=\left\|\lambda_{2}\right\|=\lambda_{1} \lambda_{2}=\alpha^{2}+\beta^{2}=1 \tag{3110c}
\end{equation*}
$$

Furthermore the angle $\gamma$ is $\cos ^{-1}(\alpha)$, or $\sin ^{-1}(\beta)$ Thus any point $(x, y)$ in the vicinty of these elliptic points would be rotated through an angle $\boldsymbol{\gamma}$ at every iteration thus forming circular orbits From table 31 the dependence of the angle $\gamma$ on $\mu_{c l}$ is shown clearly illustrating that the eigenvalues he on a unit complex circle about which they rotate anti-clockwise with increasing $\mu_{c l}$

| $\mu_{c l}$ | $\lambda_{1}$ <br> $(\alpha+\iota \beta)$ | $\lambda_{2}$ <br> $(\alpha-\iota \beta)$ | $\gamma$ <br> $\cos ^{-1}(\alpha){\text { or } \sin ^{-1}(\beta)}^{\prime}$ | Type |
| :---: | :---: | :---: | :---: | :---: |
| 10 | $09208+\iota 03900$ | $09208-\iota 03900$ | $23^{\circ} 44^{\prime}$ | Elliptıc |
| 20 | $06928+\iota 07211$ | $06928-\iota 07211$ | $46^{\circ} 89^{\prime}$ | $"$ |
| 30 | $03448+\iota 09387$ | $03448-\iota 09387$ | $69^{\circ} 498^{\prime}$ | $"$ |
| 40 | $-00752+\iota 09972$ | $-00752-\iota 09972$ | $94^{\circ} 188^{\prime}$ | $"$ |
| 50 | $-05000+\iota 08660$ | $-05000-\iota 08660$ | $120^{\circ}$ | $"$ |
| 60 | $-08432+\iota 05376$ | $-08432-\iota 05376$ | $147^{\circ} 288^{\prime}$ | $"$ |
| 70 | $-09992+\iota 00399$ | $-09992-\iota 00399$ | $177^{\circ} 425^{\prime}$ | $"$ |
| 80 | $-08432-\iota 05376$ | $-08432+\iota 05376$ | $212^{\circ} 312^{\prime}$ | $"$ |
| 90 | $-02312-\iota 09729$ | $-02312+\iota 09729$ | $256^{\circ} 348^{\prime}$ | $"$ |
| 10 | 10000 | 10000 | $360^{\circ}$ | Hyperbolcc |

Table 91 The dependence of $\lambda_{1}, \lambda_{2}$ and $\gamma$ on the kuck strength $\mu_{c l}$ The larger $\mu_{c l}$ the faster an orbit can complete a full circuit around the fixed point once $\mu_{c l}<100$ At this value of $\mu_{c l}$ the point becomes hyperbolic and the eıgenvalues satısfy eq (3111) rather than eqns ( $3110 a-c$ )

For a hyperbolic point its eigenvalues are not complex but real with the ratio between them such that their product is 1

$$
\begin{equation*}
\lambda_{1}=1 / \lambda_{2} \tag{3111}
\end{equation*}
$$

In order for eq (3111) to be satisfied one elgenvalue must be greater than 1 with the other less than 1 Whichever eigenvalue is greater than 1 describes the unstable direction whereas the eigenvalue less than 1 describes the stable direction

The eigenvalue equation for the period four fixed points is desired as this penodicity is the fundamental symmetry of the system To proceed further it is necessary to substitute, from eq ( 317 ), the values of the period four fixed points in order to elmmate the sin terms in the expressions for the 4 R M This substitution makes for a much simpler set of equations than those meqs (313) and (314) The Jacobian's elements, $J_{i j}$ where $1 \leq \imath, \jmath, \leq 2$ and $\imath, \jmath \in \mathbf{Z}$, are

$$
\begin{gather*}
J_{11}=\frac{\partial x(n+1)}{\partial x(n-3)}=1-4 K^{2} \mu_{c l}^{2}(-1)^{n+m}  \tag{3111a}\\
J_{12}=\frac{\partial x(n+1)}{\partial y(n-3)}=4 K \mu_{c l}(-1)^{n}-8 K^{3} \mu_{c l}^{3}(-1)^{2 n+m}  \tag{3112b}\\
J_{21}=\frac{\partial y(n+1)}{\partial x(n-3)}=-4 K \mu_{c l}(-1)^{m}+{ }^{2 m+n} 8 K^{3} \mu_{c l}^{3}(-1)  \tag{3112c}\\
J_{22}=\frac{\partial y(n+1)}{\partial y(n-3)}=1-12 K^{2} \mu_{c l}^{2}(-1)^{m+n}+16 K^{4} \mu_{c l}^{4}(-1)^{2(n+m)} \tag{3112d}
\end{gather*}
$$

The eigenvalue equation given previously, in eq (318), boils down, now, to

$$
\begin{align*}
& \left(1-\lambda-4 K^{2} \mu_{c l}^{2}(-1)^{m+n}\right) *\left(1-\lambda-12 K^{2} \mu_{c l}^{2}(-1)^{m+n}+16 K^{4} \mu_{c l}^{4}(-1)^{2(n+m)}\right) \\
& -\left(4 K \mu_{c l}(-1)^{n}-8 K^{3} \mu_{c l}^{3}(-1)^{2 n+m}\right) *\left(-4 K \mu_{c l}(-1)^{m}+8 K^{3} \mu_{c l}^{3}(-1)^{2 m+n}\right)=0 \tag{array}
\end{align*}
$$

This equation is solved on a computer using the formula for obtaining the roots of a quadratic equation to give the eigenvalues $\lambda_{1} \& \lambda_{2}$ for any period four fixed point ( $5 m \pi$, $5 n \pi$ ) at specified values of $\mu_{c l} \& K$ Remember that $\beta$ is fixed now at $\frac{\pi}{2}$ It has been seen from results obtaned numerically that for $\mu_{c l}>0$ the period four fixed points with $n+m$ odd are hyperbolic with unstable and stable directions given by $\lambda_{+}$and $\lambda_{-}$respectively Furthermore for two adjacent penod four hyperbolic points $A \& B$, with $A=(5 m \pi, 5 n \pi)$ and $B=(5(m+1) \pi, 5(n+1) \pi)$, the terms $J_{21} \& J_{22}$ for $A$ are equal to $-J_{21} \&-J_{22}$


Fig 31 The principle directions of travel along the stochastic layer for $\mu_{c l}=$ $65, K=01$ and $\beta=\pi / 2$ The arrows indicate the drections transport takes place in Arrows pointing away from a crossover (hyperbohc fixed point) inducate unstable directions whule stable point towards the fixed point
repsectively for $B$ The result of this is to cause the unstable direction of $A$, given by $\lambda_{+}^{A}$, to coincide with the stable direction of $B$, given by $\lambda_{-}^{B}$, and vice versa A plot showing this is given in fig 31 From this one is able to deduce that orbits in the vicinty of the unstable manfold of any hyperbohc period four fixed point, denoted P1, can be shoved away from it to one of two of its neighbours, $P 2 \& P 2^{\prime}$, because their stable directions are coincident with the unstable directions of $P 1$ Once in the vicnity of $P 2$ or $P 2^{\prime}$ the orbit is then pushed away along their unstable manufolds to one of two of their neighbours and so the process can continue until the orbit has visited all the penod four hyperbohc fixed points in the $2 D$ plane of the phase space (Lowenstein, (1991), Radons \& Prange (1990))


Fig 32 Thus phase plot shows the dumb bell type separatrix about the newly formed hyperbolic points in the cells' centers for $\mu_{c l}=110, K=01$ and $\beta=\pi / 2$ The cell eventually breaks into two preces whose centers correspond to the centers of the loops These smaller cells eventually break up in a simular way

As the period four fixed points with $n+m$ odd are hyperbolic then naturally it is safe to presume those with $n+m$ even are elliptic This is $m$ fact the case up to the value of $\mu_{c l}$ equal to 10 At this value and for values above these elliptic points become hyperbohc with their manfolds taking on a dumb bell shape as evident from fig 32 The implication of this dumb bell shape is that the unstable manfolds of these points coincide with the stable manfolds at the extreme of each loop with the result that orbits near the unstable manfold get repelled only to be attracted along the stable manufold further on This causes the orbits to remain localised within the cell. Eventually, as the kicking strength is increased, the outer orbits of the cell break up and the localised orbits become de-localised
and diffuse Furthermore, as will be shown later in the chapters relating to the quantum mapping, this dumbbell type structure and the subsequent break-up of the period four invariant cells into two smaller cells is very fortunate for us This property allows us to analyse the quantum behaviour at the origin (using low order eigenfunctions) knowing that it will become unstable and that this stabilty can be measured by examining the evolution of energy and the probability amphtudes This is done in detal in chapter 6, sections 61 and 62

Also apparent are other separatrices bordering the diffuse region These separatrices are inutially sharply defined but as the kuck strength is increased they become diffuse and eventually indistinguishable from the surrounding layer At this point all that remains are the small island cells in the layer These cells contan an elliptic point at their center with invariant orbits surrounding them As the kick strength is further increased the measure of the cells reduces due to orbit breakup at the cell boundary with the stochastic layer As the kıck strength increases further separatrices form in some of the orbits with small cells surrounding higher order ellhptic fixed points As before these separatrices become diffuse and the small cells break away from the larger mother cell which is getting smaller Eventually the elliptic points in the center of these cells become hyperbolic and all invariant cells tend to measure zero with the stochastic layer occupying the whole of the $2 D$ phase space except for the locations of the periodic points which have measure zero (Helleman (1980)) These break-ups can be best described by the following diagram

where $\mu_{c l}$ denotes increasing $\mu_{c l}$ Thus form of break up is predicted by the Kolmogorov Arnold Moser (K A M ) theorem which is used in the following section to analyse the phase space in a more quantitative manner

## 32 Of Action-Angles \& K A M

Recall how, in chapter 2, we showed how circular orbits should arise naturally in the phase space of the undriven harmonc oscillator As these carcular orbits hint at an underlying circular symmetry it is prudent for us to consider expressing the a \& a operator vaniables in terms of a set of circularly symmetric operators variables These vanables are referred to as Action-Angle $(A-A)$ vanables denoted $J$ (action) and $\theta$ (angle) Once $J$ and $\theta$ are defined with respect to $a$ and $a^{\dagger}$ they can be substituted into the onginal system Hamltoman, $H_{T}$ The advantage of thus change of vanables is to readily allow us check the influence of the non-integrability of the kick term on the oniginal integrable undriven system Our defintion of integrablity is that if the hamiltoman can be expressed as a function of $J$ only then it is integrable We proceed by defining $J$ and $\theta$ according to

$$
\begin{equation*}
a=\sqrt{J} * \exp (\iota \theta) \quad \text { and } \quad a^{\dagger}=\sqrt{J} * \exp (-\iota \theta) \tag{array}
\end{equation*}
$$

The action-angle variables can be expressed directly in terms of the $P$ and $Q$ operators with the purpose for this being our ablity to directly compare the classical $x$ and $y$ to the classical equivalent of the action-angle variables

$$
\begin{equation*}
Q=\sqrt{J} * \cos (\theta) \quad \text { and } \quad P=\sqrt{J} * \sin (\theta) \tag{3}
\end{equation*}
$$

When the Hamiltonian for the undriven oscillator, $H_{o}$, given in eq (215a) is expressed m this new set of variables it becomes independent of $\theta$ and is a function of $J$ only

$$
\begin{equation*}
H_{o}(J, \theta) \equiv H_{o}(J)=J \hbar \omega_{o} \tag{3}
\end{equation*}
$$

such that the energy of the system is related directly to the variable $J$ not $\theta$ This is essentially a canonical form of the original $H_{o}$ The variables $J$ and $\theta$ are also dimensionless as is apparent from eqs ( $\mathbf{3} 21$ ) and (3 23 ) When this Hamiltoman is transposed to the classical representation we define a new set of variables such that the following identities holds

$$
\begin{equation*}
x=\rho * \cos (\phi) \quad \text { and } \quad y=\rho * \sin (\phi) \tag{34}
\end{equation*}
$$

where $\rho$ and $\phi$ are our ciassical equvalents of $J$ and $\theta$ so that our transposition takes $\sqrt{J / 2}$ to $\rho$ and $\theta$ to $\phi$ We can now directly identify the circular orbits in the classical space to
these new variables As has been previously pointed out, the phase space of $H_{0}$ consists of concentric circular orbits which, in the $\rho, \phi$ representation, are given by

$$
\begin{equation*}
\rho^{2} \sin ^{2}(\theta)+\rho^{2} \cos ^{2}(\theta)=\rho^{2} \tag{325}
\end{equation*}
$$

where $\rho$ is essentially the radius of the orbits The larger the value of $\rho$ the larger the circle and the hagher the energy of the orbit

Such a system as that represented by the Hamultoman in eq (3 2 3) is considered integrable because the Hamiltomian can be written as a function of $J$ solely and has no direct dependence on $\theta$ A non-integrable system can be considered to be one where the hamultoman cannot be so written

Our interest here is in the kicked system so let us now proceed by expressing the complete system Hamiltonian, $H_{T}$, defined m eq (216) m terms of $J$ and $\theta$

$$
\begin{equation*}
H_{T}(J, \theta)=H_{o}(J)+\mu_{q} H_{1}(J, \theta) \tag{326}
\end{equation*}
$$

This shows $H_{T}$ to consist of an integrable term, $H_{0}$, and a non-integrable term $H_{1}$ The integrable term is that for the undriven system as described above whereas the non-integrable term is that for the driving term We have for clarity removed the kick strength variable, $\mu_{q}$, from the hamiltoman $H_{1}$ so that the proceeding analysis will be easier to understand The classical equivalent of eq (326) would be the corresponding energy equation which depends on $\rho$ and $\phi$

$$
\begin{equation*}
E_{T}(\rho, \phi)=E_{o}(\rho)+\mu_{c l} E_{1}(\rho, \phi) \tag{327}
\end{equation*}
$$

where $E_{T}, E_{0}$ and $E_{1}$ are the corresponding classical terms to $H_{T}, H_{0}$ and $H_{1}$ respectively The remander of this section is devoted to explanning what effect $\mu_{q}$ has on the integrability of $H_{T}$ and by consequence $\mu_{c l}$ on $E_{T}$ The KA M theorem is also invoked to help explan why some orbits in the classical phase space break up quicker than others

For large values of kıckıng potential, $\mu_{c l} E_{1}$, the system is certannly non- mtegrable due to a domination by this term over the $E_{o}$ term The fundamental question, however, 15 To what level is the system non-integrable for small non- zero values of the kacking potental and how does thes level of non-integrability affect the mnvariant orbats in the phase space ${ }^{?}$

This is the question that the K A M theorem addresses According to this theorem If the non-integrable term is sufficiently small to behave more as a perturbation on the mtegrable term than as a major contributory term then the invariant orbits remain if the orbits have sufficiently incommensurate frequencies $\omega_{1}$ Those with commensurate frequencies or frequencies neariy commensurate do not remain and are distorted to a lesser or greater extent with some even being destroyed

For small $\mu_{c l}$ we can see from the phase space portrait in fig $33(a)$ that the majority of the orbits are circular in shape and are thus largely unaffected by the kicking term Subsequent enlargements (figs $33(b),(c) \&(d)$ ) about the hyperbolic fixed point at $(0,5 \pi)$ show that the stochastic layer is very small indeed being confined very tightly to the manufold structure resulting from the hyperbolic fixed points of pernod four (cf fig 3 1) The only orbits with any noticeable distortion are those whose non-integrable contribution 15 comparable, or greater, to their mtegrable term As $\mu_{c l}$ is small then those orbits with large $E_{1}(\rho, \phi)$ can have non-mtegrable contributions which can not be considered as vanushingly small Indeed it is only these orbits which have some level of discernible distortion Therefore we can conclude that for small $\mu_{c l}$ the level of non-integrability in the system as a whole is so small as to be neghgible

Having addressed the problem for small $\mu_{c l}$ we now concentrate on the problem for $\mu_{c l}$ finite and beyond the perturbation realm We would expect that if $E_{1}(\rho, \phi)$ was of a sufficient size to be capable of distorting phase space orbits when $\mu_{c l}$ was small, and merely perturbing the system, then as $\mu_{c l}$ is increased the distortion would increase and orbital (toral in higher dimensions) breakup would arise probably through an overlapping of higher resonance separatrices This will be examined in detail in the next section

It is worth noting that as $E_{1}(\rho, \phi)$ is not constant then it can negate, or even reverse, the effect of an increasing $\mu_{c l}$ if it is sufficiently small In this case the non-integrable term becomes more of a perturbation and the orbits remain essentally undistorted as we can see around the elliptic point at the ongin In conclusion, the higher $E_{1}(\rho, \phi)$ for any $\mu_{c l}$, the greater the distortion and the sooner the break up of the orbits For our system we will show that the magnitude of $E_{1}(\rho, \phi)$ determnes how many terms in the senes expansion of its cosine term are needed to accurately describe the behaviour. More terms

$F_{\imath g} 93$ The phase space of the kacked oscllator for $\beta=\frac{\pi}{2}, K=01 \& \mu_{c l}=05$ The enlargements in parts (b), (c) \& (d) about the hyperbolic fixed point at $(0,5 \pi)$ show how confined the stochastic layer is but that it does exast nevertheless
allow for higher order periodic points and hence an earher break up of the orbits than for orbits with less terms and hence a smaller $E_{1}(\rho, \phi)$ In the above analysis the classical terms, $E_{1}, \mu_{c l}$ and $E_{0}$, were used However their effects on the system can be related du-
rectly back to therr quantum counterparts and the arguments presented above could quite easily have been presented in terms of $H_{1}, \mu_{q}$ and $H_{0}$ This inter-relating of quantum and classical is important to the analysis using the quantum mapping in chapters $6 \& 7$

## 33 Perıodıc Orbital Analysıs

It is our intention here to quantitatively analyse what was stated in the previous section To accomplish this we intend to construct a form of the non-mtegrable term which will allow us to predict which orbits should be present given the type of potential we are using to drive the system

Let us start by examining the driving term $H(J, \theta)$ expressed as a function of the Action-Angle varıables $J \& \theta$

$$
\begin{equation*}
H_{1}(J, \theta) \equiv \cos \left(\sqrt{\frac{J \hbar k^{2}}{M \omega_{0}}} \sin (\theta)\right) \sum_{n=0}^{\infty} \delta(t-n \tau) \tag{array}
\end{equation*}
$$

We have deliberately substituted $K$ for its constituents as defined in chapter 2 as this will permit us to explicitly equate the $\omega_{o}$ term to the driving frequency $\omega_{1}(\equiv 2 \pi / \tau)$ The cosine term is now replaced by its series expansion to give

$$
\begin{equation*}
H_{1}(J, \theta) \equiv \sum_{m=0}^{\infty}\left(\frac{-h k^{2}}{M \omega_{1}} J\right)^{m} \frac{\sin ^{2 m}(\theta)}{(2 m)^{1}} \sum_{n=0}^{\infty} \frac{\delta(t-n \tau)}{r^{m}} \tag{array}
\end{equation*}
$$

where $\omega_{o}$ has been replaced by $r \omega_{1}$ and $r$ is the ratio between $\omega_{o}$ and $\omega_{1}$ The above can now be expressed in the more readable form

$$
\begin{equation*}
H_{1}(J, \theta)=\sum_{m=0}^{\infty}(-1)^{m} f_{m}\left(J, \theta, \omega_{1}\right) \sum_{n=-\infty}^{\infty} g_{m n}\left(\omega_{1}, r\right) \tag{array}
\end{equation*}
$$

The classical equivalent for the above is identical m form but has of course the classical equivalents of the quantum operators and variables The classical form is the one we shall use in the following analysis as this analysis requires comparisons between the predicting equation and results obtanned from numerical simualtions in the classical phase space Our classical form 1 s

$$
\begin{equation*}
E_{1}(\rho, \phi)=\sum_{m=0}^{\infty}(-1)^{m} \bar{f}_{m}\left(\rho, \phi, \omega_{1}\right) \sum_{n=-\infty}^{\infty} \bar{g}_{m n}\left(\omega_{1}, r\right) \tag{33}
\end{equation*}
$$

To analyse the behaviour of this system we need to consider the relationship between $\omega_{o}$ and $\omega_{1}$ Remember that $\beta=\frac{\pi}{2}$, that is $r=1 / 4$ Therefore we have, in the function $g_{m n}\left(\omega_{1}, r\right)$, a way of describing the affect the relationship between $\omega_{o}$ and $\omega_{1}$ has on the exastence of resonances in the complete system it is evident from eq (331) that not only is the cosine term sampled by the delta function at a frequency $\omega_{1}$ but also that this term has an infinte set of intrinsic frequencies, $\omega^{m}\left(\equiv r^{m} \omega^{m}\right)$, and that the ratio between these frequencies and $\omega_{1}$ allows the various resonances to exast We can use these resonances to predict the perioduc points m the system's phase space Thus, depending on the values chosen for $m$, the resonances allowed can be vanous and not just multiples of four it is these resonances that cause the breakup of orbits in the phase space because only those orbits which are sufficiently incommensurate will survive an increase in the non-integrability of the system (as predicted by K A M theory)

How sufficient is sufficient ${ }^{\ell}$ This condition of sufficiency may become more apparent by examinng the resonances at some low values of $m$ Table 32 shows how for a given value of $m$, the number of expected resonaces is dependent on the number of factors of $r^{m} \quad$ Furthermore, the statement 'maybe all factors which almost divide evenly into 256' imples that while there are multiples of 4 which do divide into its higher powers evenly there can nevertheless exist periodic orbits which have a periodicity close enough to a divisor of 256 , or any power of $4^{m} \forall m \in \mathbf{Z}$, with a small remander These constitute the quasi-commensurate orbits which accelerate the destruction of orbits with small $E_{1}$ whose resonances otherwise would be of insufficient number to break up the orbits at the same value of $\mu_{c l}$ The periodic points actually seen in the system to date have been $1,3,4,5,6,7,8,12,16,21,24,32,48,72$ The inherent difficulty of locating, numencally, fixed points of a specific periodicity is well known especially if the points have small lyapunov exponents as this results in slow convergence to the points themselves (Auerbach et al (1987)) Therefore the list presented above is as complete as permitted by the limit on computing time and by the resolution of the search gnid used to locate the points

The non-integrable cosine term, $E_{1}$, in the system equation can only be described by its first few senes terms when its argument $(\alpha \rho \sin (\phi))$ is small As the $\rho$ can be considered, in the undriven osciliator, to be analogous to the radi of orbits then in the

| $m$ | resonances | $r^{m} \equiv \omega_{1} / \omega^{m}$ |
| :---: | :---: | :---: |
| 0 | 1-1 | $\omega_{1}$ |
| 1 | 1-4\&1-1 | 4 |
| 2 | $\begin{gathered} 1-16,1-8,1-4,1-1 \\ \text { maybe a } 1-5 \& 1-9 \end{gathered}$ | 16 |
| 3 | $\begin{gathered} 1-64,1-32,1-16,1-8,1-4,1-1 \\ \text { maybe } 1-21,1-13,1-12,1-9,1-7,1-5,1-9 \end{gathered}$ | 64 |
| 4 | all multiple of 4 up to 256 maybe all factors which almost duvide anto 256 with small remainders | 256 |

Table 32 The first few terms in the expansion of the driving term in eq ( $\begin{aligned} & 3 \\ & 3\end{aligned} \mathbf{3}$ ) showing what resonances would be expected to be present Those resonances shown in italics are those which nearly divide into the number given $r^{m}$
driven case, for those orbits near the origin, we can make a simular identification As the phase space has a definte four-fold symmetry then all invariant cells centred about a period four elliptic point satisfying the condition in Eq (317) can be transformed to the origin Therefore we can presume that the contribution from the non-integrable term $E_{1}$ about each of these elliptic fixed points is small So $\rho \sin (\phi)$ tends to zero in the vicinity of the elliptic points of period four The number of possible resonances in this region is small because of the small number of terms, in the cosine expansion of $E_{1}$, needed to adequately describe the system's behaviour (see table 32 ) Thus the possibllity of orbital breakup is small for small values of the parameter $\mu_{c l}$, but becomes increasingly more probable as the kıck strength $\mu_{c l}$ is increased The system is essentially integrable in these regions for small $\mu_{\text {cl }}$

However the further out we go from the elliptic fixed points at each invariant cell's centre then the more terms we requre to satisfactorily describe the system's behaviour and
hence the greater the number of possible resonances. There is a constant multiplier ( K ) in the $E_{1}$ term which takes on the value 0.1 in the case being considered here. As a consequence of this the $x^{\prime} s$ and $y^{\prime} s$ have prominent fixed points every $5 \pi$ (as in Eq. (3.1.7)) with hyperbolic fixed points of period four at specific multiples of $5 \pi$ (as previously explained in section 1 of this chapter). The regions of the orbits nearest these hyperbolic points have largest non-integrability and hence require the most terms in the cosine expansion of $E_{1}$. The most terms implies the greatest resonance overlap and for large $\mu_{c l}$ their contributions can be very significant.


Fig.3.4 The region of phase space surrounding the hyperbolic fixed point at $(0,5 \pi)$ for $\beta=\frac{\pi}{2}, K=0.1 \& \mu_{c l}=6.5$. Fixed points (both hyperbolic and elliptic) of various periodicities are shown: period 4 are denoted by squares, period 16 by diamonds \& period 24 by triangles

So the break up of the orbits occur in the region most non-integrable (that is regions where $K \rho \sin (\phi)$ is not small) For small $\mu_{c l}$ the driving term (regardless of how many terms are in the cosine factor) is small and the system follows the scenario of commensurate orbital breakdown as described by the KAM theorem As $\mu_{c l}$ is increased the regions following this scenario contract about the elliptic fixed points described earher The outer regions for these large $\mu_{c l}$ have many overlapping separatrices from each of the boundanes of the resonances These overlapping separatrices break up and form stochastic regions around the unbroken, albett distorted, orbits The width of the stochastic region increases with increasing $\mu_{c l}$ due to the contraction of the K A M regions about the elliptic fixed points These are evident from the phase portraits shown mf fig 3

We should expect even for very small $\mu_{c l}$ the exastence of a stochastic region around the hyperbohe points as these regions are the most non-integrable for any value of $\mu_{c l}$ It may occur that a stochastic region doesn't exast below a particular value of $\mu_{c l}$ due to the small nature of the the kicking However as this region contains the greatest number of possible resonances (due to the large $\rho$ value) it follows that separatrix break-up into stochastic layers is very probable at small $\mu_{c l}$ values The width of such a layer would be extremely thin and consequently would be very hard to locate numerically as the boundary would be quite sharp requiring high numerical precision

Fig 34 shows an enlargement of the phase space about the hyperbohc fixed point at $[0,5 \pi]$ for $K=01, \mu_{c l}=65$ and $\beta=\pi / 2$ The stochastic layer is clear in the figure as are some small invariant cells within the mother cells of the period four elliptic fixed points Supenmposed on this phase space are various opaque geometrical objects at the positions of the period 4 hyperbolic fixed point, the period 16 fixed points and the period 24 fixed points The points on the layer boundary are the penod 24 pomts whose separatrix surrounding the invariant cells are now diffuse leaving an island chain of cells with period 24 elliptic points at therr centers

## 34 The Stochastıc Layer

One of the important features of the phase space in this classical mapping is the
existence of a layer between elliptical cells which allows the unbounded growth of energy in the system In this stochastic layer (so named because of its diffusive nature) the successive iterates of the mapping can wander around the whole of phase space with the effect that the energy of the diffuse orbit can increase without any limit This aspect of the layer is examined in detanl in chapter four What we are going to analyse here is the width of the stochastic layer and also to compare the largest Lyapunov exponent to the stabilty of nearby periodic points

So let us proceed with a detaled examination of the variation of the stochastic layer width as a function of the kick parameter, $\mu_{c l}$ Generally it is accepted that for small values of the driving parameter the width increases in an exponential manner (Zaslavsky and Filonenko (1968), Zaslavsky et al (1986) and Chermikov et al (1989)) To illustrate this we present four phase plots of the layer for different values of $\mu_{c l}(10,30,50,70)$ showing both the increasing size and complexty of the layer boundary This increase in boundary complexaty caused problems with the program used to measure the width as will be explanned later

To measure the width, a program was written which sampled points on the main $x=0$ axis through the hyperbolic fixed point $(0,5 \pi)$ about which the layer exists The initial guess was to see if before $N$ iterations of the mapping the initial condition caused the $n^{\text {th }}$ 1terate (where $n \leq N \forall n \in \mathbf{Z}$ ) to leave the region of the bounded orbits and wander about the phase space If so then the distance from this point to the fixed point at $(0,5 \pi)$ was measured and this was the measure of the layer width The previous four plots show how, at the boundary layer, periodic points of increasing order, with their elliptic islands about elliptıc points and separatrices through hyperbohc points, abound at high values of $\mu_{c l}$ but are noticeably absent for low values Obviously the driving term $H_{1}(J, \theta)$ has a lesser degree of influence at low values whereas at higher values the resonance overlaps and potential for orbital break-up by said resonances bring about this prevelence of periodic islands in the layer These islands greatly affect the layer width measurement and give rise to noticeable dips in the layer width curve

The width variation, for $\mu_{c l}$ increasing from 10 to 30 inclusively, is shown in fig 36 The general trend seems to indicate an exponential increase with the data from the


Frg 35 Four plots of the stochastic layer about a hyperbolac fixed point for $\beta=$ $\frac{\pi}{2}, K=01 \&$ four different kack strengths In (a) $\mu_{c l}=05$, in (b) $\mu_{c l}=15, \ln (\mathrm{c}) \mu_{c l}=45 \& \ln (\mathrm{~d}) \mu_{c l}=50$
simulation fitting an exponential curve of the form

$$
\begin{equation*}
\Delta W=\frac{8 \pi^{3}}{\mu_{c l}} \exp \left(-\sqrt{5} \frac{\pi^{2}}{\mu_{c l}}\right) \tag{array}
\end{equation*}
$$

This form agrees exactly with that analytically predicted (for kucking strengths $\ll 1$ ) by


Fig.9.6 The stochastic layer width vs. the kick strength, $\mu_{c l}$. The line corresponds to the theoretical curve while the points correspond to the measured width. The agreement is very good considering the time it takes for each measurement to converge. For values higher then $\mu_{c l}=3.0$ we measured dips in the measured width attributed to the growth of islands on the stochastic layer boundary.

Chernikov et al. ((1989),(1987)). (Note that the $\sqrt{5}$ in the argument of the exponential arises from our explicit inclusion of $K$ in our initial definition of $\mu_{c l}$. See chapter two for our justification of this.)

For values of $\mu_{c l}$ greater than 3.4, dips occur in the curve probably due to the presence of the periodic islands growing about high order periodic points on the boundary where orbital break up occurs due to resonance overlap. These islands are evident only
at values of $\mu_{c l}$ greater than 30 and are considered by me to be the cause of the dips How a dip occurs could be as follows as the hne along which the minial points for each set of iterations passes from the bounded orbital region to the stochastic region, it may encounter separatrices and periodic islands in its way The separatrices may initially be bounded and hence non-stochastic but at some value $\mu_{c l}$ become stochastic giving rise to a jump in the width curve However the islands may also increase in size pinching these regions and causing a dip with respect to the previous value if the program takes mintial points that skup this region The only way to measure the layer for the higher values is to plot the phase space, enlarge the boundary region and manually measure the region's width

The chaos in such a conservative system as this (conservative because the Jacobian is 1 always) is attributed to this layer's exastence This is true because in this layer nearby points separate exponentially and have a positive Lyapunov exponent In fig 37 we show this sensivity to mitial conditions by measuring the separation between two points initially close to one another The mitial points taken were $\left(x_{1}, y_{1}\right)=(10,15707963)$ and $\left(x_{2}, y_{2}\right)=(1000001,15707964)$ ie an addition of 0000001 to both $x_{1}$ and $y_{1}$ As can be seen from the plot in fig 37 the separation varies widely and this with the previously mentioned positive Lyapunov exponent confirms the presence of chaos (in the classical sense) in this layer

Furthermore we can trace the exponential divergence of nearby orbits back to the manifolds for the period four fixed points (see fig 31 and the discussion following it) along which orbits can diffuse over the whole phase space Furthermore depending on how near to, or on which side of, an unstable manfold an orbit is, determines on which unstable manfold it moves along at the next period four hyperbolic point Thus nearby orbits can separate exponentially along this manufold structure We have found that the calculated positive Lyapunov exponent (for a specific orbit in the stochastic layer) vs $\mu_{c l}$ follows a weighted average of the eigenvalues of the hyperbolic fixed points the orbits visits The weighting depends on which point is visited and on which unstable mamfold it leaves For large $\mu_{c l}$ the complexaty of the layer and the proliferation of higher order hyperbolic fixed points (brought about by invariant orbit break up by overlapping resonances) brings about


Fig 97 The separation of two nearby trajectories in the stochastic layer Intially very close together, they diverge very rapidly showing clearly the sensitive dependence on intital conditions, the hallmark of classical chaos
a varnance in the above relationship between the Lyapunov exponent and the eigenvalues of the hyperbolic points This vanance is put down to the increased number of unstable manfolds from these higher order hyperbolc fixed points bringing about a more complex motion in the layer than before To our knowledge most of the analytical approaches to date have not taken into account these compexaties

## CHAPTER 4 DIFFUSION IN THE STOCHASTIC LAYER

The stochastic layer has been shown in the last chapter to hold the key to the non-linear characteristics of the system Therefore this layer is the focus of much of the preceding and proceeding analysis The unbounded nature of phase space orbits within the layer proper points to the ability of unbounded energy growth in these orbits as shown in the last chapter Furthermore as these orbits are unbounded the method of deffusion of collections of orbits is important as in any physical system it is more usual to expenence the ensemble properties rather than an individual In this chapter the diffusion of a large number of orbits is examined and some analytical arguments presented whose results fit the numencally obtaned data remarkably well

## 41 Green's Function \& Diffusion

The following approach is based on the analysis done by Rechester and White (1980) which led to an analytical form for the diffusion in the stochastic layer of the Churikov-Taylor system (Chirikov et al, 1979) Their analysis was based on the use of a Green's function solution to the Vlasov equation for turbulent system with an added diffusion term The approach taken here is simular in form We will show how the extra potential term for the kucked oscillator doesn't affect the analysis for the strength of kacking we are considering To justify any approximations we are going to make it is necessary to consider the contribution of the stable orbits to the system at large over a range of kick strengths, $\mu_{c l}$

We have already shown by lliustration how the stochastic layer increases exponentrally with increasing kıck strength, $\mu_{c l}$, and how the invariant cells break up by resonant overlap according to the K A M theorem (see chapter 3) We would therefore expect the area occupied in the phase space by the invariant (stable) cells to diminish as $\mu_{c l}$ is increased tending to zero as $\mu_{c l}$ tends to infinity In fact as this limit is approached the cells contract about any remaining elliptıc fixed points untıl they become measure zero As will be shown later, this lumit also tends to obliterate any correlations between any iterates of
the mapping (Churikov, (1987)) and a random phase approxamation can be used for the system

We do not have to go to such extremes to neglect the contributions of the stable orbits, however We can define a critical kick strength, $\mu^{C}$, above which the stable orbits' contributions are ignored as the area of phase space occupied by them has diminushed below a set level (say 10\%) (This choice is arbitrary once the dominant term is the unstable kıck term ) The periodicity of the phase space ( $\pi / K$ ) imphes that the dynamical information contained in any region of side $\pi / K$ centred on an invariant cell of order four (or one for the ongin) is indicative of the system at large Therefore we may restrict $q$ and $p$ to $\pm \pi /(2 K)$ Then the normalised momentum, $P$, evolving according to eq (2122) can be modfied to

$$
\begin{equation*}
\frac{\partial P}{\partial t}=\omega^{2} Q+\mu_{c l} \sin (2 K Q) \sum_{n=1}^{\infty} \delta(t-n \tau) \cong \mu_{c l} \operatorname{sm}(2 K Q) \sum_{n=1}^{\infty} \delta(t-n \tau) \tag{array}
\end{equation*}
$$

which is valid only when $\mu_{c l}>\mu^{C}$ Thus can then be simplefied, by normalising the time between kucks to 1 , to the following

$$
\begin{equation*}
\Delta P \cong \mu_{c l} \sin (2 K Q) \tag{array}
\end{equation*}
$$

which is a vital identity for the calculation of the diffusion coefficient $D_{c l}$ which is defined according to (Chinkov (1987))

$$
\begin{equation*}
D_{c l} \propto \lim _{t \rightarrow \infty} \frac{\left\langle(\Delta P)_{t}^{2}\right\rangle}{2 t} \tag{array}
\end{equation*}
$$

where $\langle>$ is an ensemble average and $t$ is the time (for the mapping $t \equiv n$ the iteration counter) So what value of $\mu_{c l}$ 2s the critical value for the above to be valdd ${ }^{9}$ To see this we include a table of the area occupied by both the invariant cells and the stochastic layer as a function of $\mu_{c l}$ (Refer back to chapter 2 when we defined $\mu_{c l}$ for eq (2116) to recall tts exphcit dependence on $K$ )

From table 41 we can guess at a value for $\mu^{C}$ which for our constraint (occupation $<10 \%$ ) would place it close to 14 for $K=01,70$ for $K=02$ and so on We will show how this dependence on $K$ affects the diffusion curves later We will now begin the analysis proper

| $\mu_{c l}$ <br> $(K=01)$ | $\mu_{c l}$ <br> $(K=02)$ | $\mu_{c l}$ <br> $\left(K=K_{\mathrm{arb}}\right)$ | \% area occupied <br> (1nvariant cells) | \% area occupled <br> (stochastic) |
| :---: | :---: | :---: | :---: | :---: |
| 00 | 00 | 00 | 100 | 0 |
| 20 | 10 | $20 * k$ | 9996 | 004 |
| 40 | 20 | $40 * k$ | 9210 | 790 |
| 60 | 30 | $60 * k$ | 7361 | 2639 |
| 80 | 40 | $80 * k$ | 3285 | 6715 |
| 100 | 50 | $100 * k$ | 2323 | 7677 |
| 120 | 60 | $120 * k$ | 1589 | 8411 |
| 150 | - | 75 | $150 * k$ | 615 |
| 200 | 100 | $200 * k$ | $05-1$ | 9385 |

Table \& 1 Occupied area of the phase space vs the kuck strength for the invanant cells and the stochastic layer when $\beta=\pi / 2 \quad k$ above is just ( $01 / K_{\text {arb }}$ ) allowing us to express the kack strength at arbitrary $K\left(K_{\text {arb }}\right)$ in terms of $K=01$ The dashed hne represents the cut-off below which we can neglect the stable component of the motion for the acceleration ( $d P / d t$ ) equation Beyond $\mu_{c l}=200 * k$ we can consider the system's phase space to be completely stochastic

The main thrust is to obtain an expression for the probability which satisfies the Vlasov equation and then to use this to form an expression for the diffusion coefficient, $D_{c l}$ We begin by considering the Vlasov Equation with diffusion added

$$
\begin{equation*}
\frac{\partial P}{\partial t}+v \frac{\partial P}{\partial Q}+\frac{d v}{d t} \frac{\partial P}{\partial v}-\frac{\sigma}{2} \frac{\partial^{2} P}{\partial Q^{2}}=0 \tag{array}
\end{equation*}
$$

where the probability distribution function is $P(Q, v, t)$ Noting that the time is quantised because of the periodic delta function then the system evolves across a kuck as

$$
\begin{equation*}
P(Q, v, n \tau+0)=\left\{P\left(Q, v+\omega^{2} Q+\xi \mu_{c l} \sin (Q), n \tau-0\right)\right\}_{Q \equiv Q_{t=n \tau}} \tag{array}
\end{equation*}
$$

where we have re-cast the $Q$ to eliminate the $2 K$ before it in the sine term This has the effect of changing the kick strength to $\xi \mu_{c l}$ where $\xi=K / K_{0}$ with $K_{0}=05$ Thus re-casting also has the affect of changing the system's periodicity from $\pi / K$ to $2 \pi$ allowing important identities to be made later in this analysis Between kicks the system evolves by the formula

$$
\begin{equation*}
P(Q, v, t)=\int_{-\infty}^{\infty} d v_{1} \int_{0}^{2 \pi} d Q_{1} G\left(Q-Q_{1}, v, v_{1}, t-t_{1}\right) P\left(Q_{1}, v_{1}, t_{1}\right) \tag{416}
\end{equation*}
$$

where the function $G\left(Q-Q_{1}, v, v_{1}, t-t_{1}\right)$ is a Green's function and satisfies the equation

$$
\begin{equation*}
\frac{\partial G}{\partial t}+v \frac{\partial G}{\partial Q}-\frac{\sigma}{2} \frac{\partial^{2} G}{\partial Q^{2}}=\delta\left(x-x_{1}\right) \delta\left(t-t_{1}\right) \delta\left(v-v_{1}\right) \tag{array}
\end{equation*}
$$

The Green's function is used as an evolution operator between kicks and because the velocity changes only across a kıck (see eq (412) for $d P / d t$ ) the $d v / d t$ is zero between kıcks Hence the Vlasov equation in eq (414) becomes that in eq (417) between kucks The solution to the above can be found in books on partial differential equations with boundary condtions This specific solution can be found by considering the Green's function solution to the diffusion equation $u_{t}-u_{x x}=\delta(x-\xi) \delta(t-\tau)$ (Kervorkian (1990)) Our solution is of the form

$$
\begin{equation*}
G\left(Q-Q_{1}, v, v_{1}, t-t_{1}\right)=\frac{\theta\left(t-t_{1}\right) \delta\left(v-v_{1}\right)}{\sqrt{2 \pi \sigma\left(t-t_{1}\right)}} \sum_{n=-\infty}^{\infty} \exp \left[-\frac{\left(Q-Q_{1}-v\left(t-t_{1}\right)+2 n \pi\right)^{2}}{2 \sigma\left(t-t_{1}\right)}\right] \tag{418}
\end{equation*}
$$

The initial value chosen for the probabilty function $P(Q, v, t)$ is

$$
P(Q, v, 0)=\delta\left(v-v_{o}\right)
$$

The velocity integral is non-contributory and is omitted from now on The position variable $Q$ has a constant derivative between kucks as the velocity is constant and varnes
only across a kick The integral over a kick just gives us the next $Q$ position in terms of the previous one so that the probabily at a tume $T$ for a position $Q_{T}$, where $Q_{T}$ is the value of $Q$ at a tıme $T$, is the product of the probabilities at each time step so that

$$
\begin{gather*}
P(Q, v, T)=\sum_{n_{T}=-\infty}^{\infty} \sum_{n_{1}=-\infty}^{\infty} \prod_{\imath=0}^{T-1} \int_{0}^{2 \pi} d Q_{i} \frac{1}{\sqrt{2 \sigma \pi}} \delta\left(v-v_{o}-S_{T}\right) \\
\exp \left[\frac{-1}{2 \sigma} \sum_{j=1}^{T}\left(Q_{J}-Q_{J-1}-v_{o}-S_{j-1}+2 n_{\jmath} \pi\right)^{2}\right] \tag{419}
\end{gather*}
$$

where

$$
\begin{equation*}
S_{3}=\sum_{p=0}^{J} \omega^{2} Q_{p}+\xi \mu_{c l} \sin \left(Q_{p}\right) \tag{4110}
\end{equation*}
$$

The diffusion rate is calculated from the formula

$$
\begin{equation*}
D_{c l}=\lim _{T \rightarrow \infty}(2 T)^{-1} \int_{-\infty}^{\infty} \int_{0}^{2 \pi}\left(v-v_{o}\right)^{2} P(Q, v, T) d Q d v \tag{array}
\end{equation*}
$$

Note the simularity to eq (413) which is that equation used by Chinkov (1987) Noting that the system has a stochastic component which is dominant at the values of $\mu_{c l}$ we are considering and that the system is formed by gaussians (brought about by the Green's function solution to the diffusion equation (Kervorkian (1990))) then the process can be considered gaussian with the following density function

$$
\begin{equation*}
\frac{1}{\sqrt{2 \pi \sigma}} \exp \left[\frac{(y+2 \pi)^{2}}{2 \sigma}\right] \tag{4112a}
\end{equation*}
$$

and characteristic function

$$
\begin{equation*}
\exp \left[\iota m(y+2 n \pi)-\frac{1}{2} m^{2} \sigma\right] \tag{array}
\end{equation*}
$$

with the identity

$$
\begin{equation*}
\frac{1}{\sqrt{2 \pi \sigma}} \sum_{n=-\infty}^{\infty} \exp \left[\frac{(y+2 n \pi)^{2}}{2 \sigma}\right]=\frac{1}{2 \pi} \sum_{m=-\infty}^{\infty} \exp \left[\iota m(y+2 n \pi)-\frac{1}{2} m^{2} \sigma\right] \tag{array}
\end{equation*}
$$

This changes the above duffusion expression to

$$
D_{c l}=T \xrightarrow{\operatorname{llm}_{\infty}}(2 T)^{-1} \sum_{m_{T}=-\infty}^{\infty} \sum_{m_{1}=-\infty}^{\infty} \prod_{i=0}^{T-1} \int_{0}^{2 \pi} \frac{d Q_{i}}{2 \pi} S_{T}^{2}
$$

$$
\begin{equation*}
\exp \left[\sum_{\jmath=1}^{T} \frac{-1}{2} m^{2} \sigma+\iota m_{\jmath}\left(Q_{\jmath}-Q_{\jmath-1}-v_{o}-S_{\jmath-1}\right)\right] \tag{array}
\end{equation*}
$$

We now make the approxamation that, for the turbulent diffusion to occur, the kuck term has to be such that it dominates over the stable term Remember that the phase space has less than $1 \%$ occupied by stable orbits at $\mu_{c l}=20 * \mathrm{k}$ Thus $S_{-1}$ can now be considered to be a summation over just the sine terms The number, $n$, of $m_{j}^{\prime} s$ used in any calculation determines the correlation $C(n)$ of the deffusion For all the $m_{j}^{\prime} s$ at zero we have the effective quasi-linear diffusion denoted $D_{Q L}$ and, by Rechester \& White (1980), the $S_{T}^{2}$ can be approxamated by an integral over $2 \pi$ (the effective penodicity of the system) which gives $\xi^{2} \mu_{c l}^{2} / 2$ then the quasi-linear diffusion is

$$
\begin{equation*}
D_{Q L}=\frac{1}{4} \xi^{2} \mu_{c l}^{2} \tag{4114}
\end{equation*}
$$

If we set the problem up so that just one $m_{j}$ is not zero we find that the integral term is

$$
\begin{equation*}
\exp \left[-\frac{1}{2} m^{2} \sigma-\iota m_{\imath} v_{0}\right] \int_{0}^{2 \pi} \frac{d Q_{2}}{2 \pi} \exp \left[\iota m_{\imath} Q\right]=0 \tag{4115}
\end{equation*}
$$

So the correlation term $C(1)$ is zero as predicted by Chinkov for the standard mapping (Chirıkov (1987))

In formulating the higher correlations we must remember that any expression for the diffusion coefficient $D_{c l}$ has to be independent of the velocity used as the intial probability distribution function In order for this to be satisfied the sum of all $m_{j}^{\prime} s$ for the $v_{0}$ term must come to zero Furthermore as the sum of the square of the $m_{j}^{\prime} s$ times $\sigma$ tend to zero very quickly then we have chosen to keep the $m^{\prime}, s$ within the bounds $\pm 2$ If we set two consecutive $m_{j}^{\prime} s$ not equal to zero and all the rest zero then we are effectively calculating the $C(2)$ correlation (Chirikov (1987)) For this correlation we choose $m_{i}= \pm 1$ and $m_{t+1}=-m_{1}$ ) Ourintegral in the diffusion equation can now be re-arranged to remove all terms independent of the $Q_{2}^{\prime} s$

$$
\begin{equation*}
\exp \left[-\sum_{\jmath=1}^{T}\left(\frac{1}{2} m_{\jmath}^{2} \sigma+\iota m_{\jmath} v_{o}\right)\right] \int_{0}^{2 \pi} \frac{d Q_{1}}{2 \pi} S_{T}^{2} \exp \left[\sum_{\jmath=1}^{T} \iota m,\left(Q_{\jmath}-Q_{\jmath-1}-S_{\jmath-1}\right)\right] \tag{4116}
\end{equation*}
$$

The integral now becomes

$$
\begin{align*}
& \exp \left[-\frac{1}{2}\left(m_{\imath+1}^{2}+m_{\imath}^{2}\right) \sigma-\iota\left(m_{\imath}+m_{\imath+1}\right) v_{o}-\iota\left(m_{\imath}+m_{\imath+1}\right) S_{\imath-1}\right] \\
& \quad \int_{0}^{2 \pi} \frac{d Q_{\imath}}{2 \pi} S_{T}^{2} \exp \left[\iota\left(\left(m_{\imath}-m_{\imath+1}\right) Q_{\imath}-m_{\imath+1} \sin \left(Q_{\imath}\right)\right)\right] \tag{4117}
\end{align*}
$$

From the definition of $m_{1}$ and $m_{\imath+1}$ we can determine that $m_{i}+m_{i+1}=0, m^{2}+m^{2}=2$ and $m_{2}-m_{i+1}=2 m_{i}$ Therefore the diffusion term for this $C(2)$ approximation is independent of $v_{0}$, as we wanted it to be

$$
\begin{gather*}
D_{C(2)}=\quad \lim _{T \rightarrow \infty} \neq \frac{1}{2} \exp [-\sigma]\left\{\int_{0}^{2 \pi} \frac{d Q_{2}}{2 \pi} S_{T}^{2} \exp \left[+\iota\left(-2 Q_{2}+\sin \left(Q_{2}\right)\right)\right]\right. \\
\left.+\int_{0}^{2 \pi} \frac{d Q_{2}}{2 \pi} S_{T}^{2} \exp \left[+\iota\left(2 Q_{2}-\sin \left(Q_{2}\right)\right)\right]\right\} \tag{4118}
\end{gather*}
$$

The integral can now be replaced by the second order Integer Bessel Function, $J_{2}(z)$ (see Gradshteyn and Ryzhık (1965) eq $8411(1)$ ) So our diffusion term becomes

$$
\begin{equation*}
D_{C(2)}=-\frac{1}{2} \xi^{2} \mu_{c l}^{2} J_{2}\left(\xi \mu_{c l}\right) e^{-\sigma} \tag{array}
\end{equation*}
$$

This term is added to the quasi-hnear term in eq (4 114)

$$
\begin{equation*}
D_{c l}=D_{Q L}+D_{C(2)}=\frac{1}{4} \xi^{2} \mu_{c l}^{2}\left[1-2 J_{2}\left(\xi \mu_{c l}\right)\right] \tag{4120}
\end{equation*}
$$

It is also possible to equate $m_{2-2}$ with $m_{2}$ (effectively a $C(3)$ correlation) We again set $m_{i-2}= \pm 1$ and $m_{i}=-m_{\imath-2}$ This has the effect of squaring a sum of two identical integrals

$$
\begin{gather*}
D_{C(3)}^{(1)}=\lim _{T \rightarrow \infty}-\frac{1}{2} \exp [-\sigma]\left\{\int_{0}^{2 \pi} \frac{d Q_{2}}{2 \pi} S_{T}^{2} \exp \left[+\iota\left(-Q_{2}+\sin \left(Q_{2}\right)\right)\right]\right. \\
\left.+\int_{0}^{2 \pi} \frac{d Q_{2}^{2}}{2 \pi} S_{T}^{2} \exp \left[+\iota\left(Q_{2}-\sin \left(Q_{t}\right)\right)\right]\right\} \tag{41.21}
\end{gather*}
$$

The superscript on the $D$ denotes it is the first part of the $C(3)$ correlation Again using the 1 dentity from Gradshteyn and Ryzhik (1965) we can substitute integer order Bessel functions of the first kind for these integrals The result is a $J_{1}^{2}$ term

$$
\begin{equation*}
D_{C(3)}^{(1)}=-\frac{1}{2} \xi^{2} \mu_{c l}^{2} J_{1}^{2}\left(\xi \mu_{c l}\right) e^{-\sigma} \tag{array}
\end{equation*}
$$

Again the above is in agreement with half of the $C(3)$ correlation term put forward by Chirikov (1987), the other half we will obtann now by considering a wider range of $m^{\prime}$ s This second half of the $C(3)$ correlation relates not two iterates but three such that $m_{1}$ is correlated to both $m_{i-1}$ and $m_{i-2}$, unlike the previous term where $m_{1}$ was related to just $m_{\imath-2}$ Again we need to keep the $m_{j}^{\prime} s$ small as the $\Sigma m_{j}^{2} \sigma$ term tends to zero very quckly Accounting for the $v_{0}$ constraint $\left(\Sigma m_{j}=0\right)$ we follow the example of Rechester and White (1980) and choose the following values for the $m_{j}^{\prime} s \quad m_{i}= \pm 1, m_{2+1}=-2 m_{1}$ and $m_{\imath+1}=m_{\imath} \quad$ One can see immediately that $\Sigma m_{\jmath}$ is indeed zero satisfying our $v_{0}$ independence constraint The $\Sigma m_{j}^{2}$ is six giving an $\exp (-3 \sigma)$ weighting for this term Furthermore the integrals obtained from the substitution of these $m_{j}^{\prime} s$ are of a simular form to the previous correlation terms but yield a $J_{3}$ term 1 e

$$
\begin{align*}
D_{C(3)}^{(2)}= & \xrightarrow{\operatorname{lom}} \\
& \frac{1}{2} \exp [-3 \sigma]\left\{\int_{0}^{2 \pi} \frac{d Q_{2}}{2 \pi} S_{T}^{2} \exp \left[+\iota\left(-3 Q_{2}+\sin \left(Q_{2}\right)\right)\right]\right.  \tag{4123}\\
& \left.+\int_{0}^{2 \pi} \frac{d Q_{2}{ }^{2}}{2 \pi} S_{T}^{2} \exp \left[+\iota\left(3 Q_{2}-\sin \left(Q_{2}\right)\right)\right]\right\}
\end{align*}
$$

which gives using the integral identity

$$
\begin{equation*}
D_{C(3)}^{(2)}=\frac{1}{2} \xi^{2} \mu_{c l}^{2} J_{3}^{2}\left(\xi \mu_{c l}\right) e^{-3 \sigma} \tag{4124}
\end{equation*}
$$

The complete diffusion term is now the sum of all the correlation terms thus obtaned

$$
\begin{gather*}
D_{c l}=D_{Q L}+D_{C(2)}+D_{C(3)}^{(1)}+D_{C(3)}^{(2)}= \\
\frac{1}{4} \xi^{2} \mu_{c l}^{2}\left[1-2 J_{2}\left(\xi \mu_{c l}\right) e^{-\sigma}-2 J_{1}^{2}\left(\xi \mu_{c l}\right) e^{-\sigma}+2 J_{3}^{2}\left(\xi \mu_{c l}\right) e^{-3 \sigma}\right] \tag{4125}
\end{gather*}
$$

This equation accounts for the correlations up to and including $C(3)$ We can add part of another correlation, $C(4)$, whuch has been shown by Chirkov (1987) to tend to zero slower than the $C(3)$ correlation To obtain this term we use choose values for the $m^{\prime}, s$ satisfying our constrants for $\sigma$ and $v_{0}$ The values chosen are $m_{i}=m_{1+3}= \pm 1$ and $m_{\imath+1}=m_{t+4}=-m_{\imath} \quad$ As before $\Sigma m_{\jmath}=0$ and $\Sigma m^{2}=4$ We end up with a product of two sets of integrals identical to eq ( 4118 ) (except in the multipher before the $\sigma$ term) This $C(4)$ term gives us

$$
\begin{equation*}
D_{C(4)}=\frac{1}{2} \xi^{2} \mu_{c l}^{2} J_{2}^{2}\left(\xi \mu_{c l}\right) e^{-2 \sigma} \tag{array}
\end{equation*}
$$

Our diffusion equation now takes on the form

$$
\begin{gather*}
D_{c l}=D_{Q L}+D_{C(2)}+D_{C(3)}^{(1)}+D_{C(3)}^{(2)}+D_{C(4)}= \\
\frac{1}{4} \xi^{2} \mu_{c l}^{2}\left[1-2 J_{2}\left(\xi \mu_{c l}\right) e^{-\sigma}-2 J_{1}^{2}\left(\xi \mu_{c l}\right) e^{-\sigma}+2 J_{3}^{2}\left(\xi \mu_{c l}\right) e^{-3 \sigma}+2 J_{2}^{2}\left(\xi \mu_{c l}\right) e^{-2 \sigma}\right] \tag{array}
\end{gather*}
$$

Churkov has shown (Chinkov (1987)) that the the $C(3)$ component decays as $\left|\xi \mu_{c l}\right|^{-2}$ whereas the $C(4)$ component decays as $\left|\xi \mu_{c l}\right|^{-1}$ Thus for large $\xi \mu_{c l}$ the diffusion equation is best described by

$$
\begin{equation*}
D_{c l}=\frac{1}{4} \xi^{2} \mu_{c l}^{2}\left[1-2 J_{2}\left(\xi \mu_{c l}\right)+2 J_{2}^{2}\left(\xi \mu_{c l}\right)\right] \tag{4128}
\end{equation*}
$$

where we can neglect the $\sigma$ term decay due to the fact that $\sigma$ is usually kept very small (the value chosen by Rechester and White (1980) being just $1 \times 10^{-4}$ ) This is exactly the form given by Chirikov (1987) using his simplified argument based on correlations and by Israelev (1990)

We will now show how remarkably well the above fits that data obtanned from numerical sımulations To numencally evaluate the diffusion coefficient we use the expression in eq (4 1 3) To accomplsh ths we calculate the mean energy for an ensemble of points in the stochastic layer The results showed us that the growth $m$ energy is increasing linearly with increasing $n$ ( $n$ being the iteration counter and hence time) and so can be written as

$$
\begin{equation*}
E_{n}\left(\mu_{c l}, n\right) \equiv B\left(\mu_{c l}\right) n \equiv B\left(\mu_{c l}\right) \tau \tag{4129}
\end{equation*}
$$


$F_{2 g} 41$ Four plots of the duffusion coefficient, $D_{c l}\left(\mu_{c l}\right)$, vs the kick strength, $\mu_{c l}$ for various values of the parameter $K \quad K=01 \mathrm{nn}(\mathrm{a}), 025 \mathrm{~m}(\mathrm{~b}), 05 \mathrm{in}$ (c) and 075 in (d) The sold hne is the theoretical curve predicted by eq (4128) and the asterisks are the calculated points The fit is very good except for small values of $K \mu_{c l}$ as can be seen from (a) where the fit is the worst It was the deviations in (a) which led us to discover the anomalous diffusion peaks discussed later

Care was taken in the above calculation to avoid using periodic points or quasi- periodic points as these points' non-deffuse behaviour would affect the result These points were
numencally filtered out The next step was to obtan the best fit hne to this linear graph of energy vs time ( e its denvative) This derivative is equal to twice the diffusion coefficient Why twice ${ }^{\mathcal{P}}$ The reason for this is that we calculated the energy of the system with time not the change in momentum squared The energy depends on $p$ and $q$ (or $x$ and $y$ for the mapping) and as one can show very easily that $x$ and $y$ can be interchanged without compromise then the energy is twice that value of the change in momentum So the theoretical solid curve shown in each of the figures in fig $41(a)-(d)$ is twice that of the diffusion coefficient expressed in eq ( 4128 ) To alleviate problems associated with individual orbits and not the global nature of the system a set of 60000 initial points were taken in the stochastic region close to one of the hyperbolic period four fixed points All points were tested for quasi-penodicity and filtered out if found to be such The resultant set was iterated for 20000 tıme steps The four figures in fig 41 show the diffusion coefficient, $D_{c l}$, as a function of the kıck strength, $\mu_{c l}$, for a range of the parameter $K$ The fit is remarkable except at small values of $\mu_{c l}$ where the stable component of the motion affects the assumptions made in deriving eq (4128)

Despite our success with the expression above for the diffusion coefficient we nevertheless found this coefficient to diverge from the predicted value at certan regularly spaced values of the kuck strength, $\mu_{\mathrm{cl}}$ These divergences manufested themselves in the form of delta function hke spikes on the normal turbulent duffusion predicted above with the values of the kick strength corresponding to a positive integer times the system's periodicity This discovery led to a revision of the system's behaviour when the kick strength corresponded to these periodicity related values We dubbed these spikes in the diffusion coefficient Resonant Enhanced Duffusion We now present our argument to explain these periodic spikes

## 42 Resonance Enhanced Diffusion

To proceed any further with our argument for the resonant enhanced diffusion we need to consider the nature of the effect of the system's penodicity on the dynamics of the system

We have already pointed out at the beginning of the last section that the $p$ and $q$ can be effectively restricted to region of $\pm \pi /(2 K)$ This can be clearly llustrated as follows In showing why the stable component of the motion can be ignored in the preceding section we stated that when $\mu_{c l}$ is much greater than the stable component then the turbulent analysis of Rechester and White (1980) can be used However as $p$ and $q$ effectively visit the whole phase space then they can, for a fixed value of $\mu_{c l}$, take on any value in the whole $2 D$ real plane seemingly negating our argument of the previous section This would imply that the previous argument would not describe the diffusion correctly This would be true if the dynamics in one region were unique to that region and no other However we have shown that the system is periodic and the dynamics in any periodic cell is representative of the system at large Therefore the effective values of $p$ and $q$ are restricted to the bounds imposed by the system's periodicity Thus a point $\left(q^{\prime}, p^{\prime}\right)$ in the phase space related to ( $q, p$ ) by the relation $p^{\prime}=p+n \pi / K$ and $q^{\prime}=q+m \pi / K$ will have identical dynamics to the points ( $q, p$ ) except its energy will be greater by an amount dependent on the $n \pi / K$ and $m \pi / K$ As a result of this, there is a strong dynamic correlation between the points ( $q, p$ ) and ( $q^{\prime}, p^{\prime}$ ) This reasoning is the very heart of the argument we present here We will use the correlations between sets of $\left(q^{\prime}, p^{\prime}\right)$ to show how they give rise to these anomalous peaks in the diffusion coefficient and why these peaks correspond to a kick strength equal to a positive integer times the system's periodicity

We have shown in the previous section how the turbulent diffusion (which is that when the stable component is neghgible) is derived by considening correlations between iterates at different tımes and relating them to the iterate we are interested in Here we will show, using the observation in the previous paragraph, how large summations over identical arguments give nise to the anomalous diffusion

Let us recall our equation for the diffusion coefficient, $D_{c l}$

$$
\begin{gather*}
D_{c l}=\operatorname{lom}_{T \rightarrow \infty}(2 T)^{-1} \sum_{m_{T}=-\infty}^{\infty} \sum_{m_{1}=-\infty}^{\infty} \prod_{\imath=0}^{T-1} \\
\exp \left[-\sum_{\jmath=1}^{T}\left(\frac{1}{2} m^{2} \sigma+\iota m, v_{o}\right)\right] \int_{0}^{2 \pi} \frac{d Q_{2}}{2 \pi} S_{T}^{2} \exp \left[\sum_{\jmath=1}^{T} \iota m_{\jmath}\left(Q_{\jmath}-Q_{\jmath-1}-S_{\jmath-1}\right)\right] \tag{421}
\end{gather*}
$$

Now of we have strong dynamuc correlations the term $S_{T}$, which is a summation up to $T$ of the $\xi \mu_{c l} \sin \left(K Q_{i}\right)$, will consist of a summation of the same value again and again, 1 e if the $Q_{2}$ are related by the expression

$$
\begin{equation*}
Q_{2}=Q_{o}+\jmath \pi / K \tag{422}
\end{equation*}
$$

(where $0 \leq|\jmath| \leq|\imath|, \jmath \in \mathbf{Z}$ ) then the summation gives us $T\left(\xi \mu_{c l} \sin \left(K Q_{z}\right)\right)^{2}$ We would expect this summation to tend to infinty quicker than the summation over random $Q_{\text {t }}$ This is indeed the case as we found by taking the summations of the $S_{T}^{2}$ term for various values of the kuck strength Once properly scaled, the result followed the numerically evaluated result remarkably well so we conclude that strong correlations of the $Q_{2}$ bring about the enhanced diffusion at the observed values of the kick strength Furthermore, the rapidity at which the coefficient $D_{c l}$ would tend to infinity could also depend on the behaviour of the other terms in the expression in eq (421) For the $D_{Q L}$ in the previous section the $m_{j}^{\prime} s$ are all zero so we have no $e^{-\sigma}$ terms to worry about This term we would expect to tend quackest to infinty whereas the $D_{C(2)}, D_{C(3)}$ and the $D_{C(4)}$ terms would tend to infinty slower due to the presence of non zero $m_{j}^{\prime} s$ in the $\sigma$ term As the number of non-zero $m_{j}^{\prime} s$ is increased (while still accounting for the restriction, $\Sigma m_{j}=0$ for the $v_{o}$ term to be omitted) infinty is approached slower but we see no reason why it would not be acheived eventually So the overall result is that of points in the phase space exist so that they are highly dynamically correlated then we would expect their contribution to the diffusion coefficient to be such that $D_{c l}$ would tend to infinty

We have seen this anomalous diffusion to anse when the kick strength is very close or equal to a positive integer times the system's periodicity This type of diffusion has been known to exast for this type of system since the work of Carey et al (1981a \& 1981b) on periodic area preserving maps and Karney et al (1982) on the effect of noise on the standard map We will here refer to a more recent analysis by Ishizaki et al (1989) which deals with the specific case of the kicked rotator In this paper they refer to accelerator modes islands which they found to exast in the stochastic layer of the kicked rotator The crux of their argument is that these islands impart an integer number of $2 \pi$ (their system had a periodicity of $2 \pi$ ) to the momentum of stochastic layer orbits which reside close to


Frg 42 The resonance enhanced diffusion for (a) $K=01$ and (b) $K=025$ The spikes can be seen to occur at positive integer multiples of the system's perrodicity ( $10 \pi$ for $k=01 \& 4 \pi$ for $K=025$ )
these islands This increase of $2 \pi$ translates points ( $q, p$ ) to their corresponding cousins ( $q^{\prime}, p^{\prime}$ ) whose dynamics are identical Furthermore these accelerator islands were found by Ishizakı et al (1989) to be of the same periodicity as the system This results in some orbital points ( $q, p$ ) being highly dynamically correlated to each other and to have a much higher overall mean energy than other un-accelerated orbits The stability of these islands
is crucial to the accelaration and Ishzaki et al (1989) found them to be stable for a small range of kuck strengths very close to a positive integer times the system's penodicity

Therr results are apphcable here because of the periodicity of our phase space and because hike the kicked rotator the dynamics of one periodic cell are identical to that of another Furthermore by varying $K$ we can change our phase space periodicity and hence show that any kucked system which has a penodic phase space wll have this anomalous diffusion when the kick strength is in the vicinity of the system's penodicity

We end our analysis of the system's duffusion here with the point that this resonant enhanced duffusion has, to our knowledge after searchng the relevant literature, not been reported at all for any system outside the kacked rotator This fact, despite exhaustive studies of this system by many distingushed researchers (Isralev, Berman, Chernikov etc), highlights that our present knowledge of this system is still very much incomplete

## CHAPTER 5 <br> THE QUANTUM MAPPING

Having described the classical limit of the Hamiltonan we now proceed to analyse the fully quantum limit, 1 e when $\hbar$ is finte and non-zero Our system Hamultoman is time-dependent and therefore the approach taken will be different from that for time independent systems such as the quantum billards problem (Berry et al (1986)) and others (Robnuk et al (1986), Tomsovic et al (1993)) The quantum mapping derived in this chapter is obtained using an approach sımular to that of Fox et al (Fox \& Lan (1990), Fox \& Elston (1994)) in their analysis of the kucked rotator This is, to our knowledge, the first time this has been done for the kıcked harmome oscillator and the resultant mapping is therefore unique in this regard

## 51 Derivation Of The Quantum Mapping

Thus section essentally deals with the derivation of the Fully quantum mapping from the system hamiltoman The general approach here is very simular to Fox \& Lan (1990) in their derivation for the kicked rotator but the added potential in our problem results in a much more complex system and final mapping Initially we will examine the undriven harmonic oscillator and look at its eigenfunctions and their properties because these will be needed for the driven harmome oscillator As such we call on much of the information contanned mapter 12 of Quantum Mechanics Volume 1 by Messiah (1976)

Agann we start off with the Hamiltoman of the Quantum Harmonic Oscillator driven (or Kıcked) by a temporally discrete potential

$$
\begin{equation*}
H_{T}(p, q)=\frac{p^{2}}{2 M}+\frac{1}{2} M \omega_{o}^{2} q^{2}+\mu_{q} \operatorname{Cos}(k q) \sum_{n=1}^{\infty} \delta(t-n \tau) \tag{array}
\end{equation*}
$$

We need to recall from chapter 2 our defintion of the integrable hamultoman $H_{o}(p, q)$ for the undriven system

$$
\begin{equation*}
H_{o}(p, q)=\frac{p^{2}}{2 M}+\frac{1}{2} M \omega_{o}^{2} q^{2} \tag{array}
\end{equation*}
$$

Our task now is to find a function $\psi(q, t)$ which satisfies the tıme dependent Schrodinger wave equation Through thus equation we can obtan a form for the wavefunctions of the system The equation is given below

$$
\begin{equation*}
H\left(\frac{d}{d q}, q\right) \psi(q, t)=i \hbar \frac{\partial \psi}{\partial t}(q, t) \tag{array}
\end{equation*}
$$

The functions which solve the above are many and varied but we are interested in a subclass of these solutions which have the property that $H_{o} \psi_{n}=E_{n} \psi_{n}$ where $E_{n}$ is the energy of the system's $n^{\text {th }}$ energy level These functions $\psi_{n}$ are called the eigenfunctions of the hamltonaan $H_{o}$ and will be denoted henceforth as $\phi_{n}(q)$ It should also be observed that the $\phi_{n}(q)$ are linearly independent and form a complete basis for the system Thus any other function $\psi(q, t)$ satisfying eq (513) is made up of a superposition of the $\phi_{n}(q)$ with coefficients $A_{n}(t)$ as multiphers and can be written

$$
\begin{equation*}
\psi(q, t)=\sum_{n=1}^{\infty} A(t) \phi_{n}(q) \tag{514}
\end{equation*}
$$

where the modulus of the $A_{n}$ 's gives the probability amplitudes for the $\phi_{n}$ in the function $\psi$ This way we can solve for the elgenfunctions, $\phi_{n}$, and use them to find any other function, $\psi$, which satisfies the wave equation So our first hurdle is to obtain an expression for the elgenfunctions To do this it is necessary to recall some of the defintions from chapter 2, namely those of eq (213), 1e

$$
\begin{equation*}
q=Q \sqrt{\frac{\hbar}{M \omega_{o}}} \quad \& \quad p=P \sqrt{M \hbar \omega_{o}} \tag{array}
\end{equation*}
$$

We now express the momentum operator as a function of $Q$ as per its defintion The normal $p$ operator is defined as $p=\frac{\hbar}{\iota} \frac{d}{d q}$ which allows us, employing the definitions in eq (515), to express $P$ as $\frac{1}{6} \frac{d}{d Q}$ Furthermore we can now to introduce a new set of elgenfunctions $u_{n}(Q)$ defined to be $\sqrt[4]{\frac{\hbar}{M \omega_{o}}} \phi_{n}(q)$ where the fourth root term is to allow $Q$ to be substituted for $q$ (Messiah (1976)) The effect of introducing this latter set of eigenfunctions is to change our expression of $H_{o}$ to

$$
\begin{equation*}
H_{o}\left(\frac{d}{d Q}, Q\right) u_{n}(Q)=\frac{1}{2}\left(Q^{2}-\frac{d^{2}}{d Q^{2}}\right) u_{n}(Q) \hbar \omega_{o}=E_{n} u_{n}(Q) \tag{516}
\end{equation*}
$$

Our next step is necessary for our formulation of the undriven system's eigenfunctions Our previous defintions (in chapter 2) of the anmhilation and creation operators, a and $a^{\dagger}$, are extended to include the explicit form of $P$ in terms of $Q$

$$
\begin{equation*}
a(P, Q)=\frac{1}{\sqrt{2}}(Q+\iota P)=\frac{1}{\sqrt{2}}\left(\frac{d}{d Q}+Q\right) \tag{517a}
\end{equation*}
$$

and

$$
\begin{equation*}
a^{\dagger}(P, Q)=\frac{1}{\sqrt{2}}(Q-\iota P)=\frac{1}{\sqrt{2}}\left(Q-\frac{d}{d Q}\right) \tag{517b}
\end{equation*}
$$

The form we have chosen for $a^{\dagger}$, the creation operator, is such that when $a^{\dagger}$ operates on the $n^{\text {th }}$ eigenfunction (with the system thus being solely in the $n^{\text {th }}$ state) it raises the system to the $(n+1)^{\text {th }}$ level with the $(n+1)^{\text {th }}$ elgenfunction being that wavefunction describing the system now, 1 e $\left.a^{\dagger} \mid n>=\sqrt{n+1}\right\}(n+1)>$ Thus $1 t$ creates a state and hence the name creation operator The anmbulation operator, a, does the opposite in that it lowers the state of the system from the $n^{\text {th }}$ level to the $(n-1)^{\text {th }}$ level, 1 e $a|n>=\sqrt{n}|(n-1)>$ Like the creation operator, the designation annihilation operator for a mirrors its affect on the system's state The justification for the above properties is given in Appendix $B$ where we also show that $\left(a, a^{\dagger}\right)=1$ when operating on any pure eigenstate $n$ Therefore the state remanns unchanged It also follows that $a \mid 0>=0$ (you can't go lower than the zeroth level or ground state) and $|n\rangle$, the $n^{\text {th }}$ level, is given by the recursion relationship $\left.\frac{1}{\sqrt{n^{n}}}\left(a^{\dagger}\right)^{n} \right\rvert\, 0>$ These are two definitive results that we will use to obtain the form of the elgenfunctions From the action of the anniblation operator on the zeroth level we get

$$
\begin{equation*}
a \left\lvert\, 0>\equiv \frac{1}{\iota \sqrt{2}}\left(\frac{d}{d Q}+Q\right) u_{o}(Q)=0\right. \tag{518a}
\end{equation*}
$$

or equvalently

$$
\begin{equation*}
\left(\frac{d}{d Q}+Q\right) u_{o}(Q)=0 \tag{518b}
\end{equation*}
$$

The solution of eq $(518 b)$ is $u_{0}(Q)=C * \exp \left(-\frac{1}{2} Q^{2}\right)$ with the constant $C$ being evaluated using the property $\left\langle u_{0}, u_{0}\right\rangle=1$ and that the solution to the integral $\int_{0}^{\infty} \exp$ ( $\left.Q^{2}\right) d Q=\frac{\sqrt{\pi}}{2} C$ is found to be $\frac{1}{4 \sqrt{\pi}}$ So $u_{0}(Q)=\frac{1}{4 \sqrt{\pi}} \exp \left(-\frac{1}{2} Q^{2}\right) \quad$ We now substitute this solution into our recursion relationship to obtain the general solution for all $u_{n}(Q)$,

$$
\begin{equation*}
\frac{\iota^{n}}{\sqrt{2^{n} \sqrt{\pi\left(n^{\prime}\right)}}}\left(Q-\frac{d}{d Q}\right)^{n} u_{o}(Q)=u_{n}(Q) \tag{array}
\end{equation*}
$$

where we find, after replacing $u_{0}(Q)$ with explicit its form, $u_{n}(Q)$ takes the form

$$
\begin{equation*}
\frac{\iota^{n}}{\sqrt{2^{n} \sqrt{\pi}\left(n^{\prime}\right)}}\left(Q-\frac{d}{d Q}\right)^{n} \exp \left(-\frac{1}{2} Q^{2}\right)=u_{n}(Q) \tag{array}
\end{equation*}
$$

We can now indentify part of the above expression with a type of orthogonal polynomial called Hermite Polynomials, $H_{n}(Q)$ We can do this because of the following recursion relationships (Gradshteyn \& Ryzhuk (1965))
$1 \quad \frac{d H_{n}(x)}{d x}=2 n H_{n-1}(x)$
2

$$
H_{n+1}(x)=2 x H_{n}(x)-2 n H_{n-1}(x)
$$

Combinung these two identities gives

$$
\begin{equation*}
H_{n+1}(x)=\left(2 x-\frac{d}{d x}\right) H_{n}(x) \tag{3}
\end{equation*}
$$

with the one we seek being the contmued substitution of $H_{n-k}(x)$ down to $H_{o}(x)$
4

$$
H_{n+1}(x)=\left(2 x-\frac{d}{d x}\right)^{n} H_{o}(x) \equiv\left(2 x-\frac{d}{d x}\right)^{n}
$$

because $H_{o}(x)=1$ It is easy to show that the Hermite polynomals, $H_{n}$, can be formed not just by $\left(2 x-\frac{d}{d x}\right)^{n} H_{o}(x)$ but also by the relationship $\left(Q-\frac{d}{d Q}\right)^{n} \exp \left(-\frac{1}{2} Q^{2}\right)$ (see eq 8950 in Table of Integrals, Series and Products by Gradshteyn and Ryzhu (1965) and section III in Appendix $B$ of Quantum Mechanics Volume 1 by Messiah (1976) We can thus substitute the hermite polynomials, $H_{n}(Q)$, into our equation for the eigenfunctions

$$
\begin{equation*}
\frac{\iota^{n}}{\sqrt{2^{n} \sqrt{\pi\left(n^{\prime}\right)}}} * H_{n}(Q) \exp \left(-\frac{1}{2} Q^{2}\right)=u_{n}(Q) \tag{array}
\end{equation*}
$$

This completes our expression for the eigenfunctions of the undriven integrable hamiltonuan Our task now is to see how the eigenfunction probability amplitudes vary in time for the complete non-integrable hamiltonian We are going to formulate the problem in such a way as to keep the eigenfunctions for the undriven osillator but generate a mapping which relates the probability amplitudes for these eigenfunction from one kick to the next

Between kicks the system hamiltonaan is essentially the integrable hamiltoman described above Thus any wavefunction $\psi(q, t)$ can be written as a superposition of the
$u_{n}(Q)$ or $\phi_{n}(q)$ depending on the phase space being considered The wave equation for driven (kicked) harmonic oscullator is

$$
\begin{gather*}
\iota \hbar \frac{\partial}{\partial t} \psi(q, t)=-\frac{\hbar^{2}}{2 M} \frac{\partial}{\partial q} \psi(q, t)+\frac{1}{2} M \omega_{o}^{2} q^{2} \psi(q, t)+ \\
\mu_{q} \operatorname{Cos}\left(k_{o} q\right) \sum_{n=-\infty}^{\infty} \delta(t-n \tau) \psi(q, t) \tag{array}
\end{gather*}
$$

As the delta $k z c k$ is of infinitesimal duration the integrable part of the hamultonian changes by a neglgible amount and can be neglected over a kuck duration Therefore the $\psi(q, t)$ change discontınuously across the delta kıcks by the amount $\mu_{q} \operatorname{Cos}\left(k_{o} q\right)$

$$
\begin{equation*}
\psi\left(q,(N \tau)^{+}\right)=\psi\left(q,(N \tau)^{-}\right) \exp \left(\iota \mu_{q} \operatorname{Cos}\left(k_{o} q\right)\right) \tag{array}
\end{equation*}
$$

The wavefunction between kicks is just that of the undriven oscillator and can be written in the form given by eq ( 514 ) Thus from just before the $N^{\text {th }}$ kack, when tıme $t=N \tau^{-}$, to just after 1 t, when tıme $t=N \tau^{+}$, we have

$$
\begin{equation*}
\psi\left(q,(N \tau)^{+}\right)=\sum_{p=1}^{\infty} A_{p}(N \tau)^{-} \phi_{p}(q) \exp \left(\iota \mu_{q} \operatorname{Cos}\left(k_{o} q\right)\right) \tag{array}
\end{equation*}
$$

The exponential term above can be substituted for an equivalent expression which includes a summation over bessel functions of integer order This identıty, $\exp \left(\iota \mu_{q} \operatorname{Cos}\left(k_{o} q\right)\right)=$ $\sum_{s=-\infty}^{\infty} \iota^{s} J_{s}\left(\mu_{q}\right) \exp \left(\iota s k_{o} q\right)$, can be found in Table of Integrals, Serses and Products by Gradshteyn and Ryzhik (1965) as eq 8511 (4) Our expression for the wavefunction changes to

$$
\begin{equation*}
\psi\left(q,(N \tau)^{+}\right)=\sum_{p=1}^{\infty} A_{p}(N \tau)^{-} \phi_{p}(q) \sum_{s=-\infty}^{\infty} \iota^{s} J_{s}\left(\mu_{q}\right) \exp \left(\iota s k_{o} q\right) \tag{5115}
\end{equation*}
$$

The time evolution from $(N \tau)^{+}$to $((N+1) \tau)^{-}$is that of the evolution of the harmonic oscllator itself Because each eigenfunction gives its corresponding energystate when operated on by $H_{o}, H_{o} \phi_{n}=E_{n} \phi_{n}$, then the evolution between the kicks is given by

$$
\begin{equation*}
\psi(q, t)=\sum_{p=1}^{\infty} A_{p}(t) \phi_{n}(g) \exp \left(-\iota \frac{E_{n}}{\hbar} t\right) \tag{5116}
\end{equation*}
$$

where $t$ is restricted to $N \tau^{+}$to $(N+1) \tau^{-}$and $N$ is an integer Furthermore $E_{n}=$ $\left(n+\frac{1}{2}\right) \hbar \omega_{o}$ The evolution from just before the $N^{\text {th }}$ kick to just before the $(N+1)^{\text {th }}$ kıck of the wavefuntion $\psi(q, t)$ can now be formally expressed It is the product of the eq ( 5116 ) (the evolution from the $N^{\text {th }}$ kıck to just before the next) with eq ( 5115 ) (the change over a kıck) Thus we have

$$
\begin{align*}
& \psi\left(q,((N+1) \tau)^{-}\right)= \\
& \sum_{p=1}^{\infty} A_{p}(N \tau)^{-} \phi_{p}(q) \sum_{s=-\infty}^{\infty} \iota^{s} J_{s}\left(\mu_{q}\right) \exp \left(\iota s k_{o} q\right) \exp \left(-\iota\left(p+\frac{1}{2}\right) \omega_{o} \tau\right) \tag{array}
\end{align*}
$$

However, we should also point out that we can express the wavefuntion $\psi(q, t)$ just before the $(N+1)^{\text {th }}$ kıck in terms of the $\phi_{n}(q)$ and the probability amplitudes at this later time

$$
\begin{equation*}
\psi\left(q,((N+1) \tau)^{-}\right)=\sum_{r=1}^{\infty} A_{r}\left(((N+1) \tau)^{-}\right) \phi_{r}(q) \tag{5118}
\end{equation*}
$$

This last equation allows us to express the probabilhty amplitude coefficients at time ( $(N+$ 1) $\tau)^{-}$as a function of the amplitude coefficients at time $(N \tau)^{-}$Thus

$$
\begin{align*}
& \sum_{r=1}^{\infty} A_{r}\left(((N+1) \tau)^{-}\right) \phi_{r}(q)= \\
& \sum_{p=1}^{\infty} A_{p}(N \tau)^{-} \phi_{p}(q) \sum_{s=-\infty}^{\infty} \iota^{s} J_{s}\left(\mu_{q}\right) \exp \left(\iota s k_{o} q\right) \exp \left(-\iota\left(p+\frac{1}{2}\right) \omega_{o} \tau\right) \tag{5119}
\end{align*}
$$

The problem with this expression is that it gives the relationshup between the summation over all the probabilty amplitudes times the elgenfunctions and does not give any information about individual probabilty amplitudes To rectify this we are going to take the expectation value of the above with each eigenfunction in turn This has the effect of giving us an explicit relationship between the amplitudes at time $((N+1) \tau)^{\text {- }}$ to those at tıme $N \tau^{-}$We know that the expectation value of $\phi_{m}$ with $\phi_{n},\left\langle\phi_{m} \mid \phi_{n}\right\rangle$, is $\delta_{n m}$ After taking the expectation value of the left hand side with an arbitrary eigenfunction $\phi_{m}$ we get

$$
\begin{equation*}
\sum_{r=1}^{\infty} A_{r}\left(((N+1) \tau)^{-}\right)<\phi_{m}(q)\left|\phi_{r}(q)\right\rangle=A_{m}\left(((N+1) \tau)^{-}\right) \tag{array}
\end{equation*}
$$

The right hand side does not sımplufy so easily when taking its expectation value with $\phi_{m}$ All terms which are functions of $q$ have to be brought together in the correct order inside the Bra-Ket symbohsing the expectation value

$$
\begin{align*}
& \sum_{p=1}^{\infty} A_{p}(N \tau)^{-} \sum_{s=-\infty}^{\infty} \iota^{s} J_{s}\left(\mu_{q}\right) \\
& \left.\quad \exp \left(-\iota\left(p+\frac{1}{2}\right) \omega_{o} \tau\right)<\phi_{m}(q) \right\rvert\, \phi_{p}(q) \exp \left(\iota s k_{o} q\right)> \tag{array}
\end{align*}
$$

As you can see from this expression the expectation is not between two othogonal eigenfunctions but between more compound expressions This renders vold the $\delta_{n m}$ property used for the left hand side of eq (5119) So by combining eqs (5120) and (5121) we can now obtain the expression we require to relate the ampltude coefficients at tume $(N+1) \tau^{-}$to those at tıme $t=N \tau^{-}$

$$
\begin{align*}
& A\left(((N+1) \tau)^{-}\right)=\sum_{p=1}^{\infty} A_{p}(N \tau)^{-} \sum_{s=-\infty}^{\infty} \iota^{s} J_{s}\left(\mu_{q}\right) * \\
& \left.\exp \left(-\iota\left(p+\frac{1}{2}\right) \omega_{o} \tau\right)<\phi_{m}(q) \right\rvert\, \phi_{p}(q) \exp \left(\iota s k_{o} q\right)> \tag{array}
\end{align*}
$$

The question now is What does the expression $<\phi_{m}(q) \mid \phi_{p}(q) \exp \left(\iota s k_{o} q\right)>m(5122)$ break down to and how does at affect the summation over $p$ in (5122) ${ }^{7}$ To answer this we need to refer back to eq (5111) and express the eigenfunctions' explicit form for $u_{n}(Q)$ in terms of $\phi_{n}(q), 1 e$

$$
\phi_{p}(q)=\frac{\iota^{p}}{\sqrt{2^{p} \sqrt{\pi}\left(p^{\prime}\right)}} * H_{p}(\gamma q) \exp \left(-\frac{1}{2} \gamma^{2} q^{2}\right)
$$

and with this expheit form we are going to evaluate the expectation values by integrating the wavefunctions over the whole of space thay occupy This space is effectively from $q=-\infty$ to $q=+\infty$ Remember that $q=$ real and that any polynomial in $q$ with real coefficients is also real With this in mind we can write $H_{p}^{*}$ as just $H_{p}$

$$
\begin{align*}
& \quad<\phi_{m}(q)\left|\phi_{p}(q) \exp \left(\iota s k_{o} q\right)\right\rangle= \\
& A_{m p} \iota^{m+p} \int_{-\infty}^{\infty} H_{m}(\gamma q) \exp \left(-\frac{1}{2} \gamma^{2} q^{2}\right) H_{p}(\gamma q) \exp \left(-\frac{1}{2} \gamma^{2} q^{2}\right) \exp \left(\iota s k_{o} q\right) d q \tag{array}
\end{align*}
$$

where $H_{m}(\gamma q)$ is the $m^{\text {th }}$ order Hermite polynomal and $A_{m p}=\frac{1}{\sqrt{\pi^{m+p}\left(p^{\prime}\right)\left(m^{\prime}\right)}}$ The additional term of $\exp \left(\iota s k_{0} q\right)$ can be accommodated into the other terms by multiplying
the equation by $\exp \left(-s^{2} k^{2} / 4 \gamma^{2}\right) \exp \left(s^{2} k^{2} / 4 \gamma^{2}\right)$ with the result that the $s k_{o} q$ term and the $\gamma q$ term can be grouped as a single perfect square The next step is to write $z=\gamma q-\iota \xi$ where $\xi=\frac{k_{0} s}{2 \gamma}$ and 1 s made up of system constants except for the summation counter s Hence the Hermite polynomals $H_{n}(\gamma q)$ go to $H_{n}(z+\iota \xi)$ whose argument is still real but is now expressed as a complex number Using this notation, the integral term becomes

$$
\begin{equation*}
A_{m p} \iota^{m+p} \exp \left(-s^{2} k^{2} / 4 \gamma^{2}\right) \int_{-\infty}^{\infty} H_{m}^{*}(z+\iota \xi) H_{p}(z+\iota \xi) \exp \left(-z^{2}\right) \frac{d z}{\gamma} \tag{5124}
\end{equation*}
$$

What is the solution to this integral ${ }^{8}$ How does the exponential term affect the orthogonality of the hermate polynomaals in integral term? The answer to these questions can be found in Gradshteyn and Ryzhuk's book Table of Integrals, Series and Products, eqn (7377) The form of the integral given is almost identical

$$
\begin{equation*}
\int_{-\infty}^{\infty} \exp \left(-x^{2}\right) H_{m}(x+y) H_{n}(x+z) d x=2^{n} \sqrt{\pi}\left(m^{1}\right) z^{n-m} L_{m}^{n-m}(-2 y z) \tag{5125}
\end{equation*}
$$

( $m \leq p$ ) We found that if $y$ and $z$ are replaced by $\iota \xi$ and $x$ by $z$ then we get the solution for eq (5124)

$$
\begin{equation*}
\int_{-\infty}^{\infty} \exp \left(-z^{2}\right) H_{m}(z+\iota \xi) H_{p}(z+\iota \xi) d z=2^{p} \sqrt{\pi}\left(m^{\prime}\right)(\iota \xi)^{p \sim m} L_{m}^{p-m}\left(2 \xi^{2}\right) \tag{5126}
\end{equation*}
$$

( $m \leq p$ ) where $L_{m}^{p-m}(x)$ is the Laguerre polynomal defined as $\sum_{r=0}^{m}(-1)^{r}\binom{m+(p-m)}{m-r} \frac{x^{r}}{r^{\top}}$ In order to be completely confident with this solution we had to evaluate numerically the integral above for sets of parameters and then substitute these parameters into the analytical form given This was done because we have a complex argument in the exponential and Gradshteyn and Ryzhuk's solution does not specify if the solution holds for this We found, using Mathematica for the numencal sımulations, that the solution is indeed valid for our integral' Furthermore if we stick with our onginal form, namely

$$
\int_{-\infty}^{\infty} H_{m}^{*}(\gamma q) H_{p}(\gamma q) \exp \left(-(\gamma q-\iota \xi)^{2}\right) d q
$$

we can use another identity from Gradshteyn and Ryzhuk namely eq (7374(7)) which is

$$
\begin{equation*}
\int_{-\infty}^{\infty} \exp \left(-(x-y)^{2}\right) H_{m}(x) H_{n}(x) d x=2^{n} \sqrt{\pi}\left(m^{\prime}\right) y^{n-m} L_{m}^{n-m}\left(-2 y^{2}\right) \tag{5127}
\end{equation*}
$$

( $m \leq p$ ) where by replacing $x$ by $\gamma q$ and $y$ by $\iota \xi$ we get

$$
\begin{equation*}
\int_{-\infty}^{\infty} \exp \left(-(\gamma q-\iota \xi)^{2}\right) H_{m}(\gamma q) H_{p}(\gamma q) d q=2^{p} \sqrt{\pi}\left(m^{\prime}\right)(\iota \xi)^{p-m} L_{m}^{p-m}\left(2 \xi^{2}\right) \tag{array}
\end{equation*}
$$

( $m \leq p$ ) which is the exactly the same result as from that in eq ( 5126 ) However these solutions are only defined for $m \leq \mathrm{p}$ So our equation for $A_{m}\left((n+1) \tau^{-}\right)$now takes on a more complete form

$$
\begin{gather*}
A_{m}\left(((N+1) \tau)^{-}\right)=\sum_{p=1}^{\infty} A_{p}(N \tau)^{-} \sum_{s=-\infty}^{\infty} \iota^{s} J_{s}\left(\mu_{q}\right) \exp \left(-\iota\left(p+\frac{1}{2}\right) \omega_{o} \tau\right) * \\
\frac{\iota^{m+p}}{\sqrt{\pi 2^{m+p}\left(p^{\prime}\right)\left(m^{\prime}\right)}} \exp \left(-s^{2} k^{2} / 4 \gamma^{2}\right) 2^{p} \sqrt{\pi}\left(m^{\prime}\right)\left(\iota \frac{k_{o} s}{2 \gamma}\right)^{p-m} L_{m}^{p-m}\left(s^{2} k_{o}^{2} / 2 \gamma^{2}\right) \tag{array}
\end{gather*}
$$

Some of the terms above cancel and others sumplify for example the inner summation over $s$ can be halved to contain just the positive integers because $\iota^{-s} J_{-s}\left(\mu_{q}\right)$ is equal to $\iota^{s} J_{s}\left(\mu_{q}\right)$ (the $s=0$ term is zero hence its exclusion) Furthermore only one other term in this summation depends on an odd power of $s$ (all the others depend on $s^{2}$ and are thus independent of the sign of $s$ ) With these sumplifications the final form is

$$
\begin{align*}
& A_{m}\left(((N+1) \tau)^{-}\right)=\sum_{p=m}^{\infty} A_{p}(N \tau)^{-} \sqrt{\frac{\left(m^{\prime}\right)}{\left(p^{\prime}\right)}} \sum_{s=1}^{\infty} \iota^{s+2 p} J_{s}\left(\mu_{q}\right) \exp \left(-\iota\left(p+\frac{1}{2}\right) \omega_{o} \tau\right) \\
& \quad * \exp \left(-s^{2} k_{o}^{2} / 4 \gamma^{2}\right) L_{m}^{p-m}\left(s^{2} k_{o}^{2} / 2 \gamma^{2}\right)\left\{\left(\frac{s k_{o}}{\sqrt{2} \gamma}\right)^{p-m}+\left(\frac{-s k_{o}}{\sqrt{2} \gamma}\right)^{p-m}\right\} \tag{5130}
\end{align*}
$$

which holds only for $m \leq p$ hence the lower lumit on the outer summation with $p$ now only summed from $m \leq p \leq \infty$ The complex conjugate of $A_{m}$ denoted $A^{*}$ is expected to be of a simular form but with the signs before the $\iota$ opposite to those above

To obtain the energy level spectrum as a function of time it is necessary to find the norm of the probabilty amplitudes as these are the multiphers in the summation over the energy levels of the undriven system Hence the energy of any $\psi(q, t)$, where $\psi(q, t)$
was defined in eq (514), is given by $\sum_{r=1}^{\infty}\left\|A_{r}(t)\right\|^{2} E_{r}$ As the $E_{r}$ are the energy levels of the undriven system then the energy evolves as

$$
\begin{equation*}
E_{\phi}\left(N \tau^{-}\right)=\sum_{r=1}^{\infty}\left\|A_{r}(N \tau)^{-}\right\|^{2}\left(r+\frac{1}{2}\right) \hbar \omega_{0} \tag{array}
\end{equation*}
$$

and as we can explicitly evaluate the $A_{r}$ over any kack we can calculate the energy of the system at any tıme for any set of parameters

We should bring the reader's attention to some remarkable comncidences between this mapping and the classical Firstly the presence of $\omega_{o} \tau$ which is just $\beta$ in the classical mapping (see chapter 2) Secondly we defined $\gamma$ to be the factor between $q$ and $Q$ such that $1 / \gamma=\sqrt{\hbar / M \omega_{0}}$ Therefore $k_{o} / \gamma$ is just our classical $K$ Thirdly we can relate $\mu_{q}$ to the classical $\mu_{c l}$ via the relationship defined in chapter 2 We can make comparisons quite easily now between our classical mapping and our fully quantum mapping above This we will do in the following two chapters

## 52 Numerıcal Consideratıons

Our purpose here is to hughlight the approach taken when numerically evaluating the mapping in eq ( 5130 ) Such an evaluation is necessary when one considers the compexaty of the mapping itself Futhermore by considering how functions within the mapping behave and interact with each other we intend to justify certain assumptions made Let us begin by describing the technique we use to iterate the mapping forward in tıme

If you examine closely the form of the mapping in eq ( $\left.\begin{array}{lll}5 & 1 & 30\end{array}\right)$ then you can see that only the $A_{p}$ on the right hand side is tıme dependent All other terms are time invanant and this useful observation will drastically reduce computing time in the evaluation of the mapping We can re-formulate the mapping as follows

$$
\begin{equation*}
A_{m}\left(((N+1) \tau)^{-}\right) \approx \sum_{p=1}^{\infty} U_{m p} A_{p}(N \tau)^{-} \tag{array}
\end{equation*}
$$

where $U_{m p}$ is the mp element of an infinte time-ındependent square matrix, $U$, which performs the task of stepping the system forward in time However we have a problem in
that only those elements of the matrix for $p \geq m$ are defined (from the dentity used in Gradshteyn and Ryzhik (1965) in eq (5 124 ) and (5 126)) To overcome this we present an argument to define those missing elements of the matrix U

We have already shown in the previous section that the Hermite polynomials are real as $q$ itself is real Therefore we substituted $H$ for $H^{*}$ in eq (5i23) because of the real valuedness of these polynomals As the integral in eq (5123) is a product of two polynomals then we can swap these polynomals around in the integral to give us the solution for $p \leq m$ We checked, using Mathematica, the correctness of the above and found the numerically that the two are identical Furthermore one can easly check that

$$
\begin{equation*}
U_{m p}=U_{p m} \tag{522}
\end{equation*}
$$

This symmetry allows us to calculate just one half of the matrix greatly reducing computer run tıme

Our task numerically is to evaluate thes matrix for a finte size and then use the relation in eq (5 21 ) to obtain the probabllity amplitudes at each tıme step As both the amplitudes, $A_{m}(t)$, and the evolution matrix, $U$, are complex then we have to spht the matrix into two parts (real and complex) and sımularly for the amphtudes Then we spht the evolution in eq ( 521 ) into the real and imaginary parts

$$
\begin{equation*}
A_{m_{\mathrm{re}}}\left(((N+1) \tau)^{-}\right)=\sum_{p=0}^{\infty}\left(U_{m p_{\mathrm{re}}} A_{p_{\mathrm{re}}}(N \tau)^{-}-U_{m p_{\mathrm{lm}}} A_{p_{\mathrm{lm}}}(N \tau)^{-}\right) \tag{523a}
\end{equation*}
$$

for the real

$$
\begin{equation*}
A_{m_{\mathrm{m}}}\left(((N+1) \tau)^{-}\right)=\sum_{p=0}^{\infty}\left(U_{m p_{\mathrm{re}}} A_{p_{\mathrm{im}}}(N \tau)^{-}+U_{m p_{\mathrm{tm}}} A_{p_{\mathrm{re}}}(N \tau)^{-}\right) \tag{523b}
\end{equation*}
$$

for the imaginary parts with, naturally, the sum being the complete complex probability amplatude at the specific kock

$$
\begin{equation*}
A_{m}\left(((N+1) \tau)^{-}\right)=A_{m_{\mathrm{re}}}\left(((N+1) \tau)^{-}\right)+\iota A_{m_{\mathrm{lm}}}\left(((N+1) \tau)^{-}\right) \tag{523c}
\end{equation*}
$$

The real and imaginary $U$ matrices are calculated first using a $C$ code program with a software hnk to Mathematica allowing high order Laguerre Polynomials to be evaluated
whereas a set of amphtude programs calculate the time evolution using the expressions in eqs (5 23 a ) and ( 523 b ) The complex conjugate of $A_{m}$ is then that of $A_{m}$ with the sign of $\iota$ changed We can then calculate the energy of the system at a specific time knowing that

$$
\begin{equation*}
E\left((N+1) \tau^{-}\right)=\sum_{n=0}^{\infty}\left\{A_{n_{\mathrm{re}}}^{2}\left(((N+1) \tau)^{-}\right)+A_{n_{\mathrm{rm}}}^{2}\left(((N+1) \tau)^{-}\right)\right\}(n+1 / 2) \hbar \omega_{o} \tag{524}
\end{equation*}
$$

This energy calculation is fundamental to understanding the system's evolution and the possible exastence of chaos

The numerical evaluation of the evolution matrix, $U$, is hampered by the infinite summation over $s$ (see eq ( 5130 )) Therefore we need to approxamate this by a truncated sum (if possible) It is well known (Isravev (1990)) that Integer Order Bessel functions $\left(J_{n}(z)\right)$ tend to zero if the order of the function is greater than twice the argument Therefore for any fixed argument, $z$, there exasts a cutoff for the Bessel functions at 2 z Our inner summation over $s$ could be given upper and lower hmits of $2 \mu_{c l}$ and $-2 \mu_{c l}$ respectively if no other functions exated in the summation which depended on s However we do have others which we need to consider The magnitude of the complex term could, if $s$ got large enough, hypothetically dominate but as $J_{s}(z)$ tend to zero if $|s| \gg z$ quicker than the complex term tends to $\infty$ then therr product would not diverge at large $s$ That still leaves us with a guess at the upper bound on $s$ The answer comes from the exponential term with the $-s^{2}$ argument This term dominates at large $|s|$ even over the Laguerre polynomals to the extent that the summation over $s$ can be truncated after about 20 to 30 terms (well within the scope of any computer) These properties were examined using Mathematica and we found that after 15 terms the decrease in magnitude was between 8 to 20 orders depending on the orders of the functions

All these precautions went into the numencal programming to ensure that the results obtaned would accurately mirror the matrix itself Further precautions will be outhned as necessary during the detaled analysis undertaken in the next chapter

## CHAPTER 6 <br> ANALYSIS OF THE QUANTUM MAPPING

This chapter is one of the more interesting for us in that it covers the results we obtained from our quantum mapping As this mapping has been derived using an approach never adopted for this system we presume the following analysis to be both novel and unique In this chapter we present the results obtained from the mapping obtanned in chapter 5 and compare its results with those from the classical mapping in chapters 2,3 and 4 The results presented are the time evolution of the probabilty amplatudes, the energy of the system for vanous kıck strengths, $\mu_{q}$, and the corresponding wavefunctions These results will be used along with those for the Wigner distribution in the following chapter to develop a correspondence between the quantum mapping and the classical mapping

## 61 Probability Amplitudes In Time

The quantum mapping in chapter 5 relates the probability amplitudes from one tıme step to the next Consequently, as the probablity amplitudes consist the primary data from the approach we're taking, then it is appropriate to begin our quantum analysis by examinng how they vary in time As previously discussed in chapter 5 section 2 (under Numencal Considerations) we have set the problem up so that the evolution matrix is time invariant and the size of this square matrix sets a limit on the order of the highest probability amphtude (and hence eigenfuncton) we can consider This will become important when iterating the quantum mapping over time

The diversity of our approach becomes clear when choosing which time evolutions to present as typical of the system's behaviour in this section We could, for example, choose a pure (in the sense of the undriven oscillator) even panty or odd panty eigenstate to begin with and examine its time evolution However, we have chosen a muxed state with a gaussian profile across the probability amplitudes centred on a low order undriven oscillator eigenstate, 1 e $m=10$ Our reasons for venturing so low were driven by practical considerations concernung the finite size of our evolution matrix, $U$ We restricted $U$ to a $350 \times 350$ matrix, for computational feasibility, with the result that, if the mitial state


Fig 61 The probability amplitudes for the mixed parity state of a gaussian centred on $\mathrm{m}=10$ with $\mu_{q}=000005$ Each of the successive plots (b)-(f) shows the system 300 time steps later The system remanns essentally unchanged regardless of the kıcking
spreads, or diffuses, so that the amplitudes near the $m=350$ boundary are becoming increasingly signficant, there is an artificial time limit imposed beyond which the results are boundary influenced and therefore untrustworthy

We would expect that if the kuck strength, $\mu_{g}$, was so small as to be neghgible then
its influence would be correspondungly weak In thas case the spread over the amplitudes would be very small if at all noticeable and the energy of the system would be constant This is the case for $\mu_{q}=000005$ where the gaussian profile is retanned for the duration of the time sample ( 2000 iterations) We present in figs $61(a)-(f)$ the probablity amplitudes every 300 time steps The variation from the original is shght and, as will be shown later, its energy vanes little from the initial state's energy This is what we see classically for those orbits near the ongin at low, near neghgible, kick strengths (refer to fig 22 in chapter 2) Classically, those orbits would remain circular in shape and vary little from the unkicked state

If, however, the kıcking is modest but not dominant, we would expect, classically, that those orbits very close to the ongin to reman essentally undisturbed forming an island which is invariant (the invariant cells of chapter 3) but that the orbits further out would become increasing disturbed and distorted (re Chapter 3) In the quantum case we find that the system's probabllty amphtudes settle down to specific patterns after a period of time The settling down time being dependent on the kick strength and how far away from the steady state distribution of amplitudes the initial state was We show such a case in figs $62(a)-(f)$ for $\mu_{q}=05$ Again the intervals between $(a),(b), \quad,(f)$ is 300 iterations The stable configuration is clear in figs $61(d),(e) \&(f)$ We will show in the next section how this configuration affects the energy and how simular steady configurations give rise to quast-energy levels, levels whose periodicity matches that of the kıck period and thus appear constant to the mapping (Israeilev (1990))

Lastly we present the ampltudes' evolution when the kick strength is such that it dominates the system's behaviour We saw in Chapter3 (fig 32 and table 3 1) how the elliptac fixed points of order four and the origin become hyperbohc at $\mu_{c l}=100$ This very fortunate result means that the ongin can become unstable and that orbits in its vicinity can become diffuse as they spread into the stochastic layer In the quantum case we would expect this to manfest itself as a continual spreading over all the orders of the probability amphtudes and a subsequent linear increase in energy For our third figure we present six snapshots of the evolution for $\mu_{q}=15$ The spreading out is evident and so rapid that after forty, or so, mterations we had to stop the calculations as the upper boundary on the


Fıg 62 The probability amplitudes for the same state in fig 61 except this time $\mu_{q}=05$ The interval between each plot is as before and it is worthwhule pointing out that the distributions in (d)-(f) are quite simular, ie the system is saturating naturally and the state is becoming localised
probability amplitudes had been reached The energy increase was hnear and also without saturation (within this time sample) The probability amplitudes are presented in figs $63(a)-(f)$

It can sometimes be difficult to compare successive aterations of the amplitudes


Frg 69 The amphtude spectrum for the same initial state in fig 61 but with $\mu_{q}=15$ and the time between successive plots being 4 time steps The spread over the states is apparent and boundary saturation occurs soon after (f) at around 40 tıme steps
and see obvious differences as some changes from iteration to iteration can be very subtle but nevertheless important It is clearer to examine the energy of the system as a function of the iterations to see just how substantial a change can occur a fixed time interval We perform this task in the next section with an examination of the wavefunctions' structures

## 62 Energy, Quası-Energy Levels \& Wavefunction Evolution

As with any quantum system, once you know the probabllty amplitudes and you have a complete set of eigenfunctions, then obtaining such quantities as the energy, wavefunction probabilities in space and constructing quasi-phase spaces in relatively strasghtforward In this section, the energy as a function of the time steps $1 s$ examined and the wavefunctions at each step calculated We will also show that stable energy levels exist, for moderate kıck strengths Depending on the initial state of the system, one of these levels is approached by the energy of the system We will try, within the bound specffied by the finite size of $U_{1}$ e 350 eigenfunctions, to calculate some energy level statistics as is done for most of the time-independent non-hnear quantum system examined in the journals (Berry at al (1986), Robnik et al (1986), Lewenkopf (1990) Tanner et al (1991))

For neghghbly small kicking the effect on the system's energy is very small if at all noticeable In the previous section figs $61(a)-(f)$ showed how the initial state changed almost adiabatically with little discernible change in its profile We present in fig 64 the energy corresponding to the amplutudes in fig 61 showing how little the change in energy is over the time interval The kıck strength for this case is $\mu_{q}=000005$ This case corresponds to the limit as $\mu_{q} \rightarrow 0$ and is included as a test to confirm that the system is behaving properly or to put it succintly to check that the numerics are correct

Artificial boundary effects, caused by the imposition of the finite eigenfunction set, can be seen clearly in those energy curves where $\mu_{q}$ is large enough to diffuse the intial state sufficiently so that it collides with the boundary of 350 eigenfunctions These boundary effects manfest themselves as bend overs in the energy curves, for both the off and on resonance cases, and all observed energy saturations have to be checked and thoroughly examined to exclude any such artificial effects The primary reason for this scrutiny, also the reason why we don't employ a blanket masking out of all observed saturations, is that it has been hypothesised (Chinkov et al (1981)) that energy saturations are an indication of a suppression of quantum chaos, as has been observed in the kicked rotator system


Fig 64 The energy evolution for the quantum system corresponding to the probabllyty amplitudes in fig 61 The near constant energy confirms the almost perturbative effect the kıcking is having on the system at $\mu_{q}=$ 000005
(Casatı et al (1979)), bistable systems (Ray (1990)) and simple linearly quantised systems (Scharf \& Sundaram (1991)) This is still a subject for contention with some (Berman et al 1991) argung that these systems are non-physical and that the inclusion of an extra time scale (such as in our kicked oscillator) is sufficient to counter the suppression seen in these other systems (Shepelyansly (1983) \& Adachu et al (1988)) The kucked harmonc oscillator model, on the other hand, is a realistic system and thus any saturation, and related localisation, in energy could be considered indicative of choos suppression

Unfortunately as with everything in life there is a catch, namely we need to increase the size of our matinx to eliminate the possibility of saturation Strictly speaking the size should be infinte to be absolutely positive that no saturation occurs and this is numerically impossible To date with a test matrix of $450 \times 450$ (we couldn't and wouldn't trust MathLink for orders any hugher than 450 due to its inability to handle some of the numbers Mathematica was sending it) we found no evidence of saturation In fig 65 we


Fig 65 Five energy curves for the mixed state at $\mathrm{m}=10$ The curves correspond to (from top to bottom) $\mu_{q}=17,16,15,14 \& 13$ Though not perfect for calculating the quantum duffusion coefficient, $D_{q}$, we will fit straight lines to them to estimate $D_{q}$ at each of the values The result in shown in fig 66
give a composite plot of five different kick strengths over a specified time We omit some of the energy points at higher kick strengths due to boundary saturation of the energy at these higher levels

The linear nature of these energy curves, we found the energy to scale as

$$
\begin{equation*}
E\left(\mu_{q}, n\right)=B\left(\mu_{q}\right) n \tag{array}
\end{equation*}
$$

suggest it might be possible to evaluate a quantum diffusion coefficient, $D_{q}\left(\mu_{q}\right)$ The problems associated with this exercise are, the small time interval over which the energy doesn't saturate by impinging on the boundary, the requirement to calculate a new evolution matrix for each value of the kick strength and the need to average over a set of states which is further limited by the finite size of the untary matrices However we do have a number of matrices of size $350 \times 350$ calculated and for these matrices with linear
energy increases we attempted to calculate a quantum diffusion coefficient, $D_{q}\left(\mu_{q}\right)$ where we define this diffusion coefficient to be the partial derivative of the energy with respect to tıme for a fixed kıck strength

$$
\begin{equation*}
D_{q}\left(\mu_{q_{o}}\right)=\left\{\frac{\partial E\left(\mu_{q}, t\right)}{\partial t}\right\}_{\mu_{q} \equiv \mu_{q_{o}}} \tag{array}
\end{equation*}
$$



Fig 66 The quantum diffusion coefficient, $D_{q} \equiv \frac{\partial E(t)}{\partial t}$, vs $\mu_{q}$ The increase is certainly non-lnear and found generally to increase as $A\left(\mu_{q}\right)^{b}$ where $A=02$ and $b=468$ The solid hne is the fit associated with $A \& b$ above and fits the data very well The bump is not a gltch and was found to occur at $\mu_{q}=15$

The result is presented in fig 66 We found a glitch to occur at $\mu_{q}=15$ and at finer resolution we found the neighbouring points to be well behaved (in the sense they followed the general trend of the curve) but the slope of the energy at $\mu_{q}=15$ took a definite jump This occurance at close to $\frac{\pi}{2}$ cannot be discarded and may be a kınd of enhancement simular to the Resonantly Enhanced Diffusion in Chapter 4 One of the methods this ghtch will
be examined in the future is to examine the elements of the evolution matrix at specific kick strengths, their hmists, as the matrix size tends to infinty, and therir influence on the momentum (and hence duffusion) It should be possible to do this hand in hand with the future analysis proposed for the quasi-energy levels described later

This finishes our examining of those diffuse states with apparently boundless energy increase Such states may exast the kind of hypersensitivity referred to by Schack \& Caves (1993) as these states are responsible to the extended structure we shall see in the the Wigner distributions in Chapter 7 These states are (over our limited range of scrutiny) de- localised and do not give us any information on possible stable energy levels withn our kacked system To do this we need to examune localised states

The final set of energy curves dealt with those which saturate naturally (those which don't impinge on the finte boundary dicatated by the numerics) and saturate at different levels for different intial conditions These levels are steady state solutions for the quantum mapping at a specific value of $\mu_{q}$ They are also reffered to as Quası-Energy levels because the wavefunctions are periodic wrt the kick duration Therefore at each kick the wavefunction returns to 1 ts value at the previous kick and the energy stays constant The wavefunction has become localised in terms of the kicks Israilev (1990) introduces such levels in the kicked rotator where he refers to the quasi-energies first introduced by Zeldovich (Isralev (1990)) and Ritus (Isralev (1990)) back in 1966 We use his notation here We introduce a set of quasi-energies $\epsilon$ determined by the relationships

$$
\begin{equation*}
\psi_{\epsilon}(q, t)=e^{-\iota \epsilon t / T} \phi_{\epsilon}(q, t) \tag{623}
\end{equation*}
$$

where

$$
\begin{equation*}
\phi_{\epsilon}(q, t+T)=\phi_{\epsilon}(q, t) \tag{6}
\end{equation*}
$$

In this sense the $\phi_{\epsilon}(q, t)$ are eigenfunctions of the evolution operator matrix $U$ for the specific value of $\mu_{q}$ we are dealing with The dagonalisation of the evolution matrix $U$ for any kicked system should present us with a complete set of these eigenfunctions (Berry et al (1986)) but as $U$ is an infinte square matrix we need to be sure that the decay of off diagonal terms to zero is rapid enough so that their contribution to the calculation of the eigenvalues of $U$ is neghgible $O n$ examınation of the necessary matrix data files for
the real and imaginary components at $\mu_{q}=05$ and $\beta=(\sqrt{5}+1) / 2$, whuch is the case we considered, we found that the off diagonal terms decay to more 15 orders of magnitude less than the principle diagonal terms within 170 terms of the principle diagonal If we set thus hmit to 10 orders less then the limit is 100 terms to the dagonal and reduces to 50 if the lumit is set at 5 orders below the principle dagonal term So the question is what accuracy do we require to determine the cut-off for the off-dzagonal terms ${ }^{\curvearrowright}$ This cut-off will also determine how many eigenvalues we can calculate from a fixed $350 \times 350$ matrix This is only one of several considerations necessary for the solving of the matrix, $U$ To date we have not yet attempted this beyond a brief trial using the routines in Mathematica (which crashed) yet despite this the idea does provide an interesting direction in which this research can proceed in the near future such as obtainng the energy level statistics (Berry et al (1986), Lewenkopf (1991)) However, as shown by Lewenkopf (1991), these statistics are not conclusive evidence of chaos We continue now with the quasi-energy levels which we obtained from the evolution matrices

When investigating the existence of steady state levels (quasi-energy levels) we found that not all initial states give unique steady state levels and so we have degeneracy present We found these steady state levels to exist in the system at a kıck strength of $\mu_{q}=05$ for an off resonance case $(\beta=(\sqrt{5}+1) / 2)$ From the plot of the results, in fig 67 , one can see quite clearly the three levels attanned by the four starting states The lowest state (that of a gaussian distribution of probability amplitudes centered on the lowset state $m=0$ ) evolved to the same steady state level as that for a state with gaussian distribution centred on $m=11$ This was included to illustrate the degeneracy present and also to highlight that this steady state level is probably the lowest level in the system We can see that the levels do not appear to be equaliy spaced but, with the matrix size avaulable to us, we do not have enough levels present to construct a relationshp between them With the future diagonalisation of the evolution matrix, $U$, we should be capable of relating levels (or groups of levels) together

Our final discussion in this section is centered about the structure of the wavefunctions themselves and how the probability in $Q$ is related to the diffuse or localsed nature of the system at the values of the kıck strength being considered


Frg 67 The three lowest levels obtanned for $\mu_{q}=05$ and $(\beta=(\sqrt{5}+1) / 2)$ The levelling off is obvious and indicative of a stable solution of the evolution matrix These levels are the quasi-energy levels discussed in the text

Any wavefunction which satısfies Schrodinger's wave equation (eq (5 13 )) for the undriven system, whose Hamultonian is given by eq (512), can be expressed as a linear combination of the eigenfunctions of that system, $c f$ eq (514) For the kicked system (eq (5 1 1) ) the wavefunctions are related to the undriven system's eigenfunctions with time dependent amphtudes given by the relationship in eq (5130) So once we have the amplitudes we can calculate what the wavefunction would look like at each time step The importance of examining the wavefunctions is principally to see how the evolution of system affects the exastence probabilities of the system at different $Q$ values As the elgenfunctions themselves have an $\exp \left(-\frac{1}{2} Q^{2}\right)$ term present they will tend to zero as $Q$ tends to infinity The higher the order of the elgenfunction, the further out along the $\pm \boldsymbol{Q}$ axes the wavfunction will exast for Thus for a finite set of amphtudes (as we have here with our $350 \times 350$ matrix) the wavefunctions will tend to zero as $Q$ increases beyond a critical value dependent on the highest eigenfunction order present in the evolution matrix with non-zero probability ampltude Therefore we expect that, for those states which diffuse

$F_{2 g} 68$ The wavefunction, $\psi(Q)$, for the system when the kuck strength, $\mu_{q}$, is so small as to be negligrble The system is on resonance and $\beta=\frac{\pi}{2}$ The time interval between plots is 300 iterations The wavefunction is symmetric about the origin ( $Q=0$ ) so only the positive half is shown
rapidly up through the orders of the elgenfunctions, the wavefunctions would be the most spread out (diffuse) in a given time scale Those which saturate naturally at a given energy above their intial state would be less spread out whereas those whose energy stays approximately constant and equal to their mitial state would have little or no spreading at all The degree of spreading out over the eigenfunctions orders would also determine the degree of diffusivity and lack of structure present in the wavefunction To llustrate these three cases we present the wavefunctions for the three cases considered so far for the


F2g 69 The wavefunction, $\psi(Q)$, for when the kıck strength, $\mu_{q}=05$ and $\beta=$ $\frac{(\sqrt{5}+1) \pi}{2}$ The time interval between plots is 300 iterations The system has saturated naturally and the wavefunction can be seen to be time independent in the final 3 plots supporting this natural saturation, $\mathrm{c} f$ fig 67
energy evolution and the probability amplitude evolution
The initial state we choose for the following set of three figures is the mixed state consisting of a gaussian distribution of probability amplitudes centred on the $m=10$ elgenstate The figures 6 8-6 10 show $\left|\psi\left(Q, \mu_{q}, t\right)\right|^{2}$ vs $Q$ at specified times of the system's evolution The wavefunction is defined according to eq 514 and the dependence of the probabilty amphtudes on tıme and $\mu_{q}$ is given by the evolution equation, eq ( 5130 )

The figures show just the positive half of the $Q$ axas but due to the property that $H_{-n}(Q)$ $=(-1)^{n} H_{n}(Q)$, and as we're plotting $\left|\psi\left(Q, \mu_{q}, t\right)\right|^{2}$, the resultant figures are symmetric about the origin because $\left(H_{-n}(Q)\right)^{2}=\left(H_{n}(Q)\right)^{2}$


Fig 610 The wavefunction, $\psi(Q)$, for when the kuck strength, $\mu_{q}=15$ and $\beta=\frac{\pi}{2}$ The parameter values correspond to the unbounded growth we discussed and presented in fig 65 The time between each plot is 6 time steps of the mapping Even over such a short penod of time the spread over $Q$ and the increasing lack of structure indicates a signoficant diffusion for thus state

As before we begin with the case where the kick strength 16 so small as to be a perturbation of the undriven system We use this case as a test to ensure that in the limit
$\mu_{q} \rightarrow 0$ the system reverts to the undriven case This is what figs $61 \& 64$ llustrate for the probability amphtudes and energy respectively The corresponding figure for the wavefunction itself is shown is fig 68 The change is negligible, if at all, and the system is behaving like the undriven oscillator as we would expect This limit is important as it ensures a continuty from the undriven to the driven systems and, as the undriven oscillator has been well studed, it serves as a test to ebsure the numencs are functioning properly

The plots presented in fig 69 are those for the wavefuntion evolution when the kuck strength 1505 and the system is off resonance, 2 e $\beta=\frac{(\sqrt{5}+1) \pi}{2}$ This corresponds to the system in fig 67 where we illustrated the natural saturation of the system and the existence of quasi-energy levels The intial state chosen for the wavefunctions in fig 69 corresponds to the lowest energy level in fig 67 The temporal periodicity of the wavefunction, $\psi(Q)$, which ensures its steady state behaviour at each time step makes this wavefunction a eigenfunction of the driven system The energy corresponding to this eigenfunction would constitute one of the qausi-energy levels we discussed previously One of the more interesting observations that can be made of the probability distributions in fig 69 is the structure which exists in the wavefunction in the last three plots The peaks near the origin in the second plot ( 300 time iterations after the start) remann right through the time evolution up to 2000 iterations These, plus the smaller and less prominent peaks, indicate that certain eigenfunction are preferred over others and these preferential eigenfunctions are those with the greatest overlap with the peaks seen Such stable structures in the wavefunction is important as it is proof that the wavefunction is not going to spread out over the complete basis of eigenfunctions but will remain essentially restricted to a subset of the basis Thus the evidence of stable structure in the wavefunction can be used as a method for determining the diffusvveness of an initial state for a given set of parameters

The unbounded like spread in the probabllty amplitudes in fig 63 and the corresponding unsaturating energy increases in fig 65 suggest that the wavefunction for such a class of state would diffuse from the initial bounded (in $Q$ ) state and tend to a state which was spread out near unformly over the whole $Q$ axis This tendency to spread out over $Q$ would necessitate the inclusion of higher (and higher) order eigenfunctions, $u_{n}(Q)$, of the undriven oscillator as these elgenfunctions are themselves bounded $m$ a finte section of the
$Q$ axas for a finte order This to spread over the full axas (or even tend to such a spreading) the initial state would have to diffuse over all the eigenfunctions of the undriven system which would subsequently lead to an unbounded increase in the waveunction's energy We present, in fig 610 , the closest we can come to such an unbounded diffuse state Our size of matrix ( $350 \times 350$ ) limits our analysis and the number of iterations we can perform on the state before the state reaches the limit mposed by our finte matrix and our results become meaningless Therefore the time duration of our analysis of such diffuse states is, on average, about 40 iterations (depending on the kick strength and whether we're on or off resonance) The wavefunction evolution shown in fig 610 is of such a diffuse state with an interval of 6 time steps between each plot, it is trivial for one to work out the total evolution illustrated encompasses 30 interations The kick strength, $\mu_{q}$, is set equal to 15 , about mid-way along the curves shown in fig 65 , and the system is on resonance with $\beta=\frac{\pi}{2}$ The plots show how, after a short tume, the mitial well-structured state becomes more diffuse and spreads out over $Q$ with increasing time and how the wavefunction itself is losing any structure and tending to a flat distribution across the $Q$ axis The origin has a still somewhat high probability but even this is decreasing with increasing time and we would expect, on this evidence and trend, for the probability, $|\psi(Q)|^{2}$, to tend to become independent of $Q(\imath e$ a flat distribution)

## CHAPTER 7 <br> THE WIGNER DISTRIBUTION OF THE QUANTUM MAPPING

In this chapter we derive a form for the Wigner distribution (Wigner (1932), Hillery et al (1984)) and use it to construct a quasi-phase space for the quantum mapping By using thus Wigner distribution we hope to compare like with hke (the observed structures in the quantum quasi-phase space with those in the classical phase space) The principle aum here is to develop a correspondence between the quantum and the classical structures with the intention of using thus correspondence to argue the exastence of quantum chaos

## 71 The Wigner Distribution

The effect of classical orbits scarring the quantum wavefunctions can be best examined using those distributions which generate a quasi phase space for the system When we use the term scar, we refer to the correspondence in the quantum eigenfunctions and quası-phase space of a classical periodic orbıt (Heller (1989)) By quası we refer to the fact that trajectories as we understand them $m$ the classical regime have no meaning in the corresponding quantum regime Therefore we use a specific form of distribution whose task is to represent the quantum wavefunction in a momentum and position space ( $P \& Q$ in our case) The most widely used of this type of distribution is that attributed to Wigner (1932) He proposed a distibution, denoted $P_{W}(p, q)$ whose form is given by

$$
\begin{equation*}
P_{W}(\boldsymbol{q}, \boldsymbol{p})=\frac{1}{\pi \hbar} \int_{-\infty}^{\infty} d y<\boldsymbol{q}-y|\hat{\rho}| q+y>\exp (2 \iota p y / \hbar) \tag{array}
\end{equation*}
$$

which for a pure state, $\psi$, becomes (Hullery et al (1984))

$$
\begin{equation*}
P_{W}(q, p)=\frac{1}{\pi \hbar} \int_{-\infty}^{\infty} d y \psi^{*}(q+y) \psi(q-y) \exp (2 \iota p y / \hbar) \tag{712}
\end{equation*}
$$

To proceed with this denvation we follow the steps that Hillery et al use in their calculation of the Wigner distribution for a pure eigenstate of the harmome oscillator (Hallery et al (1984)) The eigenstates of the harmonic oscullator are given by

$$
\begin{equation*}
\phi_{n}(q)=\left(\frac{\alpha^{2}}{\pi}\right)^{1 / 4}\left(\frac{1}{2^{n} n^{1}}\right)^{1 / 2} \exp \left(-\alpha^{2} q^{2} / 2\right) H_{n}(\alpha q) \tag{array}
\end{equation*}
$$

where $\alpha=\sqrt{m \omega_{o} / \hbar}$ Therefore any tume-dependent wavefunction can be expressed as a linear combination of the eigenstates

$$
\begin{equation*}
\psi(q, t)=\sum_{n=0}^{\infty} a_{n}(t) \phi_{n}(q) \tag{714}
\end{equation*}
$$

The Wigner distribution of this mixed state is thus

$$
\begin{equation*}
P_{W}(q, p)=\int_{-\infty}^{\infty} d y \exp (\iota 2 p y / \hbar) \psi^{*}(q+y) \psi(q-y) \tag{715}
\end{equation*}
$$

Expanding $\psi(q, t)$ (writing $a_{n}^{*}(t)$ as just $a_{n}^{*}$ and sımularly for $\left.a_{m}(t)\right)$ we get

$$
\begin{gather*}
P_{W}(q, p)=\sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\alpha}{\sqrt{2^{n} 2^{m} n^{\prime} m^{\prime} \pi}} a_{n}^{*} a_{m} \\
\int_{-\infty}^{\infty} d y \exp (\iota 2 p y / \hbar) \exp \left(-\alpha^{2}\left(q^{2}+y^{2}\right)\right) H_{n}(\alpha(q+y)) H_{m}(\alpha(q-y)) \tag{716}
\end{gather*}
$$

where we have replaced the $\exp \left(-\alpha^{2}(q-y)^{2} / 2\right) * \exp \left(-\alpha^{2}(q+y)^{2} / 2\right)$ term arising from the $\exp \left(-\alpha^{2} q^{2} / 2\right)$ part of the elgenfunctions, $U_{n}(\alpha q)$, by the more compact $\exp \left(-\alpha^{2}\left(q^{2}+y^{2}\right)\right)$ Removing all terms not dependent on $y$ from the integral, we obtain

$$
\begin{gather*}
P_{W}(q, p)=\sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\alpha \exp \left(-\alpha^{2} q^{2}\right)}{\sqrt{2^{n} 2^{m} n^{\prime} m^{\prime} \pi}} a_{n}^{*} a_{m} \\
\int_{-\infty}^{\infty} d y \exp (\iota 2 p y / \hbar) \exp \left(-\alpha^{2} y^{2}\right) H_{n}(\alpha(q+y)) H_{m}(\alpha(q-y)) \tag{717}
\end{gather*}
$$

To be consistent with the other derivations we have for the Harmonic Oscillator problem we now replace $p \& q$ with their dimensionless cousins $P \& Q$ where

$$
\begin{gathered}
Q=\alpha q \\
P=p /(\alpha \hbar)
\end{gathered}
$$

and we replace $y$ with a dımensionless form $Y=\alpha y$ so that $d Y / \alpha=d y$ Our expression for $P_{W}(q, p)$ changes to $P_{W}(Q, P)$ given by

$$
P_{W}(Q, P)=\sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\exp \left(-Q^{2}\right)}{\sqrt{2^{2} 2^{m} n^{\prime} m^{\prime} \pi}} a_{n}^{*} a_{m}
$$

$$
\begin{equation*}
\int_{-\infty}^{\infty} d Y \exp (\iota 2 P Y) \exp \left(-Y^{2}\right) H_{n}(Q+Y) H_{m}(Q-Y) \tag{718}
\end{equation*}
$$

Noting that $\exp \left(\iota 2 P Y-Y^{2}\right)$ can be written as $\exp \left(-(Y-\iota P)^{2}\right) \exp \left(-P^{2}\right)$ we find that the expression for $P_{W}(Q, P)$ takes the form

$$
\begin{gather*}
P_{W}(Q, P)=\sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\exp \left(-\left(Q^{2}+P^{2}\right)\right)}{\sqrt{2^{n} 2^{m} n^{\prime} m^{\prime} \pi}} a_{n}^{*} a_{m} \\
\int_{-\infty}^{\infty} d Y \exp \left(-(Y-\iota P)^{2}\right) H_{n}(Q+Y) H_{m}(Q-Y) \tag{array}
\end{gather*}
$$

We now define the new variable $z=Y-\iota P$ so that the $d z=d Y$ The arguments inside the Hermite polynomals become $Q+z+\iota P$ instead of $Q+Y$ and $Q-z-\iota P$ instead of $Q-Y$ The argument for the $H_{m}$ is thus depending negatively on the variable $z$ so we use the identity that $H_{m}(-z)=(-1)^{m} H_{m}(z)$ With these two changes, noted $P_{W}(P, Q)$ becomes

$$
\begin{align*}
& P_{W}(Q, P)=\sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\exp \left(-\left(Q^{2}+P^{2}\right)\right)}{\sqrt{2^{n} 2^{m} n^{\prime} m^{\prime} \pi}} a_{n}^{*} a_{m}(-1)^{m} \\
& \int_{-\infty}^{\infty} d z \exp \left(-z^{2}\right) H_{n}(z+\iota P+Q) H_{m}(z+\iota P-Q) \tag{7110}
\end{align*}
$$

We now invoke that well used (by us anyhow) identity in Gradshteyn and Ryzhuk (1965) (eq 7377 pg 838 ) which gives an analytical solution for the integral above with a certain restriction, namely it exasts for $n \leq m$ However (as will be shown later) the order in the integral can be swapped keeping the arguments the same (the order they appear in the integral changes but the arguments of $H_{m}$ and $H_{n}$ are the same as above) so we can obtain a solution for the case $m \leq n$ Using the identity from Gradshteyn and Ryzhik (1965) we obtain the following solutions for the case $n \leq m$

$$
\begin{gather*}
P_{W}(Q, P)=\sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\exp \left(-\left(Q^{2}+P^{2}\right)\right)}{\sqrt{2^{n} 2^{m} n^{\prime} m^{\prime} \pi}} a_{n}^{*} a_{m} \\
(-1)^{m} 2^{m} \sqrt{\pi} n^{\prime}(\iota P-Q)^{m-n} L_{n}^{m-n}\left(2\left(Q^{2}+P^{2}\right)\right) \tag{7111}
\end{gather*}
$$

This simpifies to

$$
P_{W}(Q, P)=\exp \left(-\left(Q^{2}+P^{2}\right)\right)
$$

$$
\begin{equation*}
\sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \sqrt{\frac{2^{m} n^{\prime}}{2^{n} m^{\prime}}} a_{n}^{*} a_{m}(-1)^{m}(\iota P-Q)^{m-n} L_{n}^{m-n}\left(2\left(Q^{2}+P^{2}\right)\right) \tag{7112}
\end{equation*}
$$

whach of course has the restriction that $n \leq \mathrm{m}$ For the case $m \leq n$ we swap the order in the integral for $H_{n}$ and $H_{m}, 1$ e

$$
\begin{gathered}
\int_{-\infty}^{\infty} d z \exp \left(-z^{2}\right) H_{n}(z+\iota P+Q) H_{m}(z+\iota P-Q) \equiv \\
\int_{-\infty}^{\infty} d z \exp \left(-z^{2}\right) H_{m}(z+\iota P-Q) H_{n}(z+\iota P+Q)
\end{gathered}
$$

in our case the above holds but in general the swapping of functions within an integral like the above depends on the individual properties of the functions (or operators) being considered In this case the expression takes on the slightly different form

$$
\begin{gather*}
P_{W}(Q, P)=\exp \left(-\left(Q^{2}+P^{2}\right)\right) \\
\sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \sqrt{\frac{2^{n} m^{\prime}}{2^{m} n^{\prime}}} a_{n}^{*} a_{m}(-1)^{m}(\iota P+Q)^{n-m} L_{m}^{n-m}\left(2\left(Q^{2}+P^{2}\right)\right) \tag{7113}
\end{gather*}
$$

which can be changed to a more comparable form for the case of $n \leq m$ by noting that the $(-1)^{m}(\iota P+Q)^{n-m}$ can be replaced by $(-1)^{n}(-\iota P-Q)^{n-m}$ Grouping all real components together and separating out and grouping those elements containing both real and complex parts we get for $P_{W}(Q, P)$

$$
\begin{array}{cc}
P_{W}(Q, P)=\exp \left(-\left(Q^{2}+P^{2}\right)\right) \\
\sum_{m=0}^{\infty} \sum_{n=0}^{m} \sqrt{\frac{2^{m} n^{1}}{2^{n} m^{\prime}}}(-1)^{m} L_{n}^{m-n}\left(2\left(Q^{2}+P^{2}\right)\right)\left\{a_{n}^{*} a_{m}(-Q+\iota P)^{m-n}\right\} & n \leq m \\
P_{W}(Q, P)=\exp \left(-\left(Q^{2}+P^{2}\right)\right) & \\
\sum_{n=0}^{\infty} \sum_{m=0}^{n} \sqrt{\frac{2^{n} m^{\prime}}{2^{m} n^{\prime}}}(-1)^{n} L_{m}^{n-m}\left(2\left(Q^{2}+P^{2}\right)\right)\left\{a_{n}^{*} a_{m}(-Q-\iota P)^{n-m}\right\} & m \leq n \tag{array}
\end{array}
$$

It is quite straughtforward to see that when $m=n$ the two expressions are the same All terms outside the $\}$ brackets on the RHS in both cases are equal in value when $m$ in the first equals $n$ in the second If we think of each element outside the $\}$ (and exclusive of the summations) as a matrix element $A_{n m}$ of some matrix $A, 1$ e

$$
\begin{equation*}
A_{n+1, m+1}(Q, P)=\exp \left(-\left(Q^{2}+P^{2}\right)\right) \sqrt{\frac{2^{m} n!}{2^{n} m^{!}}}(-1)^{m} L_{n}^{m-n}\left(2\left(Q^{2}+P^{2}\right)\right) \tag{7116}
\end{equation*}
$$

then what we're stating here is that $A_{1 j}=A_{j 2}$ or that $A^{T}=A$ where $T$ signifies transpose What follows from this is that if the pars of elements $A_{y_{2}}$ and $A_{\imath j}$ are summed together first, and then all these pairs added for the complete sum, then the terms in the $\}$ brackets (for each of these parr additions) are complex conjugates of each other and the complex parts subsequently cancel leaving a real valued distribution which is one of the more important properties of Wigner distributions A proof of how these terms give rise to complex conjugates can be found in Appendix $C$ as well as how the $P$ and $Q$ values in each of the four quadrants in the $2 D$ position-momentum plane are related

It is also worth adding here that as the matrix $A$ is equal to its transpose then we only have to calculate all of the elements $A_{i j}$ on one side of the principle diagonal Thus the equation for the Wigner distribution is

$$
\begin{gather*}
P_{W}(Q, P)= \\
2 \sum_{m=0}^{\infty} \sum_{n=m}^{\infty} A_{n+1, m+1}(Q, P) *(a(Q, P) * \gamma(Q, P)-b(Q, P) * \delta(Q, P)) \tag{array}
\end{gather*}
$$

where we have used eq ( $C 3$ ) obtained in Appendux $C$ in substituting for the complex terms This is the expression we numerically evaluate to obtain the Wigner distribution

Most of the literature today quote using the Husimi distribution as the source of their quasi-phase space portrats (Radons \& Prange (1990), Balazs (1990), Scharf \& Sundaram (1991), Kuś et al (1991)) The Husim distribution (Husim (1940)) is a special case of the scheme proposed by Cohen (Hillery et al (1984)) for generating distribution functions Cohen proposed the use of the expression

$$
\begin{equation*}
P_{g}(q, p)=\int_{-\infty}^{\infty} d q^{\prime} \int_{-\infty}^{\infty} d p^{\prime} \tilde{g}\left(q-q^{\prime}, p-p^{\prime}\right) P_{W}\left(q^{\prime}, p^{\prime}\right) \tag{7118}
\end{equation*}
$$

for calculating distributions other than the Wigner distribution One can see from the above expression that the distribution function $P_{W}$ is simply smeared by the function $\tilde{g}$ (Hillery et al (1984)) The Husimi distribution is such that this smearing function $g$ has a form simular to the coherent state representation used by investigators analysing systems with E-M radiation playing a major role (lasers etc) The explecit form of $\tilde{g}$ is

$$
\begin{equation*}
\tilde{g}(q, p, \alpha)=\frac{1}{\pi \hbar} \exp \left(-q^{2} / \alpha\right) \exp \left(-\alpha p^{2} / \hbar^{2}\right) \tag{71.19}
\end{equation*}
$$

which leads to a distribution $P_{H}$ ( $H$ for Husim) which is positive for all $p$ and $q$ The corresponding smeanng function, $\tilde{g}$, for $P$ and $Q$ is

$$
\begin{equation*}
\tilde{g}(Q, P)=\frac{1}{\pi \hbar} \exp \left(-Q^{2}\right) \exp \left(-P^{2}\right) \tag{7120}
\end{equation*}
$$

so that the Husim distribution becomes

$$
\begin{equation*}
P_{H}(Q, P)=\frac{1}{\pi} \int_{-\infty}^{\infty} d Q^{\prime} \int_{-\infty}^{\infty} d P^{\prime} \exp \left(-\left(Q-Q^{\prime}\right)^{2}\right) \exp \left(-\left(P-P^{\prime}\right)^{2}\right) P_{W}\left(Q^{\prime}, P^{\prime}\right) \tag{7121}
\end{equation*}
$$

This is all very nice on paper but implementing the above is another matter Our exact analytical expression for $P_{W}(Q, P)$ aliows us to choose any point ( $Q_{o}, P_{o}$ ) on the phase space and calculate its $P_{W}$ at that point For the Husimi distribution, however, we have a problem Following a search of as many textbooks as we could find on the subject, we couldn't come up with an analytical form for eq ( 7121 ) and its integrals contanned within Therefore the calculation of $P_{W}$ rests solely on approximating the double integral over the phase space by a crude summation over the $P_{W}$ for the points we considered This is acceptable if the number of calculated $P_{W}$ values is quite large but our calculated $P_{W}$ consisted of a $64 \times 64$ gnd over a square of side $25 Q \& P$ units centered on $(0,0)$ So we, with the time available and the computing power at our disposal, decided to use the Wigner distribution as our guide to the quantum phase space

Now that we have our expression for $P_{W}$ we will now present some of the results obtaned by numerically evaluating the expression for $P_{W}$ in $e q$ ( 7117 ) We would hope for a correspondance to exist between the regions of high probability m the quantum space and those of hagh densities in the classical space This scarring of the quantum quassphase space by classical orbits is considered by many to be the signature of a quantum correspondance to classical chaos

## 72 The Quası-Phase Space

The almost overwhelming diversity of chorce for the initial conditions for the quantum mapping presented us with many dulemmas up to now but this was no preparation for
the problems we faced here in evaluating the wigner distributions for the quantum phase space Each $64 \times 64$ gnd of distribution values for one tıme step for a specific set of parameters took 20 hours computing time ${ }^{\prime}$ To examine even crudely the system's behaviour at various parameter values, at specific times and for vanous initial states we had to make some decisions regarding the extent of our examination at various parameter values, times etc In the following sets of figures we present seven pairs of phase spaces with each parr consisting of an zmage of the phase space with the corresponding contour below it The tımes, parameter values and intial states are given when appropriate and the reasons for their choices also explauned

Our first pair of figures show the wigner distribution, $P_{W}(Q, P)$ for the chosen initial pure state By pure we refer to a state which is made up of a single eigenfunction of the undriven system The state chosen is the $10^{\text {th }}$ eigenfunction ( $\mathrm{m}=10$ ) of the harmonc oscillator, 1 e

$$
\begin{equation*}
(\psi(Q))_{2}=u_{10}(Q)=\frac{\iota^{10}}{\sqrt{2^{10} \sqrt{\pi}\left(10^{\prime}\right)}} * H_{10}(Q) \exp \left(-\frac{1}{2} Q^{2}\right) \tag{721}
\end{equation*}
$$

where the subscript $\imath$ for the $\psi$ indicates that its corresponds to the initial state of the system The image shown in fig 71 shows the circularly symmetric property of a pure elgenstate of the undriven system It is interesting to compare this figure (and its proceeding companion contour plot in fig (72)) with the phase space for the corresponding classical system when there is an absence of the driving potential, $c f$ fig (21) The circular orbits of the classical system are strikingly simular to the pure eigenstates of the undriven system as we would expect The Husimı distribution for a $400 \times 400$ grid over the phase space of the Wigner distribution, which can be obtaned because a pure state Wigner distribution requires many times less calculations then a muxed state distribution, shows a single smeared orbit circling a peak at the origin This is even more hike a classical orbit which is normally a single circular orbit centered on the ongin whose radus is the measure of its energy ( $c f$ chapter 2 section 2 and chapter 3 section 2)

So as expected we have a direct correspondence between the quantum and classical system's for the undriven harmonic oscillator The question remains as to the extent of the correspondence between the two when the kucking is present and beyond a perturbation

Frg 71 Image of $P_{W}(Q, P)$ for the intial system state corresponding to the $10^{\text {th }}$ eigenfunction of the undriven system One can clearly see the circular symmetry for this pure state and how thas symmetry will be seen to contrast with the distributions for later times


Fıg 72 The corresponding contour plot of $P_{W}(Q, P)$ for the system as described for the amage plot in fig (71) above

We make the note here that the scale used on the $Q \& P$ axes of all the following figures (both images and contours) is the same so compansons can be made Furthermore each Wigner distribution is calculated at a specific time after the initial state The distribution is time dependent due to the tıme dependence of the probability amplitudes, $a_{m}(t)$, in constructing the wavefunctions, $\psi(Q, t)$, refer to eq (514) for the explicit definition Therefore the distribution will change in time, whether by a substantial amount or by a negleghble amount depends on the system parameters and initial state Some change drastically such as the case for the on-resonance scenario with an intial pure state (compare fig (71) to fig (73)) Others change very little with the intial state's quasi-phase space being quite stable even after sigmficant time penods (compare fig (711) with fig (713)) So we can expect the comparison of quantum to classical to be difficult but possible once the time dependence of these distributions is considered

The image and accompanying contour plot in figs $73 \& 74$ illustrate how the phase space has evolved from the intial state in figs $71 \& 72$ after a tıme lapse of 38 iterations of the quantum mapping for $\mu_{q}=20$ In the natural time scale of the harmonic oscillator, $\tau_{o}$ where $\tau_{o}=\frac{2 \pi}{\omega_{o}}$, this would correspond to 95 full period oscillations of the undriven oscillator This is so because for this resonance case the value of $\beta=\frac{\pi}{2}$ which by eq (3 111 ) makes the period between the kucks, $\tau$, to be one-quarter the natural time period of the undriven system, $\tau_{o}$ The spread over first 110 eigenfunctions, (out of 350 for the matrix) used for calculating the Wigner distribution, from the initial single pure eigenstate is represented by the spreading out over the phase space and the obvious fourfold symmetry from the figures The dark regions represent hugher probability areas while the loght represent less probable zones There are four obvious islands surrounding the origin at 90 degree intervals which can be related back to the mvanant islands of the classical phase space The regions joining such islands are probably related to the classical stochastic layer and exhibit a striking hne type structure within them This could be a manifestation of the graphics package (ID L) joining regions of hke probablity as the number of calculated points for each of the figures presented in this chapter (except figs $71 \& 72$ ) is 6400 (an $80 \times 80$ gnd) Whatever the answer is the symmetry is authentic and not a product of I D L 's management of the data


Fıg 79 Image of $P_{W}(Q, P)$ for system 38 time steps after the mintal state shown in figs (71) \& (72) The system is on resonance with $\beta=\frac{\pi}{2}$ and $\mu_{q}=20$ The spreading out and 4-fold structure scarring the initial state is very evident

$F_{2 g} 7 \&$ The corresponding contour plot of $P_{W}(Q, P)$ for the system as described for the image plot in fig (73) above

There is one obvous difference between the Wigner distribution llustrated in figs $73 \& 74$ and the resonant classical phase spaces in chapters $2 \& 3$ and that is the 45 degree orientation of the two wrt each other One possible reason why the could be is that the initial choice of $a$ and $a^{\dagger}$ in chapter 2 relate the original basis ( $P \& Q$ ) to a new basis ( $a \& a^{\dagger}$ ) which is onientated 45 degrees to the onginal This could be just conncidence or, maybe, a factor in the onentation dufferences

The energy growth for the system with $\mu_{q}=20$ and on resonance ( $\beta=\frac{\pi}{2}$ ) is of a linear type increase simular to those shown in fig 65 Within a short time (around 40 interations) the spread of the initial state is such that the boundary of 350 eigenfunctions is hit and the energy curve saturates and bends over (refer to Chapter 6 sections $1 \& 2$ ) The rate of spreading can be related to the rate of increase in the energy curve with time (our quantum diffusion coefficient) and turns out to be hagher than the rates plotted in fig 6 If the matrix was infinty large and we could iterate the mapping for an infintely long time then we would expect the spread (from the trends given for the smaller matrix) to continue over all the eigenfunctions and the quasi-phase space for thus case would extend over the whole $2 D Q, P$ plane

Another method of checking the dynamics of the phase space, at specific points, would be to construct an asymmetric wavefunction in the $Q, P$ plane and iterate it forward in time and examine its evolution By asymmetric, we mean a wavefunction which is not symmetric about the phase space onign lake the harmonic oscillator's eigenfunctions but which is located at some point ( $Q, P$ ) in the phase space and occupies some space about this point Unfortunately, to accomphsh this, we would need to know the probability amplitudes for this wavefunction and a fourter type expansion of the wavefunction in terms of a basis of the harmonic oscillator's eigenfunctions is requred and to date we have not been able to obtain an analytical expression for the probability amplitudes of such a asymmetnc wavefunction, if any exist This is one shortcoming of our technique but one which is outweighed by the importance and necessity of relating the probability amplitudes from one iteration to the next

The next parr of figures (figs $73 \& 74$ ) show the system for an off-resonance kıcking The kick strength, $\mu_{q}$, is kept constant at 20 and the number of iterations of the


Frg 75 Image of $P_{W}(Q, P)$ for system 38 time steps after the initial state shown in figs (71) \& (72) The system is off resonance this tıme with $\beta=$ $\frac{(\sqrt{5}+1) \pi}{2}$ and $\mu_{q}=20$ It is clear that the dominant symmetry is circular as we would expect ( $c f$ fig 24 )

$F_{2 g} 76$ The corresponding contour plot of $P_{W}(Q, P)$ for the system as described for the umage plot in fig (75) above
mapping is also the same as that for the previous pair of diagrams This is to faciltate comparisons between the off and on resonance cases For this parr the parameter $\beta$ is set equal to $\frac{(\sqrt{5}+1) \pi}{2}$ which will exclude any commensurate resonance between $\tau$, the kick period, and $\tau_{o}$, the natural time period of the oscillator The differences between the Wigner distribution for this case and the distribution for the previous case in quite starthng The symmetry of the distribution is definitely circular (except at the very ongin which has a dumb-bell type shape) which makes it very like the off resonance classical phase space portrait in fig 24 The spread over the $2 D$ plane is a lot less and an energy plot for this set of parameters would reveal it to saturate naturally in a manner simular to the cases shown in fig 67 This saturation occurs quite rapidly and the final time independent distribution is identical to that in the figure

From the last two pairs of figures the following conclusions can be made concerning the evolution of the quantum system with a pure initial state Firstly, the resonance case shows conclusively that the spreading over the eigenfunctions is significantly greater than the off resonance case The result is of this is that, secondly, the energy increase is correspondingly greater giving rise to a high diffusion coefficient for the resonance case (The off resonance case has a zero diffusion coefficient when it reaches saturation) Thirdly, the general symmetries of the quantum distributions correspond to the general forms of the classical phase space for both cases We conclude from this that correspondence between the quantum mapping and the classical one is very good and, further, that the observed energy increase on resonance, with the limited matrix size, points to possible unbounded growth in energy as we would expect in a classical phase space ensemble which includes some orbits in the stochastic layer

The success with the pure initial state prompted a further analysis with a muxed intial state The muxed state chosen is made up of a gaussian distribution of pure eigenstates and includes both even and odd parities As already mentioned previously in chapter 5 there is no interaction between odd panty states and even parity states Whle we expect this we have, up to now, only investigated an even pure eigenstate as our initial state so it could be argued that, without evidence to the contrary, the results for the Wigner distributions to date could be characteristic of the even states This is the reason why a

## hom

Frg 77 Image of $P_{W}(Q, P)$ for the initial state corresponding to the a gaussian distribution of the probability ampltudes centered on the $10^{\text {th }}$ eigenfunction of the undriven system Note the dufferences between thes and fig (71)


Fig 78 The corresponding contour plot of $P_{W}(Q, P)$ for the system as described for the image plot in fig (77) above
mixed state is now considered to allay these fears The evolution of thus mixed state in tıme can be seen from the results presented in chapter 6 for the probability amplitudes, the energy evolution and the wavefunction evolution The initial state is identical to that chosen for the mixed state used in the chapter 6, a gaussian distribution of probability amplitudes centered on the $10^{\text {th }}$ eigenfunction The distribution of this mitial state is lllustrated in figs $77 \& 78$ The differences between it and the pure state in figs $71 \&$ 72 are immediately apparent There is no obvious circular symmetry, except for the outer rim caused by the finte set of eigenfunctions, and the inner pattern is almost flower like in its form The only observable symmetry from these figures is 4 -fold but the fine detall of the distribution is absent due to the restrictions placed on the size the grid of calculated points in this phase space can take In the analysis presented here all but the first parr of figures have a $80 \times 80$ grid of calculated distribution points spread over the same size section of phase space centered on the origin As a result, some of the smaller structures are absent because the grid's resolution is too course to pick them out

The result of iterating forward $m$ time this muxed intial state when the kick strength, $\mu_{q}$, is 15 and the parameter $\beta$ is $\frac{\pi}{2}$ can be seen in figs $79 \& 710$ The time interval from initial state to the state shown is 20 iterations of the mapping (equivalent to 5 full periods of oscllation of the undriven oscillator as $\beta$, with its present value, gives a 14 ratio between $\tau_{o}$ and $\tau$ The simularity to the on resonance case for a pure even initial state is obvious with the four zslands being more prominent than for the previous case in fig 73 Furthermore, as we shall haghlight later in fig 713 , there is evidence in fig 79 that four more slands could be present each situated on the corners of the sides the prominent islands make up Only half of these new islands exust but their presence would point to a possible continued (or extended) 4-fold structure in this quantum distribution The contour plot, in fig 710 , too shows a half island type structure but it is not as clear as the image plot in fig 79 Thus we have some evidence to suggest a quantum periodic tiled structure existing in the Wigner distribution of the system of the order of the resonance we're considering

The absolute test would be to construct a very large evolution matrix (say $5000 \times$ 5000 ) and check to see how far with this matrix this extended structure continues into the


F2g 79 Image of $P_{W}(Q, P)$ for system 20 time steps after the intial state shown in figs (75) \& (76) The system is on resonance with $\beta=\frac{\pi}{2}$ and $\mu_{q}=15$ Some of the structure seen in figs $73 \& 74$ is not as clear but the dominant 4 -fold symmetry is stll very evident


Fig 710 The corresponding contour plot of $P_{W}(Q, P)$ for the system as described for the image plot in fig (79) above
quantum quast phase space At present thes is beyond the computing and software hmints avalable to us but with more powerful computers and a more efficient lunk to Mathematica to generate the Laguerre polynomals we should be capable of exploring the phase space more thoroughly and probing this apparent 4-fold structure

Serendipity in science is quite common and indeed we found something by accident which highlights the extended structure we have just mentioned The onginal distributions we obtained had a beautiful tiled structure up to the limit of the eigenfunctions extent in the phase space This structure was taken by us to be the proof we needed of a quantum like tiled phase space untll we discovered the wrong file was used by the program The stunning simulanty between these results and the classical phase space plagued us so much that we decided to incorporate the tuled structure as an initial condition and iterate the map forward, properly this time, and observe the evolution of the structure The next two pars of figures show the intial state and then the result after 2000 iterations of the mapping The stability of the pattern is startling and leads us to beheve that the preferred distribution is close to that in the images in figs $711 \& 713$ So the mistake made initially has paid dividends in that it has highlighted the existence of an extended tiled structure, which appears quite stable, in the quantum quasi-phase space

We end our examination of the quantum phase space by highlighting the correspondence between the qauntum and classical and by briefly mentioning further improvements and research directions The correspondence can be considered quite strong between the quantum system and the classical The on resonance and off resonance cases yeld phase spaces in both regimes which are definitely related The only minus is the rotation factor mentioned but a change of basis negates this The possible existence and stability of the extended tiled structure in the Wigner distribution for the resonance case when $\beta=\frac{\pi}{2}$ suggests that such a structure is a preferred by the system over others and that if the matrix was larger than its current $350 \times 350$ then we could possibly explore this more thoroughly As for future directions, a more detaled look at the finer structure of the Wigner distribution to check for further correspondences between the classical and quantum cases and an attempt to confirm, using large evolution matnces, the unboundedness of certain on resonance scenarios These results, though important, are only illuminating


Fig 711 Image of $P_{W}(Q, P)$ for the initial state corresponding to the diffuse distribution of the probability amplitudes as discussed in the text Note the distinctive tiled structure and the close relationship between this type of pattern and the classical phase space tuled structure

$F_{2 g} 712$ The corresponding contour plot of $P_{W}(Q, P)$ for the system as described for the 1 mage plot in fig (711) above


Fig 719 Image of $P_{W}(Q, P)$ for system 2000 time steps after the imitial state shown in figs (711) The system is on resonance with $\beta=\frac{\pi}{2}$ and $\mu_{q}=15$ The stability of the tuled pattern is evident from this images and contour plot below


Fag 714 The corresponding contour plot of $P_{W}(Q, P)$ for the system as described for the image plot in fig (713) above.
a very small region of what could turn out to be a very nich and beautiful area of quantum mechanıcs

## CHAPTER 8 CONCLUSIONS

The conclusions that can be drawn from the work presented in this thesis are many as are the number of routes further work can be persued The Harmonc Oscllator was chosen primarily because it is one of the most fundamental physical systems and not a mathematically-inspired mapping, such as the precewise-hnear standard map of Scharf \& Sundaram (1991) Though some conclusions can be drawn from such unphysical systems they cannot be attributed as purely physical phenomena as such systems are physically unreal The harmonic oscillator is not and as such is an ideal candidate for analysis

In our classical analysis of the system's phase space we showed how the stochastic layer is the principle mechanism behind the diffuse transport and unbounded energy growth of certain orbits which he wathin the layer itself The principle directions of transport lllustrated in fig 31 show the diffusing of orbits along the (stable and unstable) manfolds of the period four hyperbolic points whose separatrix net is the layer itself for infinitesimal kıcking We have have found for the kıcked oscillator, at resonance, that the exastence of periodic orbits in or around the stochastic layer influences the breakup of the invariant orbits as would be expected from the K A M theorem As a result of ths breakup, the boundary of the layer becomes increasingly comphcated as the kick strength is increased Our analysis allows us to predict which resonances and hence which periodic orbits should be present in the system's phase space by means of a function obtained via a polynomial expansion of the kicking term The correlation between prediction and numerical results was found to be very good despite the practical problems in numerically locatmg complete sets of periodic orbits We further showed that the elliptic fixed points of period four at the center of the main invariant islands become hyperbolic for a kuck strength of 100 causing these islands to break in two for kick strengths above 100 (cf table $31 \&$ fig 3 2) This result holds the key to our quantum analysis, that being the origin becoming unstable above a certann kick strength allowing diffuse orbits to exist close to it for reasonable kıck strengths Therefore we can analyse the quantum mapping using low order eigenfunctions knowing that if there is a correspondence between the quantum and classical we should be
able to observe diffuse type wavefunctions for reasonable kicking We do in fact see this and therefore can conclude the two hmits do have a direct correspondence

The detailed study of the evolution of the stochastic layer as the kick strength is increased is of fundamental importance as it enables us to see the effect of increasing nonintegrability on the system at large The most striking change in the layer is its variation in width with kick strength, which we studied in detail and compared to that predicted by perturbative analytical arguments (Chernikov et al 1989) We found that the width of the layer as a function of the kacking strength follows an exponential variation of the form $u e^{-u}$, where $u=1 /($ kuck strength ) This is the form predicted by Chernkov et al (1989) for kıck strengths $\ll 1$ We find that this result is true for a range of kick strengths considerably in excess of the theoretical prediction This validity is surprising when one considers the increased complexity of the layer itself and the overall increase m structure at the layer boundary with increased kicking

The diffusion in the classical regime, analysed in chapter 4 based on the argument of Rechester \& White (1980), was found to correspond exactly with that for the kicked rotator and this correspondence was put down to the dominance of the turbulent motion over the invariant periodic \& restricted motion in the cells (refer to table 41 in chapter 4) This is important because such sımulanty, when the turbulent motion is dominant, is evidently generic of all one-dımensional hamultoman systems when the kicking term is cosinusoidal Furthermore, the lower asymptotic limit was found to be identical to that found by Lachtenberg and Wood (1989) who constructed an approxamate theory of diffusion on the stochastic web which is valid in the small $\mu_{c l}$ region Our results on the Resonance Enhanced Diffusion are unque in that no-one has yet shown the spikes in the diffusion coefficient to exist for the kicked harmonic oscillator (Ishizaki et al (1989) showed this diffusion to exast for the kucked rotator) or explucitly shown them to vary as the phase space's periodicity changes Our argument for this anomalous diffusion is based on strong correlations between successive points and follows on logically from the arguments of Rechester \& White The main conclusion to be drawn from this diffusion analysis is that correlations do exist for reasonable kıck strengths but normally die off when the kick strength tends to infinty However the anomalous diffusion has strong correlations
regardless of how strong the kyck stength is once it equals a positive integer tımes the system's phase space

The quantum mapping we derived has not been, to our knowledge, used before Furthermore the approach we use, to obtain thes mapping, based on the approach of Fox et al (1990 \& 1994) has not been apphed to this system etther The approach is quite simple and was chosen for its directness, that of relating the probability amplitudes for successive kıcks The success of this approach can be gauged from our abllty to calculate energies, quası-energy levels as well as wavefunctions and Wigner distributions From our analysis of the system we have been able to make some interesting conclusions about the quantum system We have seen certain states naturally saturate giving rise to quası energy levels (Israilev 1990) in the kicked system The corresponding localisation in the wavefunction has also been examined and prominent peaks found to exast in these local states For larger kıck strengths the wavefunctions spread out with an increasing rapidity with kick strength and their energy increase was found to be quite linear (up to the point they hit the boundary imposed by the finte matrix size) We used this chearity to calculate a quantum deffusion coefficient and found it to have a definite non linear form In fact we found the power-law to be of the form $02\left(\mu_{q}\right)^{47}$ The wavefunctions for these kuck strengths were found to be quite diffuse with little structure on them except for a gradual fall off $Q$ tends to $\infty$ (caused by the finite number of eigenfunctions used) We conclude from these latter results that the energy saturation is definitely artificial (caused by the finte matrix size) and that the energy was increasing linearly in a seemingly unbounded fashion and not following the nature of the suppresion of chaos in the kicked rotator

The Wigner distribution for the quantum phase space can be seen to have a definte four-fold symmetry when the system is on resonance ( $\beta=\frac{\pi}{2}$ ) and, for a simular initial state and kıck strength, the resonance case is more spread out and defintely of higher energy than its off resonance cousin The off resonance case also has a more circularly symmetric nature about it than the resonance case These point to a definite correspondence between the classical case (the off resonant \& resonant phase space structures visible in chapter 2) and the quantum case (the Wigner distributions in chapter 7) Furthermore we observed from an muxed imitial state, at a kick strength of 15 , that the resonant structure formed
is very close to the tiled structure we used to test for the exastence of a stable extended tiled structure ( re figs 711-714) We found that this mixed mintial state, at that kıck strength, became diffuse as time evoloved (as shown in chapter 6) This, coupled with the similarity to the extended tiled structure, is indicative of continuing diffusion of this wavefunction over the phase space Our reasoning is quite simple, such an extended structure bears a striking resemblence to the classical tiled phase space structure and, as the wavefunction is spreading out without any indication of saturation, we conclude that this is proof that the system is behaving as if it has a quantum analog to the classical stochastic layer Therefore we conclude the existence of quantum chaos in this system as the quantum regime has wavefunctions which, from the results presented in chapters $6 \& 7$, mımic classical ensembles which have elements in the stochastic layer It is using ensembles like this that the classical phase space is constructed and to have quantum wavefunctions construct a tiled structure so close is proof that some correspondence exasts between the dynamics of both limots

We have accomplished our main objective, that of showing the existence of quantum chaos in the kıcked harmonic oscillator The research undertaken here has brought to light many unknowns about the system itself and we have also highlighted some of the ways in which future research on this system can proceed The kicked harmonic oscillator is a truly physical system, one whose properties should more fully increase our knowledge on thes fledghing area of quantum chaos

## APPENDIX A EXPLICIT FORMULATION OF THE POTENTIAL TERM

The form of the equation that we have chosen for the potential is $V(p, q)=$ $\mu_{q}(\cos (k q))$ (after Berman et al. (Berman et al. (1989)). This term is in effect the non-linearity in the system Hamiltonian and as such makes the integrable oscillator equations non-integrable. The degree of non-linearity depends on how many terms in the series expansion of $\cos (x)$ are required to accurately describe the potential term. As is well known, the larger the argument present in the cosine then the higher the number of terms needed to be accurate. Using the cosine series expansion we will show how the potential term in eq. (2.1.15) is arrived at. From the definitions in eqs. (2.1.3) and (2.1.4), for the operators $a(t)$ and $a^{\dagger}(t)$, we have for $q(t)$

$$
\begin{equation*}
q(t)=\frac{1}{\iota} \sqrt{\frac{\hbar}{2 M \omega_{0}}}\left(a^{\dagger}(t)-a(t)\right) \tag{A.1}
\end{equation*}
$$

So our equation for the potential becomes

$$
\begin{equation*}
V\left(a, a^{\dagger}\right)=\mu_{q}\left(\cos \left(\frac{k}{\iota} \sqrt{\hbar /\left(2 M \omega_{0}\right)}\left(a^{\dagger}(t)-a(t)\right)\right)\right) \tag{A.2}
\end{equation*}
$$

For convenience we define $K$ to be $\left(k \sqrt{\hbar /\left(2 M \omega_{0}\right)}\right)$. We show in chapter 2 how the $\hbar$ dependence is removed from $K$. Thus eq.(A.2) is simplified to

$$
\begin{equation*}
V\left(a, a^{\dagger}\right)=\mu_{q}\left(\cos \left(\frac{K^{\prime}}{t}\left(a^{\dagger}(t)-a(t)\right)\right)\right) \tag{A.3}
\end{equation*}
$$

To proceed further we have to change the compound argument of the cosine into simpler arguments of cosines and sine. This we accomplish using the well known trigonometric identity:

$$
\operatorname{Cos}(A-B)=\operatorname{Cos}(A) \operatorname{Cos}(B)+\operatorname{Sin}(A) \operatorname{Sin}(B)
$$

Eq.(A.3) now becomes:

$$
\begin{equation*}
V\left(a, a^{\dagger}\right)=\mu_{q}\left(\cos \left(\frac{K}{\iota} a^{\dagger}(t)\right) \cos \left(\frac{K}{\iota} a(t)\right)+\sin \left(\frac{K}{\iota} a^{\dagger}(t)\right) \sin \left(\frac{K}{\iota} a(t)\right)\right) \tag{A.4}
\end{equation*}
$$

We do not have any explicit rules for the commutation of a and $a^{\dagger}$ with trigonometric functions whose arguments are either a or $a^{\dagger}$. To allow the commutation to proceed we
find it necessary to express the sines and cosines in terms of their series expansions These series expansions, for the sine and cosine functions, are well known and for our arguments break down to

$$
\begin{gather*}
\cos \left(\frac{K}{\iota} a^{\dagger}(t)\right)=\sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n}}{(2 n)^{\prime}}\left(a^{\dagger}(t)\right)^{2 n}  \tag{A5a}\\
\sin \left(\frac{\hbar}{\iota} a^{\dagger}(t)\right)=\sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n+1}}{(2 n+1)^{\prime}}\left(a^{\dagger}(t)\right)^{2 n+1} \tag{A5b}
\end{gather*}
$$

So when expressed m this series form eq (A 6) takes on the formidable but more useful appearance

$$
\begin{array}{r}
V\left(a, a^{\dagger}\right)=\mu_{q}\left\{\sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n}}{(2 n)^{\prime}}\left(a^{\dagger}(t)\right)^{2 n} \sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n}}{(2 n)^{\prime}}(a(t))^{2 n}\right\} \\
+\mu_{q}\left\{\sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n+1}}{(2 n+1)^{1}}\left(a^{\dagger}(t)\right)^{2 n+1} \sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n+1}}{(2 n+1)^{1}}(a(t))^{2 n+1}\right\} \tag{A6}
\end{array}
$$

The next task is to commute this with $a(t)$ for the explicit form of $-\iota\left(a, V\left(a, a^{\dagger}\right)\right)$ Eqs (2 $18 a-d$ ) will be put to good use here to simphfy the equations immensely For the cosine terms we have a commutation of the form

$$
\begin{gather*}
\left(a(t), \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n} \sum_{n=0}^{\infty}(a(t))^{2 n}\right) \\
=\left(a(t), \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n}\right) \sum_{n=0}^{\infty}(a(t))^{2 n}+\left(a(t), \sum_{n=0}^{\infty}(a(t))^{2 n}\right) \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n} \tag{A7}
\end{gather*}
$$

using the relation in eq ( $218 c$ ) Taking the summation after the commutation makes no difference so for the first half of the nght hand side in eq ( $A 7$ ) we have

$$
\begin{equation*}
\left(a(t), \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n}\right) \sum_{n=0}^{\infty}(a(t))^{2 n} \equiv \sum_{n=0}^{\infty}\left(a(t),\left(a^{\dagger}(t)\right)^{2 n}\right) \sum_{n=0}^{\infty}(a(t))^{2 n} \tag{A8}
\end{equation*}
$$

whereas for the latter half we get

$$
\begin{equation*}
\left(a(t), \sum_{n=0}^{\infty}(a(t))^{2 n}\right) \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n} \equiv \sum_{n=0}^{\infty}\left(a(t),(a(t))^{2 n}\right) \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n} \tag{A.9}
\end{equation*}
$$

The parts to be commuted with $a(t)$ consist now of an operator to the power of 2 n There is one problem, however, and that is the $n=0$ term of the summation This gives ( $a(t), 1$ ) which is zero So the summation should be performed from $n=1$ to $n=\infty$ with the $n=0$ term separate Noting that, using eq (218d), $\left(A, B^{2 n}\right)=2 n(A, B) B^{2 n-1}$ we get for eq (A8)

$$
\begin{equation*}
\left(a(t), \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n}\right) \sum_{n=0}^{\infty}(a(t))^{2 n} \equiv \sum_{n=1}^{\infty} 2 n\left(a^{\dagger}(t)\right)^{2 n-1} \sum_{n=0}^{\infty}(a(t))^{2 n} \tag{A10}
\end{equation*}
$$

for eq (A9) the story is different because $a(t)$ commutes with itself so that $\left(a(t),(a(t))^{2 n}\right)$ $\equiv 2 n(a(t), a(t)) a(t)^{2 n-1} \equiv 2 n * 0 * a(t)^{2 n-1} \equiv 0$ So substıtutıng these identities into eq ( $A$ 7) we get

$$
\begin{equation*}
\left(a(t), \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n} * \sum_{n=0}^{\infty}(a(t))^{2 n}\right)=\sum_{n=1}^{\infty} 2 n\left(a^{\dagger}(t)\right)^{2 n-1} \sum_{n=0}^{8}(a(t))^{2 n} \tag{A11}
\end{equation*}
$$

So the complete cosine term from eq (A6) becomes

$$
\begin{equation*}
\sum_{n=1}^{\infty} \frac{(-1)^{n}(K / L)^{2 n}}{(2 n)^{\prime}} 2 n\left(a^{\dagger}(t)\right)^{2 n-1} * \sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n}}{(2 n)^{\prime}}(a(t))^{2 n} \tag{A12}
\end{equation*}
$$

The summation from 1 to $\infty$ presents no problem as by introducing a new dummy variable $u=n-1$ we get

$$
\begin{equation*}
\sum_{n=1}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n}}{(2 n)^{\prime}} 2 n\left(a^{\dagger}(t)\right)^{2 n-1} \rightarrow-\frac{K}{\iota} \sum_{u=0}^{\infty} \frac{(-1)^{u}(K / \iota)^{2 u+1}}{(2 u+1)^{1}}\left(a^{\dagger}(t)\right)^{2 u+1} \tag{A13}
\end{equation*}
$$

which 15 Just $-\frac{h}{\iota} \sin \left(\frac{K}{\iota} a^{\dagger}(t)\right)$ So the $\cos \left(\frac{K}{\iota}\left(a^{\dagger}(t)\right)\right) \cos \left(\frac{K}{\iota}(a(t))\right)$ term becomes $-\frac{K}{\iota} \sin \left(\frac{K}{\iota}\left(a^{\dagger}(t)\right)\right) \cos \left(\frac{K}{\iota}(a(t))\right)$ following the commutation with $a(t)$

For the sine terms we have a simalar argument with the commutation being of the form given below in eq (14)

$$
\begin{gather*}
\left(a(t), \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n+1} * \sum_{n=0}^{\infty}(a(t))^{2 n+1}\right)= \\
\left(a(t), \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n+1}\right) \sum_{n=0}^{\infty}(a(t))^{2 n+1}+\left(a(t), \sum_{n=0}^{\infty}(a(t))^{2 n+1}\right) \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n+1} \tag{A14}
\end{gather*}
$$

Taking the approach we used for the cosine terms we get

$$
\begin{equation*}
\left(a(t), \sum_{n=0}^{\infty}\left(a^{\dagger}(t)\right)^{2 n+1} * \sum_{n=0}^{\infty}(a(t))^{2 n+1}\right)=\sum_{n=0}^{\infty}(2 n+1)\left(a^{\dagger}(t)\right)^{2 n} \sum_{n=0}^{\infty}(a(t))^{2 n+1} \tag{A15}
\end{equation*}
$$

The term modfied $\left(a^{\dagger}(t)\right)$ by the commutation with $a(t)$ affects the total sine sum as follows

$$
\begin{equation*}
\sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n+1}}{(2 n+1)^{\prime}}(2 n+1)\left(a^{\dagger}(t)\right)^{2 n} \rightarrow \frac{K}{\iota} \sum_{n=0}^{\infty} \frac{(-1)^{n}(K / \iota)^{2 n}}{(2 n)^{\prime}}\left(a^{\dagger}(t)\right)^{2 n} \tag{A16}
\end{equation*}
$$

which is Just $\frac{K}{\iota} \cos \left(\frac{K}{\iota} a^{\dagger}(t)\right)$ Thus the term, $\sin \left(\frac{K}{\iota}\left(a^{\dagger}(t)\right)\right) \sin \left(\frac{K}{\iota}(a(t))\right)$, becomes $\frac{K}{\iota} \cos \left(\frac{K}{\iota}\left(a^{\dagger}(t)\right)\right) \sin \left(\frac{K}{\iota}(a(t))\right)$ after the commutation

So we have now obtained expressions for both the cosine product term and the sine product term after commuting with $a(t)$ What remains now is to combine these results to obtain an expression for the potential after commutation ( a e the term $f\left(a, a^{\dagger}\right)$ ) We find that, in its final form, the potential term satisfies

$$
=\mu_{q}\left\{K \operatorname { s i n } \left(\frac{h}{\iota}\left(a(t), V\left(a(t), a^{\dagger}(t)\right)\right) .\right.\right.
$$

or more succintly

$$
\begin{equation*}
f\left(a(t), a^{\dagger}(t)\right)=-\iota\left(a(t), V\left(a(t), a^{\dagger}(t)\right)\right)=\mu_{q}\left\{K\left(\sin \left(\frac{h}{\iota}\left(a^{\dagger}(t)-a(t)\right)\right)\right)\right\} \tag{A17}
\end{equation*}
$$

This is the substitution we make for $f\left(a, a^{\dagger}\right)$ in chapter 2

## APPENDIX B THE ANNIHILATION \& CREATION OPERATORS

In eqs (517) and (518) we defined the $a \& a^{\dagger}$ operators in terms of $Q$ and $\frac{d}{d Q}(\equiv P)$ The eigenfunctions of the hamultoman can be found without using the operator technique, previously, by explicitly solving the differential equation with trial solutions This way the solutions found are identical to those found using the more direct operator approach The eigenfunctions found $u_{n}(Q)$ have the form

$$
\begin{equation*}
u_{n}(Q)=\frac{1}{\sqrt{2^{n} \sqrt{\pi}\left(n^{\prime}\right)}}\left(Q-\frac{d}{d Q}\right)^{n} \exp \left(-\frac{1}{2} Q^{2}\right) \tag{array}
\end{equation*}
$$

where $u_{o}(Q)=\frac{1}{\sqrt{\sqrt{\pi}}} \exp \left(-\frac{1}{2} Q^{2}\right)$ and $u_{1}(Q)=\frac{1}{\sqrt{2}}\left(Q-\frac{d}{d Q}\right) u_{o}(Q) \equiv a^{\dagger} u_{o}(Q)$ Likewise it can be shown that for any $n$

$$
\begin{align*}
& u_{n+1}(Q)= \\
& \quad \frac{1}{\sqrt{2^{n+1} \sqrt{\pi}\left((n+1)^{\prime}\right)}}\left(Q-\frac{d}{d Q}\right)^{n+1} \exp \left(-\frac{1}{2} Q^{2}\right)=\frac{1}{\sqrt{(n+1) 2}}\left(Q-\frac{d}{d Q}\right) u_{n}(Q) \tag{B2}
\end{align*}
$$

thus

$$
u_{n+1}(Q) \equiv \frac{1}{\sqrt{n+1}} a^{\dagger} u_{n}(Q) \quad \text { or } \quad \sqrt{n+1} u_{n+1}(Q) \equiv a^{\dagger} u_{n}(Q)
$$

The anmhilation operator $a=\frac{1}{\sqrt{2}}\left(\frac{d}{d Q}+Q\right)$ does the opposite of the $a^{\dagger}$ operator by taking a polynomial of order $n$ down to a polynomial of order $n-1$ with a $\sqrt{n}$ coefficient How this comes about is by looking at the commutation properties of $a$ and $a^{\dagger}, 1$ e

$$
\left(a, a^{\dagger}\right)=a a^{\dagger}-a^{\dagger} a=1
$$

so that if we operate on $u_{n}$ then

$$
\begin{equation*}
\left(a, a^{\dagger}\right) u_{n}(Q)=a a^{\dagger} u_{n}(Q)-a^{\dagger} a u_{n}(Q)=u_{n}(Q) \tag{B4}
\end{equation*}
$$

Hence $\sqrt{n+1} * a u_{n+1}(Q)-a^{\dagger} a u_{n}(Q)=u_{n}(Q)$ Assuming the operator $a$ changes $u_{n}(Q)$ to $\sqrt{n+r} u_{n+p}(Q)$ then we have that

$$
\begin{equation*}
\sqrt{n+1} \sqrt{n+1+r} u_{n+p+1}(Q)-\sqrt{n+r} \sqrt{n+p+1} u_{n+p+1}(Q)=u_{n}(Q) \tag{B5}
\end{equation*}
$$

As the $u_{n}(Q)$ form an othogonal basis then $n+p+1=n \Leftrightarrow p=-1 \quad$ Furthermore $\sqrt{n+1} \sqrt{n+1+r}-\sqrt{n+r} \sqrt{n+p+1}=1 \quad n+p+1=n$ so $\sqrt{n+1} \sqrt{n+1+r}-$ $\sqrt{n+r} \sqrt{n}=1 \Leftrightarrow r=0$ Thus $a u_{n}(Q)=\sqrt{n} u_{n-1}(Q)$ as stated earker in this section

This ends our brief justıfication of the $a \& a^{\dagger}$ definitions Though not very elegant, this justafication serves for me as a way to visualise how these operators manipulate eigenfunctions of the quantum system

## APPENDIX C SIMPLIFICATION OF THE COMPLEX TERM IN $P_{w}(\mathrm{Q}, \mathrm{P})$

Here we want to sumplify the complex terms contained within the $\}$ brackets in eqs (6214-15) so that we can show the Wigner distribution, $P_{W}(Q, P)$, to be real valued (ff not always positive) To do this we shall spht the complex term into two parts the first will be the product $a_{n}^{*} a_{m}$ whule the second will deal with the $(-Q+\iota P)^{m-n}$ term

Let the complex probability amplitude $a_{k}=a_{k}^{(r)}+\iota a_{k}^{(1 m)}$ then the expression for $a_{n}^{*} a_{m}$ becomes

$$
\begin{equation*}
a_{n}^{*} a_{m}=\left(a_{n}^{(r)} a_{m}^{(r)}+a_{n}^{(1 \mathrm{~m})} a_{m}^{(1 \mathrm{~m})}\right)+\iota\left(a_{n}^{(r)} a_{m}^{(\mathrm{lm})}-{ }_{n}^{(1 \mathrm{~m})} a_{m}^{(r)} a\right) \equiv \alpha+\iota \beta \tag{C1}
\end{equation*}
$$

For the case when $m$ and $n$ are interchanged we get $\alpha-\iota \beta$ Furthermore for any complex number $z$ taken to the power $m$ we find that it still has a complex conjugate relationship to its complex conjugate taken to the same power

$$
(x+\iota y)^{n}=
$$

$$
\begin{equation*}
x^{n}+\iota n x^{n-1} y+(\iota)^{2}\binom{n}{2} x^{n-2} y^{2}+\quad+(\iota)^{n-2}\binom{n}{2} x^{2} y^{n-2}+(\iota)^{n-1} n x y^{n-1}+(\iota y)^{n} \tag{C2a}
\end{equation*}
$$

and
$(x-\iota y)^{n}=$
$x^{n}-\iota n x^{n-1} y+(-\iota)^{2}\left({ }^{n}\right) x^{n-2} y^{2}+\quad+(-\iota)^{n-2}\left({ }^{n}\right) x^{2} y^{n-2}+(-\iota)^{n-1} n x y^{n-1}+(-\iota y)^{n}$
As $(\iota)^{2 m}$ is equal to $(-\iota)^{2 m} \forall m \in \mathbf{Z}$ and $(\iota)^{2 m+1}=-(-\iota)^{2 m+1}$ again for $m \in \mathbf{Z}$ then the above two equations give a real part which is identical but imaginary parts which are the negative of each other thus making $(x-\iota y)^{n}$ the complex conjugate of $(x+\iota y)^{n}$ Therefore $(-Q+\iota P)^{m-n}$ is complex conjugate to $(-Q-\iota P)^{m-n}$ or, in our case when $m$ and $n$ are interchanged, $(-Q-\iota P)^{n-m}$ Therfore we have the sum of the product of two complex number with the product of ther complex conjugates below $a_{n}^{*} a_{m}$ is denoted $(a+\iota b)$ and $(-Q+\iota P)^{m-n}$ is replaced by $(\gamma+\iota \delta)$ so that the complete complex term, $a_{n}^{*} a_{m}(-Q+\iota P)^{m-n}$, ss

$$
\begin{align*}
& a_{n}^{*} a_{m}(-Q+\iota P)^{m-n}=(a+\iota b)(c+\iota d)  \tag{C3}\\
& (a+\iota b)(\gamma+\iota \delta)=(a \gamma-b \delta)+\iota(b \gamma+a \delta) \\
& (a-\iota b)(\gamma-\iota \delta)=(a \gamma-b \delta)-\iota(b \gamma+a \delta)
\end{align*}
$$

which when summed gives a real number, $2(a \gamma-b \delta)$ Thus the distribution is real To numerically evaluate the equation(s) leading to the Wigner distribution would require an immense amount of computing power so any tricks to reduce this cost in time would be welcome The phase space of $P \& Q$ is spht into four quadrants but here we shall restrict our argument to Just two of them ( 1 and 3 ) and then apply the results to the other two (we can do this because any point $(a, b)$ in the quadrant 1 is just $(-1)(a, b)$ in 3 and likewise for 2 and 4)


## Schematic of the four quadrants constatuting the phase space

The distributions for the respective quadrants are

$$
\begin{array}{ll}
P_{W}^{1}(Q, P)=\sum_{m=0}^{\infty} \sum_{n=0}^{m} A_{n+1, m+1}\left\{a_{n}^{*} a_{m}(-Q+\iota P)^{m-n}\right\} & n \leq m \\
P_{W}^{1}(Q, P)=\sum_{n=0}^{\infty} \sum_{m=0}^{n} A_{n+1, m+1}\left\{a_{n}^{*} a_{m}(-Q-\iota P)^{n-m}\right\} & m \leq n \tag{C4b}
\end{array}
$$

for quadrant 1 (note the superscript 1 on $P_{W}(Q, P)$ and

$$
\begin{align*}
& P_{W}^{3}(Q, P)=\sum_{m=0}^{\infty} \sum_{n=0}^{m}(-1)^{m-n} A_{n+1, m+1}\left\{a_{n}^{*} a_{m}(-Q+\iota P)^{m-n}\right\}  \tag{C4c}\\
& n \leq m  \tag{C4d}\\
& P_{W}^{3}(Q, P)=\sum_{n=0}^{\infty} \sum_{m=0}^{n}(-1)^{n-m} A_{n+1, m+1}\left\{a_{n}^{*} a_{m}(-Q-\iota P)^{n-m}\right\}
\end{align*} \quad m \leq n
$$

for quadrant 3 where we have defined $A_{n+1, m+1}$ to be the matrix element

$$
\exp \left(-\left(Q^{2}+P^{2}\right)\right)(-1)^{m} \sqrt{\frac{2^{m} n^{1}}{2^{n} m^{1}}} L_{n}^{m-n}\left(2\left(Q^{2}+P^{2}\right)\right) \quad n \leq m
$$

or

$$
\exp \left(-\left(Q^{2}+P^{2}\right)\right)(-1)^{n} \sqrt{\frac{2^{n} m^{\prime}}{2^{m} n^{!}}} L_{m}^{n-m}\left(2\left(Q^{2}+P^{2}\right)\right) \quad m \leq n
$$

It is clear from eqns $(C 4 a-d)$ that for $m-n(n \leq m)$ even, the terms in the thurd quadrant equal that in the first whereas for $m-n$ odd the terms are of equal magnitude but opposite signs The same is true for $n-m$ when $m \leq n$ To lllustrate things lets restrict $m$ and $n$ to values less than 4 The matrices would be as follows
$\left(\begin{array}{llll}A_{11} & A_{12} & A_{13} & A_{14} \\ A_{12} & A_{22} & A_{23} & A_{24} \\ A_{13} & A_{23} & A_{33} & A_{34} \\ A_{14} & A_{24} & A_{34} & A_{44}\end{array}\right)$ for quadrant 1 $\left(\begin{array}{cccc}A_{11} & -A_{12} & A_{13} & -A_{14} \\ -A_{12} & A_{22} & -A_{23} & A_{24} \\ A_{13} & -A_{32} & A_{33} & -A_{34} \\ -A_{14} & A_{24} & -A_{34} & A_{44}\end{array}\right)$ for quadrant 3

In general these two matrices are very different however we will now show that if we restrict the initial condition to a PURE eigenstate then the matrices for both quandrants one and three are the same To see this refer back to chapter 5 where we derive the quantum mapping In eq ( 5131 ), we show that there is included in the expression for $A_{m}((N+1) \tau)$, a term contanned within $\}$ which forces the expression to zero if $p-m$ is odd Therefore if the imtial state is a pure odd numbered state then, morder for $\boldsymbol{p}-\boldsymbol{m}$ to be even, only odd states exist in the resulting system (similarly if we begin with a pure even numbered state) Thus we end up in eqs ( $C 4 c-d$ ) with the same expression for $P_{W}$ as we have in eqs ( $C 4 a-b$ ) The importance of this is to allow us to cut down on time when calculating the phase space itself by obtaining two quandrants while only calculating one

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